BRITISH GEOLOGICAL SURVEY Hartland Observatory Monthly Magnetic Bulletin May 2009 09/05/HA



Hotel

HARTLAND OBSERVATORY MAGNETIC DATA

1. Introduction

This bulletin is published to meet the needs of both commercial an d acad emic u sers o f g eomagnetic data. Magnetic observatory data is p resented as a series of plots of one -minute, hour ly a nd daily values, followed by tabulations of monthly values, geomagnetic activity i ndices a nd r eports of r apid variations. T he ope ration of t he observatory and presentation of data are described in the rest of this section.

Enquiries about the data should be addressed to:

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2. Position

Hartland Observatory, one of the three geomagnetic observatories operated and m aintained in the UK by BGS, is situated on the N W boundary of the village of H artland i n N orth D evon. T he observatory co-ordinates are:

Geographic:	50°59.7′N	355°30.96′E
Geomagnetic:	53°54.42 'N	080°06.54′E
Height above m	ean sea level:	95 m

The geomagnetic co-ordinates are calculated using the 10th ge neration I nternational G eomagnetic Reference Field at epoch 2009.5.

3. The Observatory Operation

3.1 GDAS

The observatory operates under the control of the Geomagnetic Data Acq uisition System (GDAS), which was developed by B GS st aff, i nstalled i n 2002, a nd be came f ully ope rational i n January 2003. The system operates under the control of data acquisition s oftware r unning on QNX computers, which c ontrol t he da ta l ogging a nd communications.

There are two sets of sensors used for making magnetic m easurements. A t ri-axial l inear-core fluxgate magnetometer, m anufactured by t he Danish Meteorological Institute, is used to measure the variations in the horizontal (H) and vertical (Z) components of t he f ield. T he t hird s ensor i s oriented p erpendicular t o t hese, and measures variations, which are proportional to the changes in declination (D). Measurements are made at a rate of 1 Hz.

In addition to the fluxgate sensors there is a proton precession magnetometer making measurements of the a bsolute to tal field in tensity (F) at a r ate o f 0.1Hz.

The raw unfiltered data are retrieved automatically via I nternet c onnections to t he B GS of fice i n Edinburgh in near real-time. The fluxgate data are filtered t o pr oduce one -minute values using a 61 point c osine f ilter w hilst th e to tal f ield in tensity samples ar e f iltered u sing a 7-point cosine f ilter. These o ne-minute values are u sed t o u pdate t he Geomagnetism I nformation and Forecast Service (GIFS), an on-line information system accessed via the World Wide Web at the address given in Section 1.1. GIFS also pr ovides i nformation on geomagnetic and solar activity.

3.2 Back-up Systems

There are two ot her f ully i ndependent i dentical systems, GDAS 2 and GDAS 3, o perating at the observatory. The data from these are also processed in n ear r eal-time a nd us ed f or quality control purposes. They can also be used to fill any gaps or replace any corrupt values in the p rimary system, GDAS 1.

3.3 Absolute Observations

The GDAS f luxgate m agnetometers accurately measure variations in t he c omponents of t he geomagnetic field, but not the absolute magnitudes. Two sets of absolute measurements of the field are made m anually o nce p er week. A f luxgate sen sor mounted on a theodolite is used to determine D and inclination (I); the GDAS PPM measurements, with a site difference correction applied, are used for F. The absolute observations are used in conjunction with t he GDAS v ariometer m easurements to produce a continuous record of the absolute values of t he g eomagnetic field el ements as if they had been measured at the observatory reference pillar.

4. Data Presentation

The data presented in the bulletin are in the form of plots and tabulations described in the following sections.

4.1 Absolute Observations

The absolute o bservation m easurements m ade during the month are tabulated. A lso included are the corresponding b aseline v alues, wh ich ar e t he differences between the absolute measurements and the variometer measurements of D, H and Z (in the sense absolute–variometer). These are also plotted (markers) along with the derived preliminary daily baseline v alues (line) t hroughout the year. Daily mean differences between the measured absolute Fand the F computed from the baseline corrected Hand Z values are plotted in the fourth panel (in the sense measured–derived). The bot tom pa nel shows the daily mean temperature in the fluxgate chamber.

4.2 Summary magnetograms

Small-scale magnetograms are plotted which allow the month's data to be viewed at a glance. They are plotted 16 days a page and show the variations in D, H and Z. The scales are shown on the right-hand side of the page. On disturbed days the scales are multiplied by a factor, which is indicated above the panel for that day. The variations are centred on the monthly mean value, shown on the left side of the page.

4.3 Magnetograms

The daily magnetograms a re pl otted us ing one minute values of D, H and Z from the fluxgate sensors, with an y g aps filled u sing b ack-up da ta. The magnetograms are plotted to a v ariable scale; scale bars are shown to the right of each plot. The absolute level (the m onthly m ean v alue) i s indicated on the left side of the plots.

4.4 Hourly Mean Value Plots

Hourly mean values of D, H and Z for the past 12 months a re pl otted i n 27 -day seg ments corresponding to the Bartels solar rotation number. Magnetic disturbances asso ciated wi th act ive regions on the surface of the Sun may recur after 27 days: the sam e is t rue for geomagnetically quiet intervals. Plotting the d ata in th is way h ighlights this r ecurrence, an d al so i llustrates seaso nal an d diurnal variations throughout the year.

4.5 Daily and Monthly Mean Values

Daily m ean v alues of D, H, Z and F are p lotted throughout the year. In addition, a table of monthly mean v alues o f al 1 t he g eomagnetic el ements i s provided. These va lues de pend on a ccurate specification of t he f luxgate sen sor b aselines. Provisional a nd definitive v alues ar e i ndicated i n the t able as **P** or **D** respectively. It i s anticipated that provisional values will not be altered by more than a few nT or tenths of arcminutes before being made definitive.

4.6 Geomagnetic activity indices

The Observatory K index. This su mmarises geomagnetic activity at an observatory by assigning a code, an integer in the range 0 to 9, to each 3-hour Universal Time (UT) interval. The index for each 3-hour UT interval is determined from the ranges in H and in D (scaled in nT), with allowance made for the regular (undisturbed) diurnal va riation. The conversion from r ange t o a n index value is made using a qua si-logarithmic scal e, with the scale values dependent on the geomagnetic latitude of the observatory. The K index r etains the lo cal tim e (LT) and seasonal dependence of activity associated with the position of the observatory.

The provisional aa index. A num ber of 3 -hour geomagnetic indices are computed by combining K indices f rom ne tworks of obs ervatories to characterise global activity levels and to eliminate LT and seaso nal effects. The simplest of these is the *aa* index, c omputed us ing the K indices from two a pproximately a ntipodal observatories: Hartland in the UK and Canberra in Australia. The *aa* index is cal culated f rom linearisations of t he Hartland and Canberra K indices, and has units of nT. The daily mean value of *aa* (denoted *Aa*), the mean values of *aa* for the intervals 00-12UT and 12-24UT and the daily mean values f or Hartland alone (Aa_n) and Canberra alone (Aa_s) are tabulated.

Although the *aa* index is based on data from only two observatories, provided averages over 12 hours or longer are used, the index is strongly correlated with the *ap* and *am* indices, wh ich ar e d erived using da ta f rom m ore extensive observatory networks.

The *aa* indices lis ted in th is publication are provisional only; the definitive values are published by t he I nternational S ervice f or Geo magnetic Indices, C RPE/CNET - CNRS, 4 Av enue d e Neptune, F-94107 Saint Maur Cedex, France.

4.7 Rapid Variations

Charged particles stream from the Sun in the solar wind. T he so lar wind i nteracts with the geomagnetic field to cr eate a cav ity, t he magnetosphere, in which the field is confined. When a region of enhanced velocity and/or density in the solar wind arrives at the day-side boundary of the magnetosphere (at about 10 e arth radii) the boundary is pushed towards the Earth. Currents set up on the boundary of the magnetosphere can cause an abrupt change in t he ge omagnetic f ield measured on t he ground a nd t his is recorded on observatory m agnetograms a s a Sudden Impulse (SI). If, following an SI, there is a change in the rhythm of activity, the SI is termed a Storm Sudden Commencement (SSC). A classical magnetic storm exhibiting initial, main and recovery phases (shown by, f or i nstance, t he *Dst* ring c urrent i ndex) c an often occur after a S SC, in which case the start of the storm is taken as the time of the SSC.

Solar flares, seen at o ptical wav elengths as a sudden brightening of a small region of the Sun's surface, are also r esponsible f or i ncreased X -ray emissions. The X-rays cause increased ionisation in the ionosphere, which leads to absorption of shortwave r adio si gnals. On an observatory magnetogram a Solar Flare Ef fect (SFE), o r "crochet" may b e observed. This i s an enhancement to the diurnal variation of the order of 10 nT, lasting about an hour.

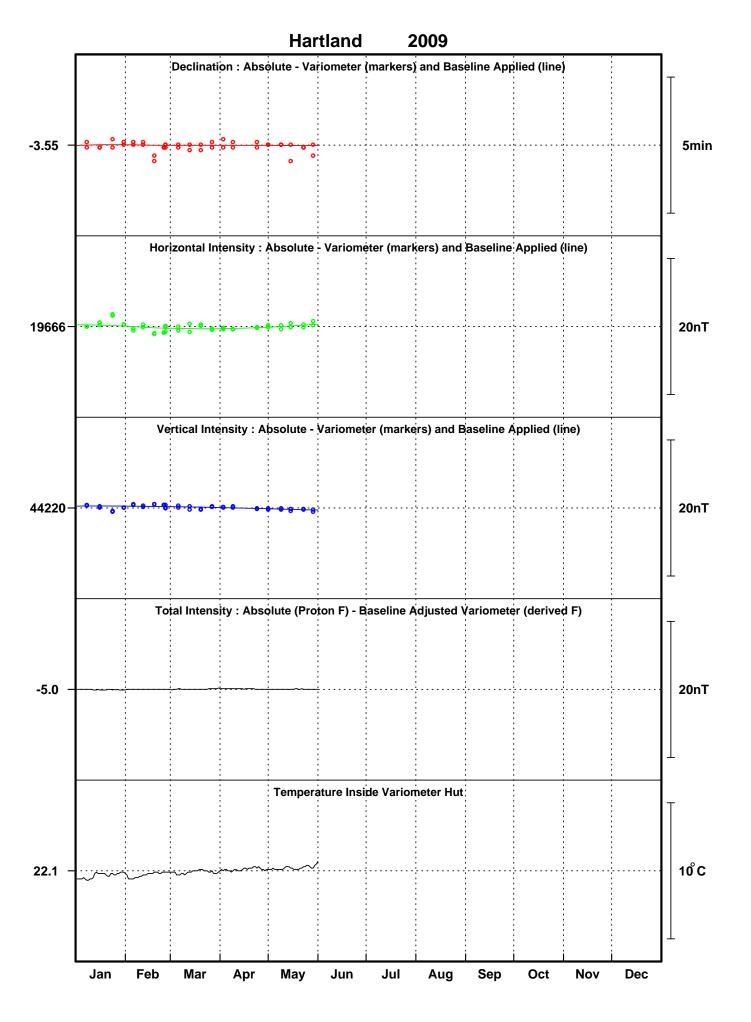
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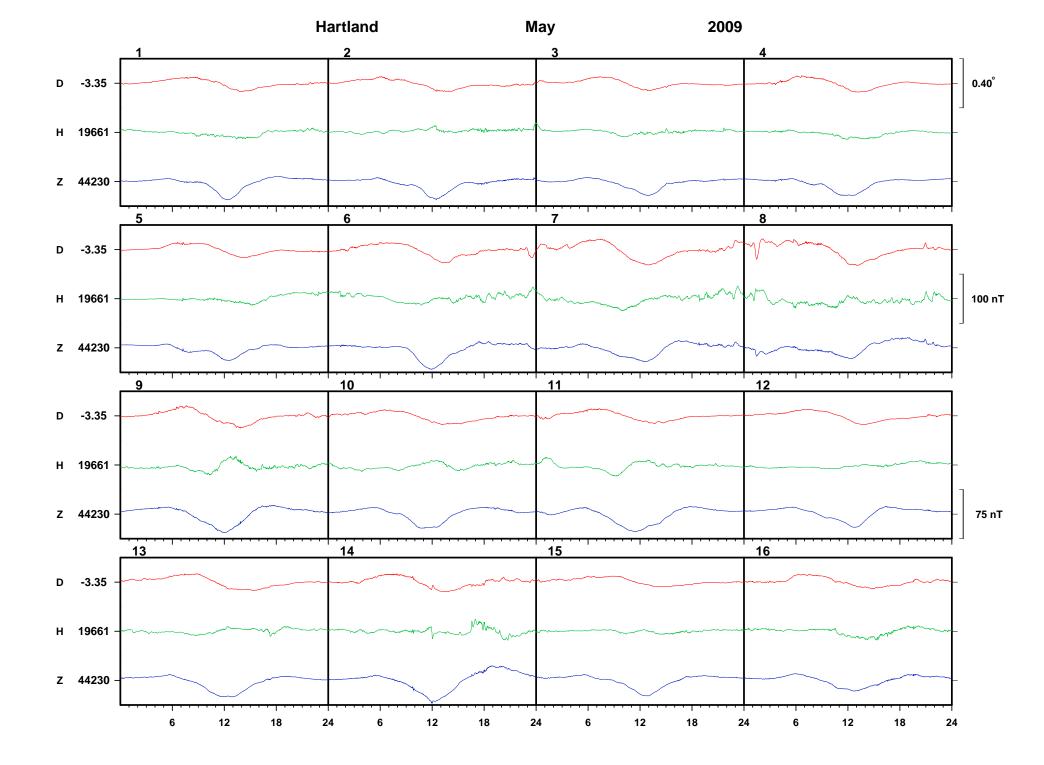
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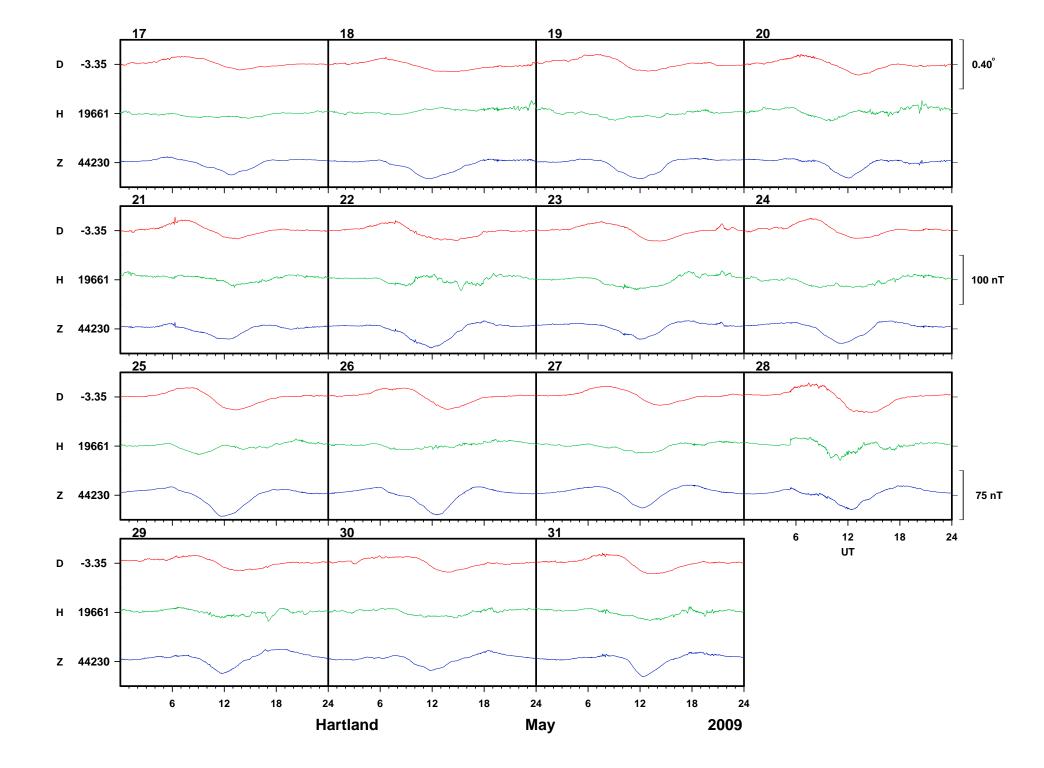
HARTLAND OBSERVATORY

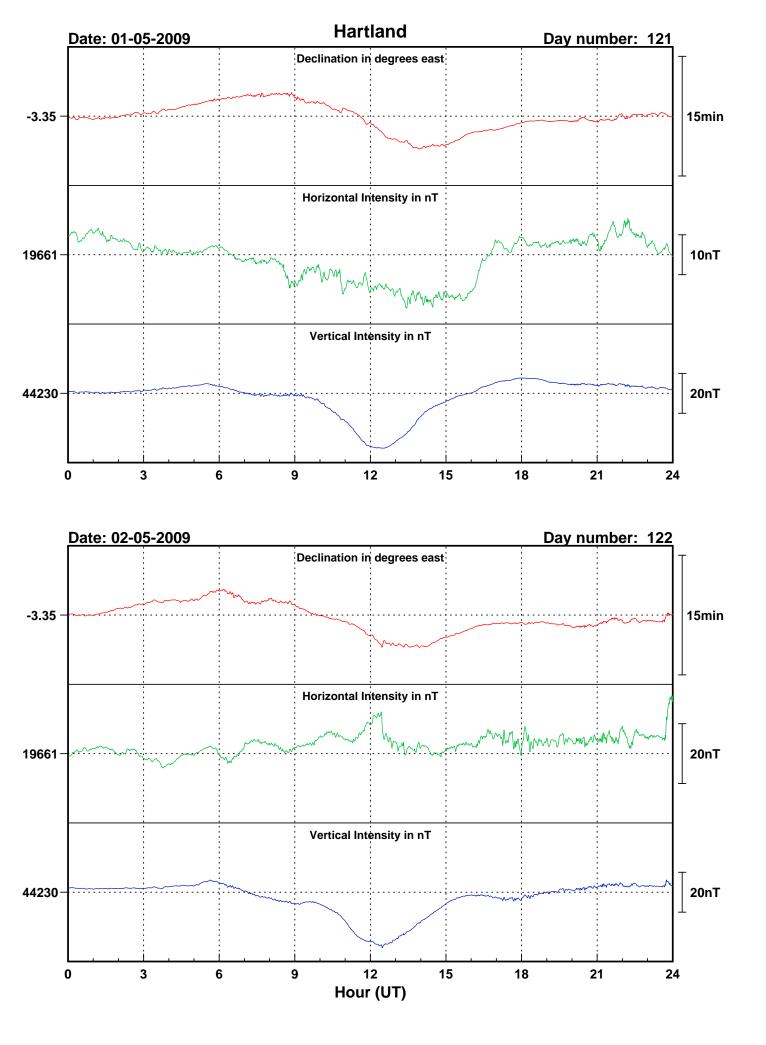
ABSOLUTE OBSERVATIONS

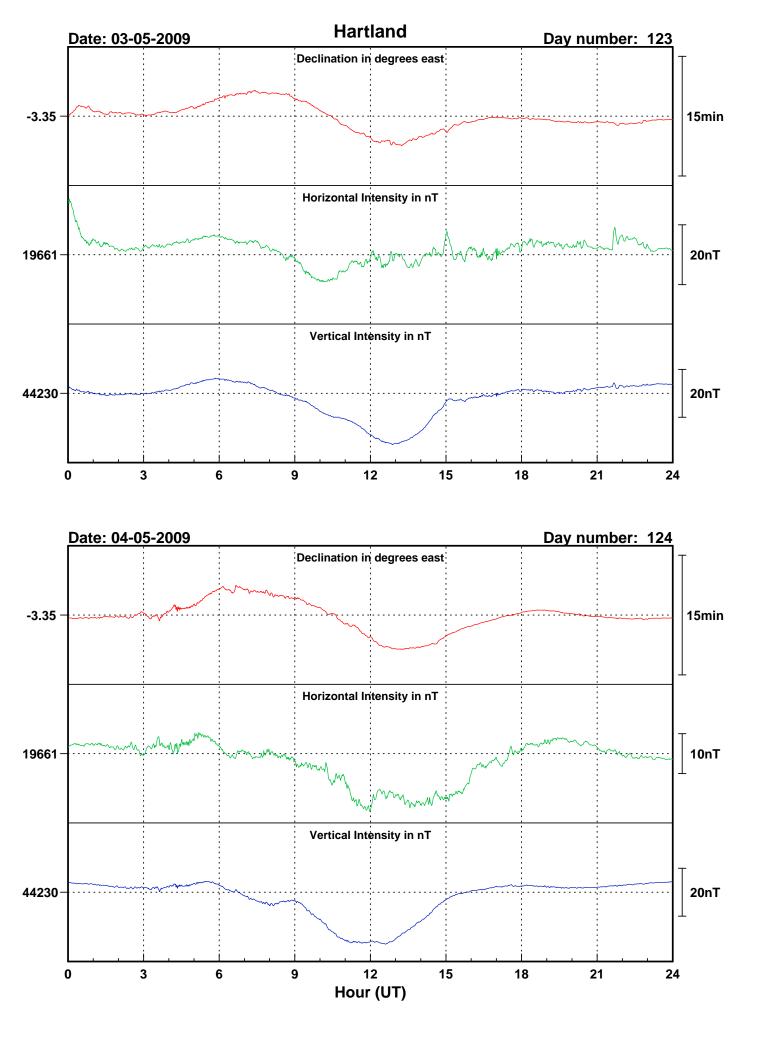
		D	DECLINATION			INCLINATION						
Date	Day Number	Time (UT)	Absolute (°)	Baseline (°)	Time (UT)	Inclination (°)	Total Field Intensity (nT)	H Absolute (nT)	H Baseline (nT)	Z Absolute (nT)	Z Baseline (nT)	Observer
08-May-09	128	08:49	-3.3101	-3.5467	08:57	66.0413	48395.8	19652.5	19666.0	44226.0	44219.7	ST
08-May-09	128	09:04	-3.3140	-3.5467	09:13	66.0425	48393.8	19650.9	19666.6	44224.5	44219.5	ST
14-May-09	134	08:38	-3.3116	-3.5567	09:16	66.0309	48395.5	19660.4	19666.3	44222.1	44219.6	ST
14-May-09	134	08:51	-3.3016	-3.5467	09:00	66.0287	48398.8	19663.4	19666.9	44224.3	44219.3	ST
22-May-09	142	08:09	-3.2872	-3.5483	08:18	66.0414	48395.8	19652.5	19666.3	44225.9	44219.6	ST
22-May-09	142	08:24	-3.2984	-3.5483	08:32	66.0420	48394.1	19651.3	19666.7	44224.6	44219.5	ST
28-May-09	148	09:33	-3.2998	-3.5533	09:41	66.0467	48396.2	19648.5	19667.2	44228.2	44219.2	ST
28-May-09	148	09:47	-3.2963	-3.5467	09:56	66.0558	48389.5	19638.7	19666.6	44225.2	44219.5	ST

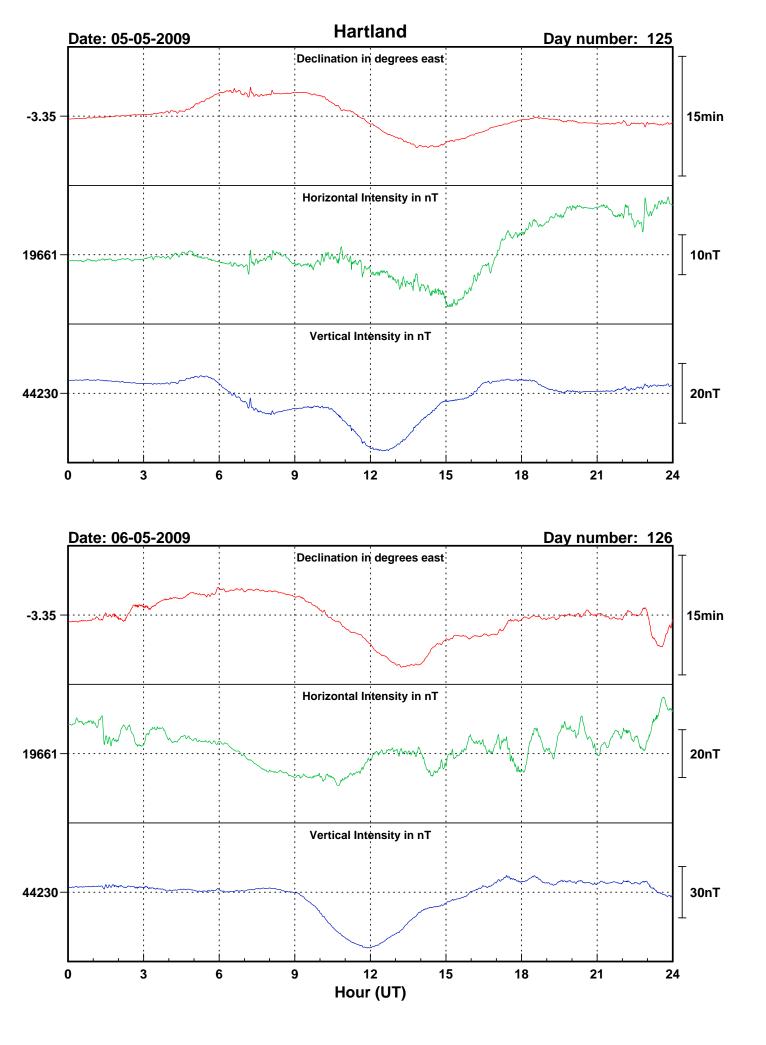


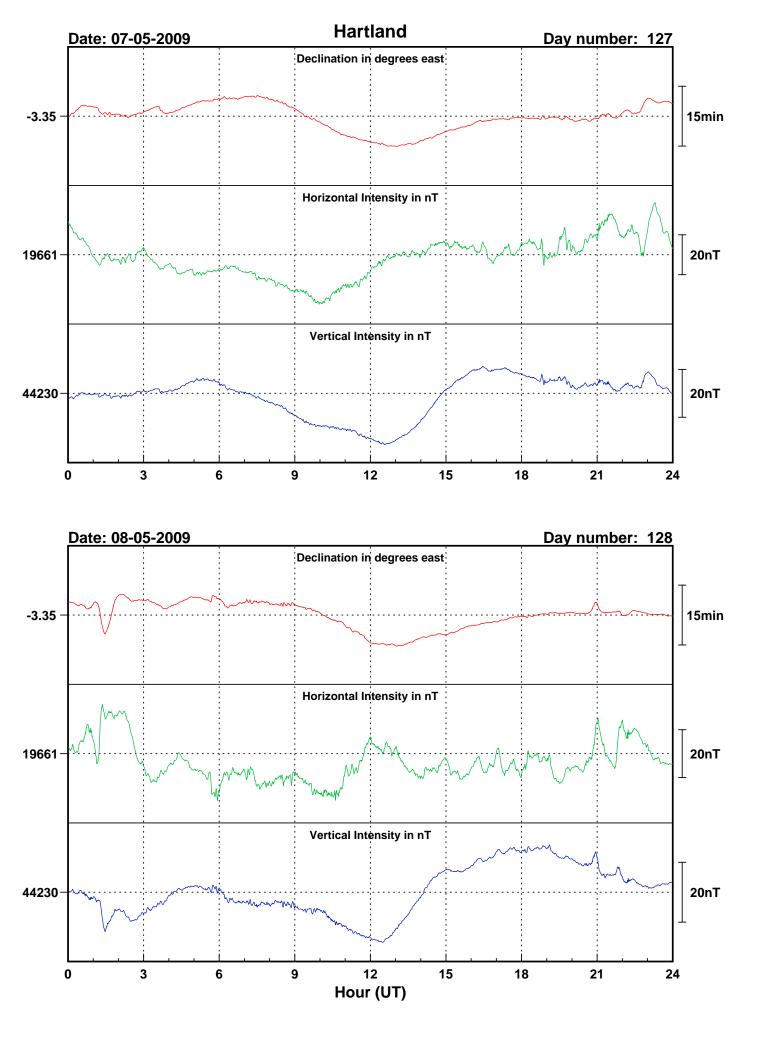


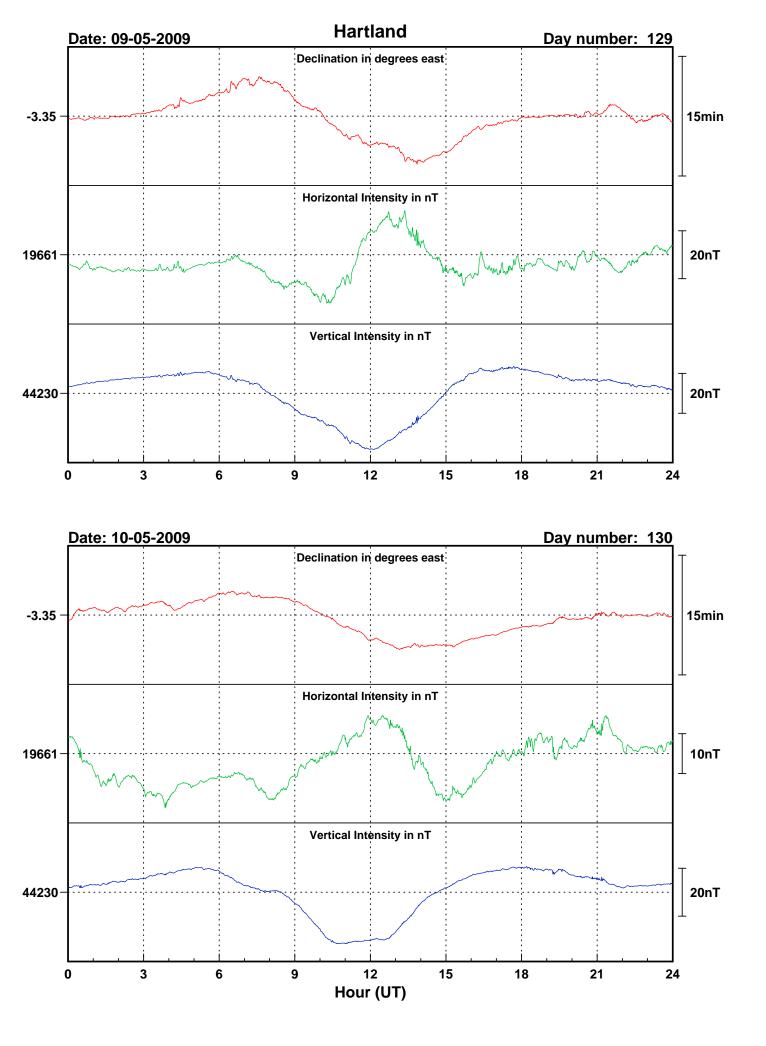


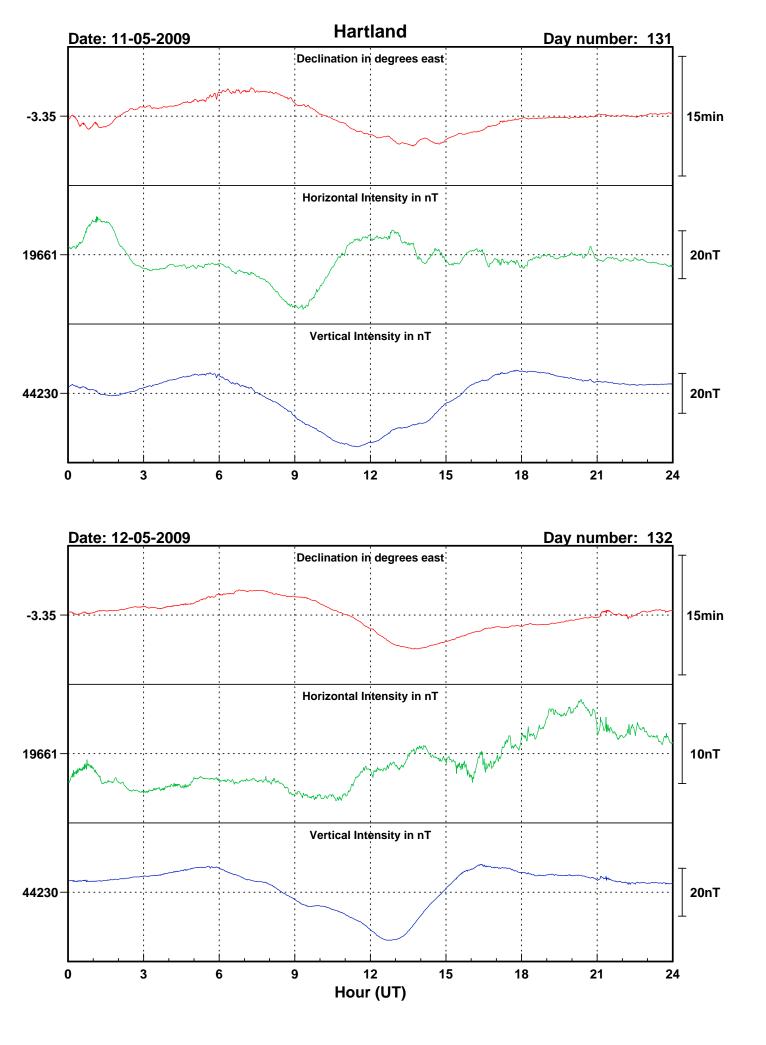


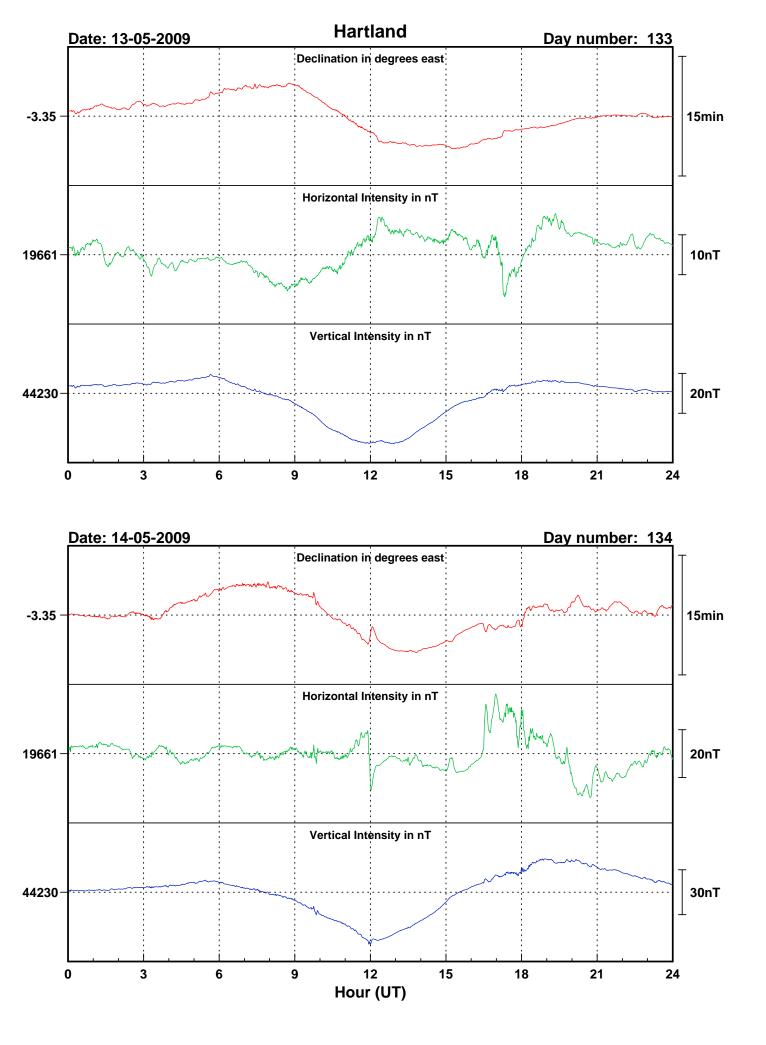


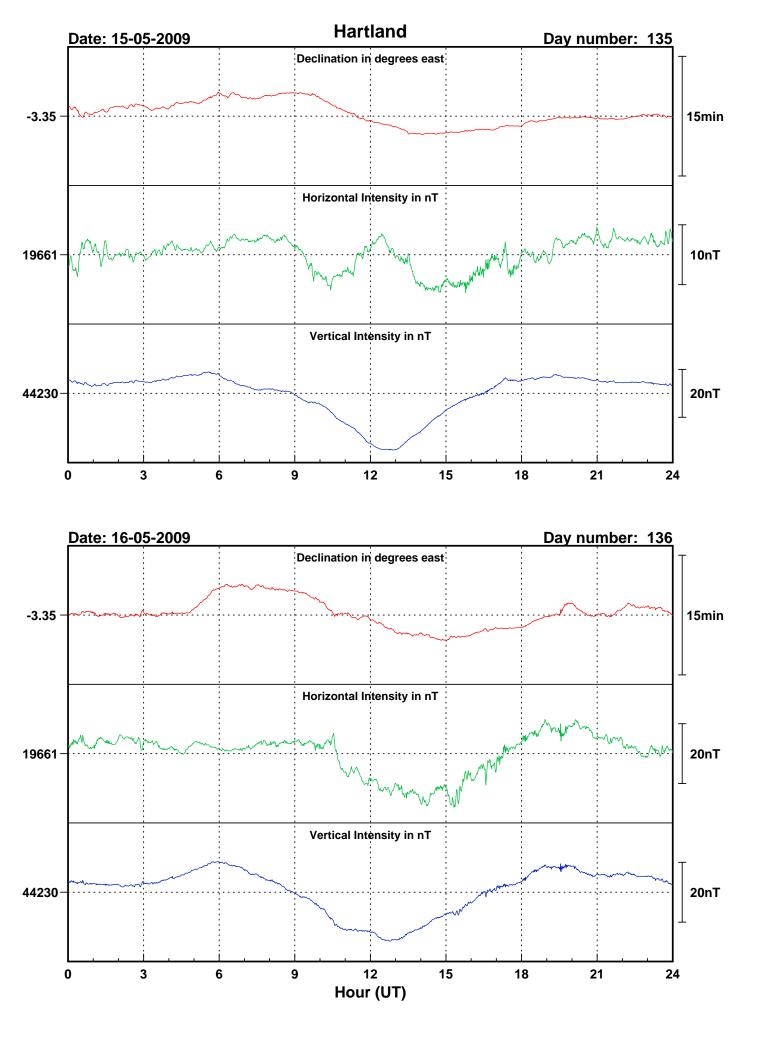


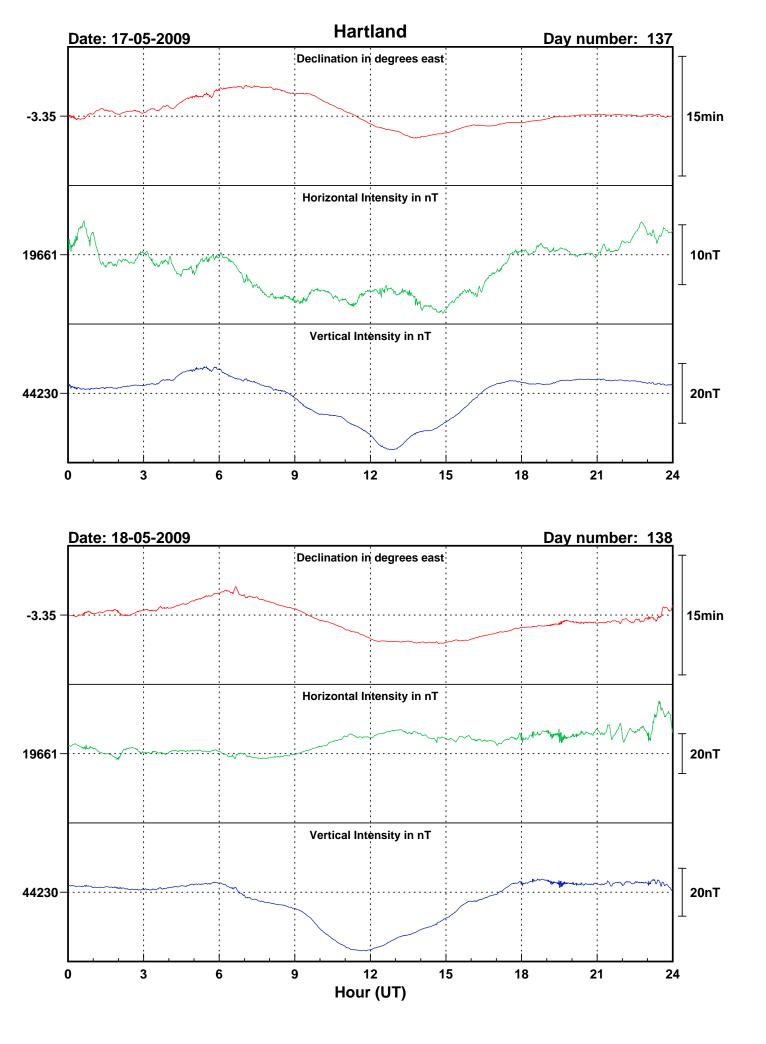


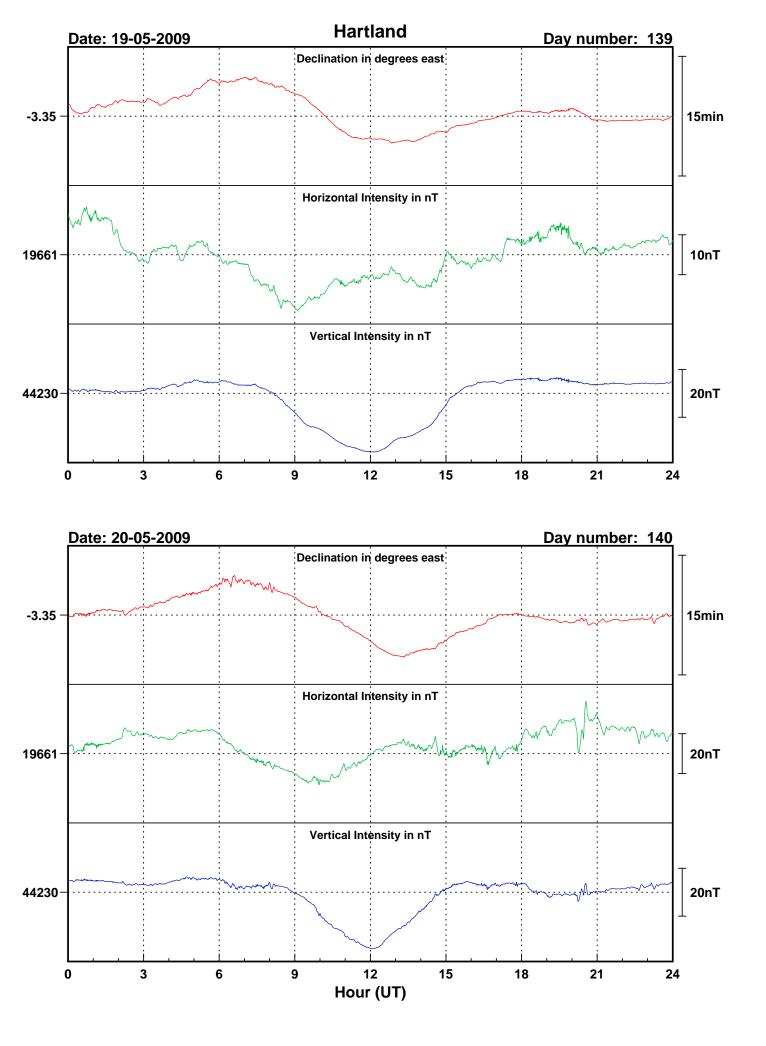


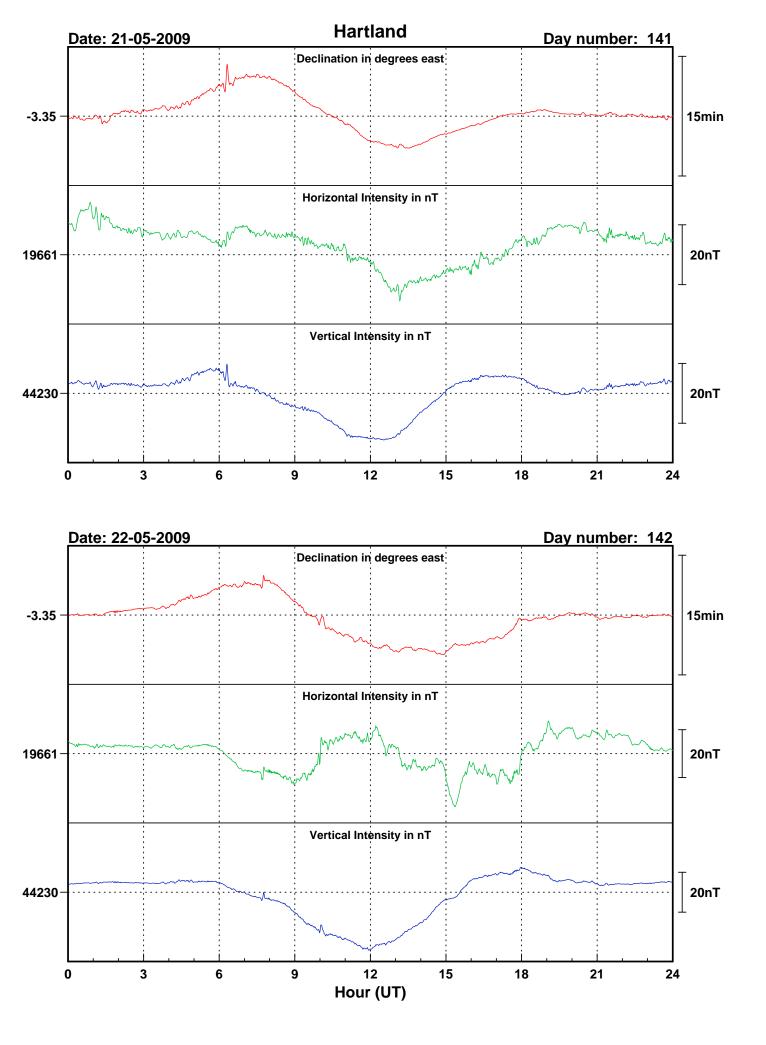


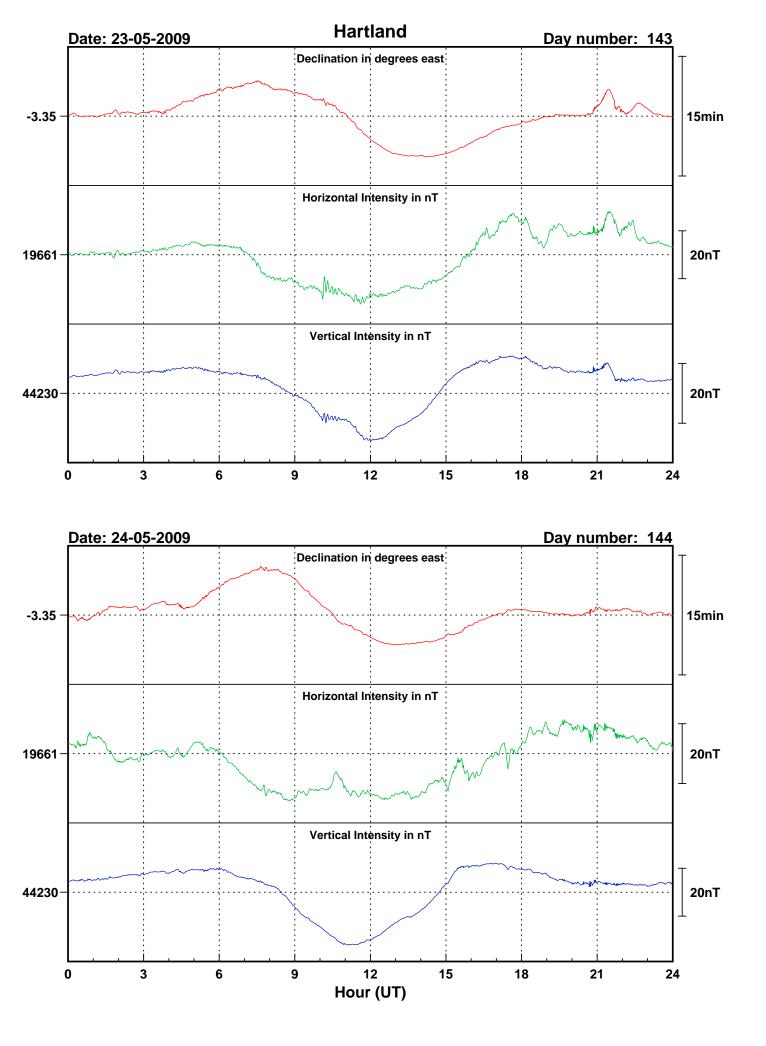


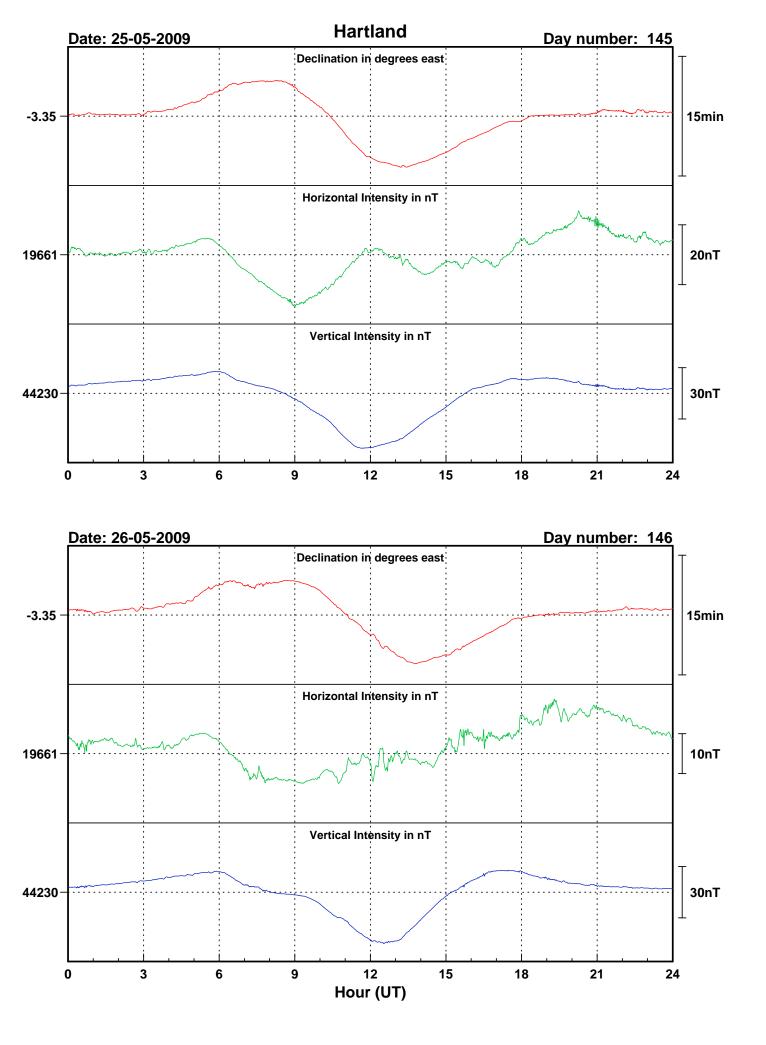


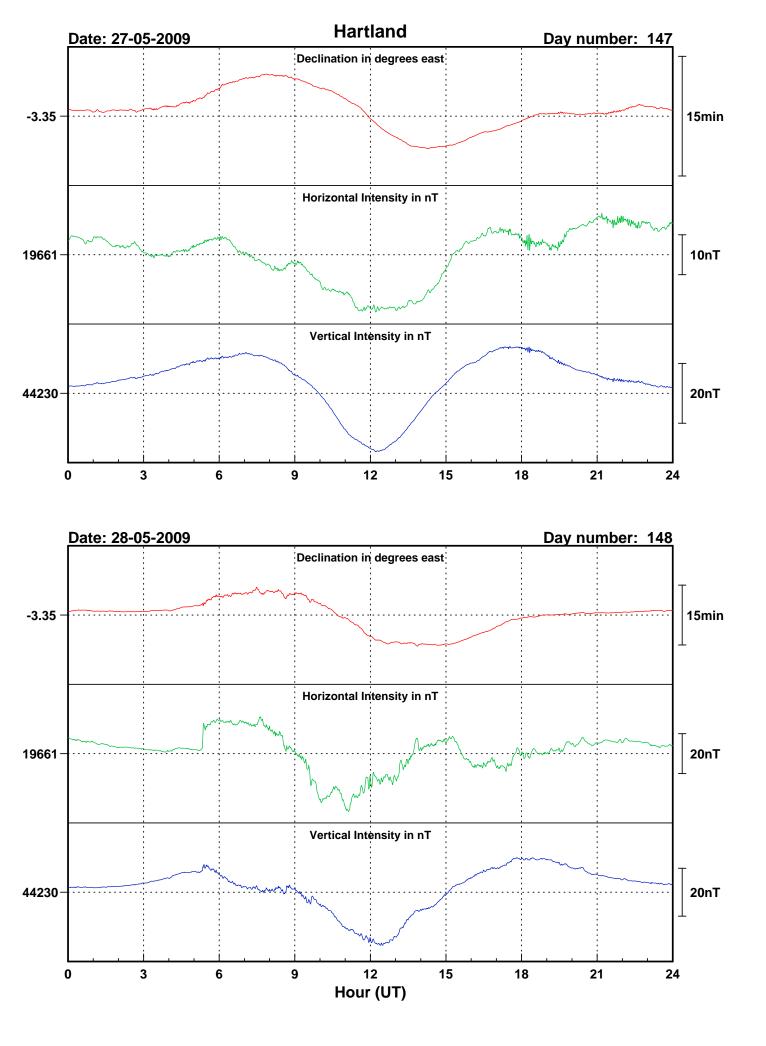


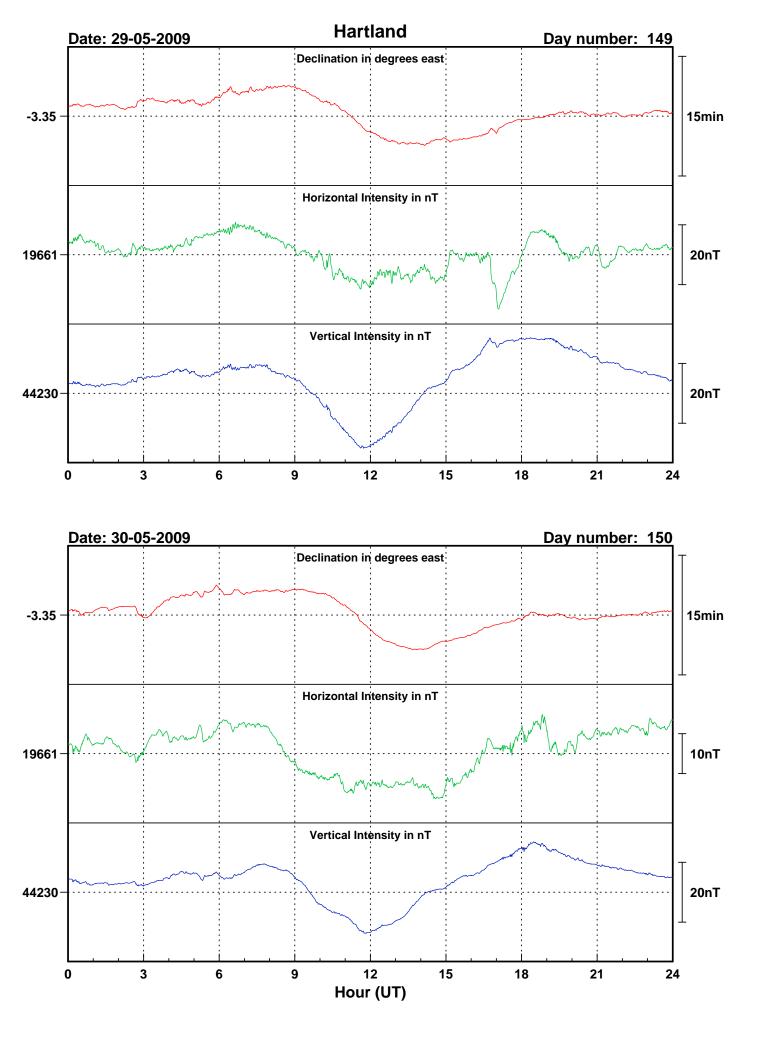


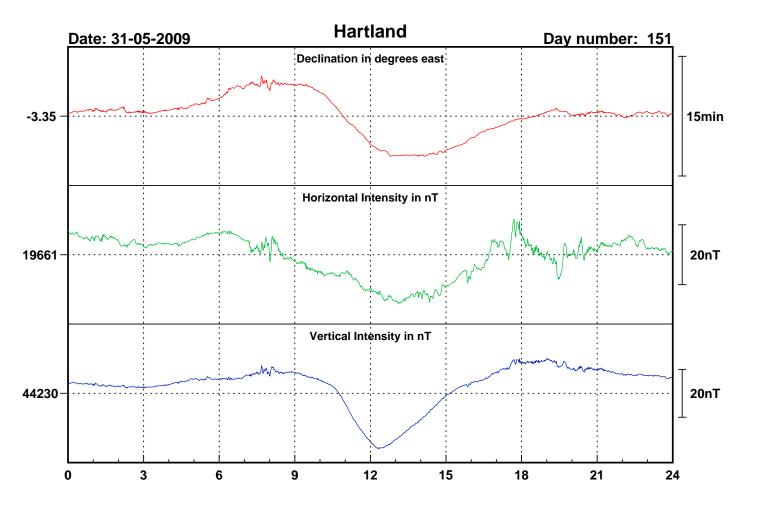




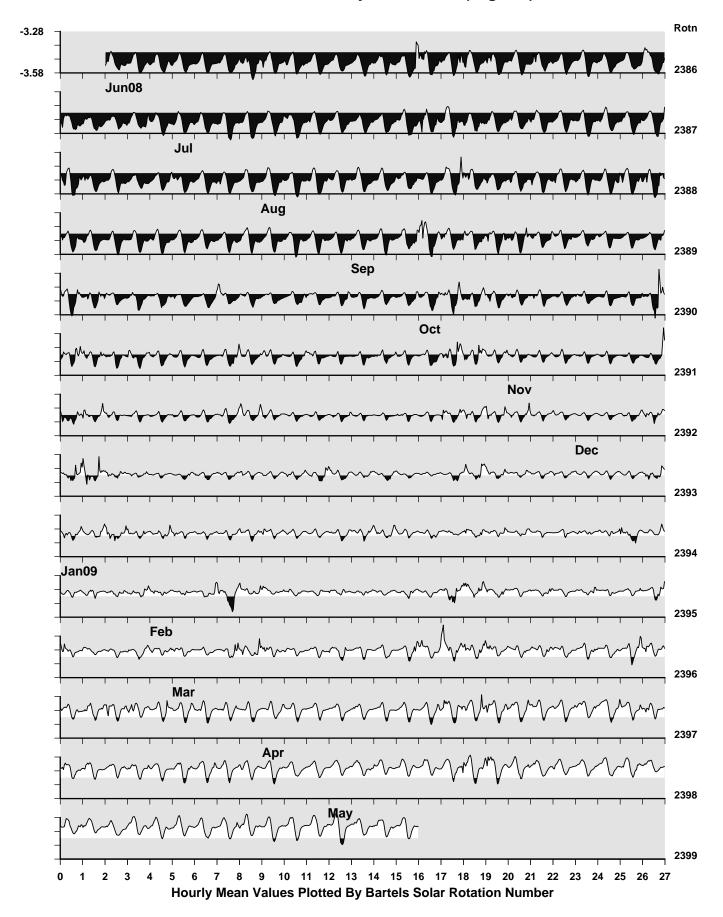




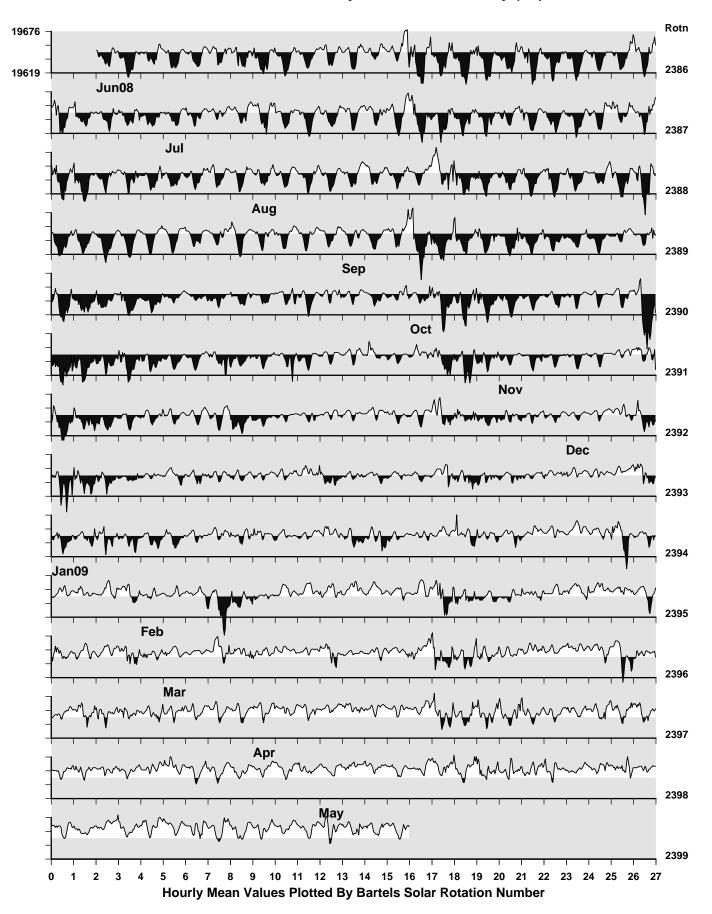




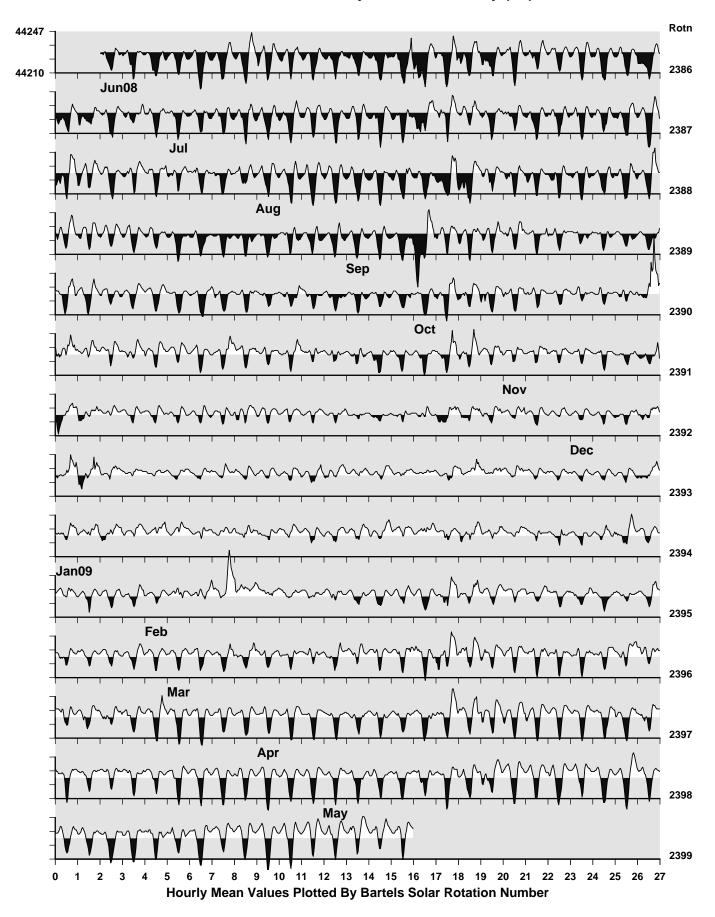
Hartland Observatory: Declination (degrees)

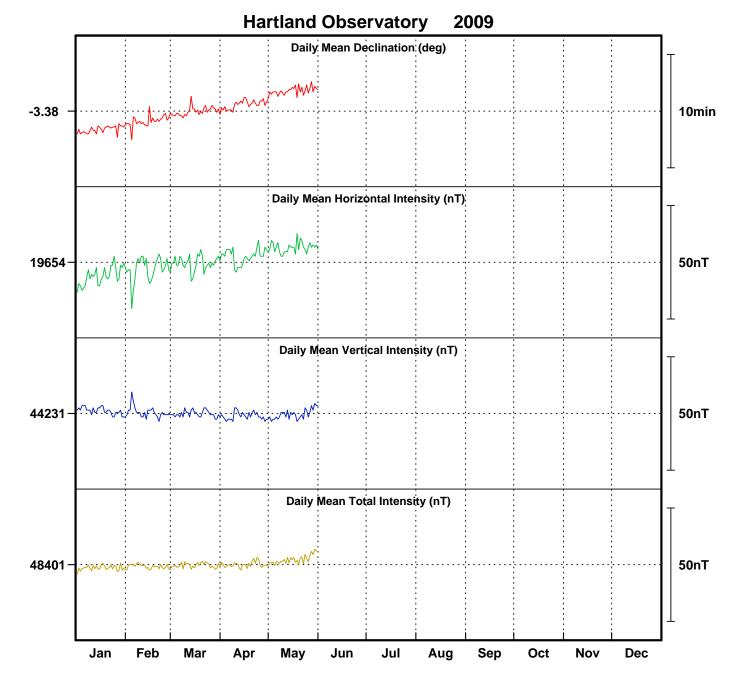


Hartland Observatory: Horizontal Intensity (nT)



Hartland Observatory: Vertical Intensity (nT)





Monthly Mean Values for Hartland Observatory 2009

Month	D	H	Ι	X	Y	Ζ	F
January February March April May	-3° 24.6′ -3° 23.9′ -3° 23.0′ -3° 22.4′ -3° 21.2′	19648 nT 19651 nT 19654 nT 19657 nT 19661 nT	66° 2.9' 66° 2.7' 66° 2.5' 66° 2.3' 66° 2.1'	19614 nT 19617 nT 19620 nT 19623 nT 19627 nT	-1169 nT -1165 nT -1160 nT -1157 nT -1150 nT	44232 nT 44231 nT 44230 nT 44229 nT 44230 nT	48400 nT 48400 nT 48401 nT

Note

i. The values shown here are provisional.

HARTLAND RAPID VARIATIONS

SIs and SSCs

Date	Time (UT)	Туре	Quality	H (nT)	D (min)	Z (nT)
14-05-09	11 53	SI*	С	-9.6	2.16	-2.6
28-05-09	05 18	SSC*	В	12.9	-0.71	1.4

Notes:

An asterisk (*) indicates that the principal impulse was preceded by a smaller reversed impulse. The quality of the event is classified as follows:

A = very distinct

 $\mathbf{B} = \mathbf{fair}$, ordinary, but unmistakable

C = doubtful

The amplitudes given are for the first chief movement of the event.

SFEs

Date	Universal Time			H (nT)	D (min)	Z (nT)
	Start	Maximum	End			
			NONE			

Note:

The amplitudes given are for the first chief movement of the event.

INDICES OF GEOMAGNETIC ACTIVITY

The K Index

Hartl	land Observatory								May 2009		
		K – INDICES FOR THREE-HOUR INTERVAL									
Day	00-03	03-06	06-09	09-12	12-15	15-18	18-21	21-24	SUM		
1	0	0	1	1	0	1	0	1	4		
2	0	1	1	1	2	1	1	2	9		
3	2	0	0	0	1	2	0	1	6		
4	0	1	1	1	1	0	0	0	4		
5	0	1	1	1	1	1	0	1	6		
6	2	1	0	1	2	2	3	3	14		
7	3	2	1	1	1	2	2	3	15		
8	4	2	2	2	2	2	3	3	20		
9	0	1	1	2	3	2	1	2	12		
10	1	1	1	0	1	1	1	1	7		
11	2	1	0	0	1	1	1	0	6		
12	0	0	0	0	0	1	0	1	2		
13	1	1	0	0	1	2	1	0	6		
14	1	1	1	3	2	3	3	1	15		
15	1	1	1	0	0	1	0	0	4		
16	1	1	1	2	1	1	2	1	10		
17	1	1	0	0	0	0	0	0	2		
18	1	0	1	0	0	1	1	2	6		
19	1	1	0	0	1	1	1	0	5		
20	1	1	1	1	1	2	3	2	12		
21	2	1	2	1	1	1	1	1	10		
22	0	0	1	2	2	2	2	0	9		
23	0	0	1	1	0	1	2	2	7		
24	1	1	0	1	0	2	1	0	6		
25	0	0	0	0	0	1	1	0	2		
26	0	1	1	1	1	1	1	0	6		
27	0	0	0	0	0	0	1	1	2		
28	0	2	2	2	2	2	1	0	11		
29	1	1	1	1	1	2	2	1	10		
30	1	2	1	0	1	1	1	0	7		
31	0	0	2	0	0	2	2	1	7		

Lower bound (nT) for the range for each index value at Hartland Observatory										
K-Index										
0	1	2	3	4	5	6	7	8	9	
0	5	10	20	40	70	120	200	330	500	

The aa Index

Date	Day	K-North	K-South	(a)	(b)	(c)	(d)	(e)
01-05-09	121	00110101	00111100	5	5	5	5	5
02-05-09	122	01112112	01112000	9	6	6	9	8
03-05-09	123	$2\ 0\ 0\ 0\ 1\ 2\ 0\ 1$	$1\ 0\ 0\ 0\ 0\ 0\ 0$	7	3	5	5	5
04-05-09	124	01111000	$0\;1\;1\;2\;2\;0\;0\;0$	5	7	8	5	6
05-05-09	125	01111101	$0\ 0\ 0\ 0\ 0\ 1\ 0\ 0$	6	3	4	5	5
06-05-09	126	21012233	1 1 1 1 1 2 2 1	16	10	8	18	13
07-05-09	127	3 2 1 1 1 2 2 3	22121212	17	13	15	15	15
08-05-09	128	4 2 2 2 2 2 3 3	3 2 3 1 2 2 2 1	25	18	24	19	22
09-05-09	129	01123212	01132210	13	12	11	14	12
10-05-09	130	11101111	$0\ 0\ 0\ 1\ 2\ 0\ 1\ 0$	7	6	5	8	6
11-05-09	131	21001110	$1\ 1\ 1\ 2\ 1\ 2\ 0\ 1$	7	9	9	8	8
12-05-09	132	$0\ 0\ 0\ 0\ 0\ 1\ 0\ 1$	$0\ 0\ 0\ 0\ 0\ 1\ 0\ 0$	4	3	2	4	3
13-05-09	133	$1\ 1\ 0\ 0\ 1\ 2\ 1\ 0$	$0\ 0\ 0\ 0\ 0\ 1\ 0\ 0$	7	3	4	6	5
14-05-09	134	1 1 1 3 2 3 3 1	0 1 1 2 2 2 2 1	18	11	11	18	15
15-05-09	135	$1\ 1\ 1\ 0\ 0\ 1\ 0\ 0$	$1\ 0\ 0\ 0\ 0\ 1\ 0\ 0$	5	4	5	4	4
16-05-09	136	11121121	$1\ 1\ 0\ 1\ 1\ 1\ 1\ 0$	10	7	8	8	8
17-05-09	137	$1\ 1\ 0\ 0\ 0\ 0\ 0$	$0\ 1\ 0\ 0\ 0\ 0\ 0$	4	3	4	2	3
18-05-09	138	$1\ 0\ 1\ 0\ 0\ 1\ 1\ 2$	$0\ 0\ 0\ 0\ 0\ 0\ 0\ 1$	7	3	4	6	5
19-05-09	139	11001110	$1\ 1\ 2\ 2\ 1\ 0\ 0\ 0$	6	8	9	5	7
20-05-09	140	1 1 1 1 1 2 3 2	22100111	13	9	9	12	11
21-05-09	141	21211111	$1\ 0\ 0\ 1\ 1\ 0\ 0\ 0$	10	5	9	6	7
22-05-09	142	0 0 1 2 2 2 2 0	0 0 1 2 2 2 2 1	10	11	7	13	10
23-05-09	143	00110122	11210111	8	8	8	9	8
24-05-09	144	11010210	$1\ 1\ 0\ 1\ 0\ 0\ 0\ 0$	7	5	6	5	6
25-05-09	145	00000110	000000000	4	2	2	4	3
26-05-09	146	01111110	00101000	6	4	5	5	5
27-05-09	147	00000011	000000000	4	2	2	4	3
28-05-09	148	0 2 2 2 2 2 1 0	03222000	12	11	15	8	11
29-05-09	149	1 1 1 1 1 2 2 1	01122110	10	9	8	10	9
30-05-09	150	12101110	11010110	7	6	8	6	7
31-05-09	151	0 0 2 0 0 2 2 1	00000110	8	4	4	8	6
	Monthl	y mean value =	7.	8				

(a)

The northern daily mean value, Aa_n The southern daily mean value, Aa_s (b)

The mean value of aa for the interval 00-12 UT (c)

The mean value of aa for the interval 12-24 UT (d)

The daily mean value of aa (Aa) (e)

<u>Notes</u>

i. The values are rounded to the nearest integer.

The units of the aa index are nT. ii.

iii. The values shown here are provisional. The definitive values are computed and published by the International Service for Geomagnetic Indices, Paris