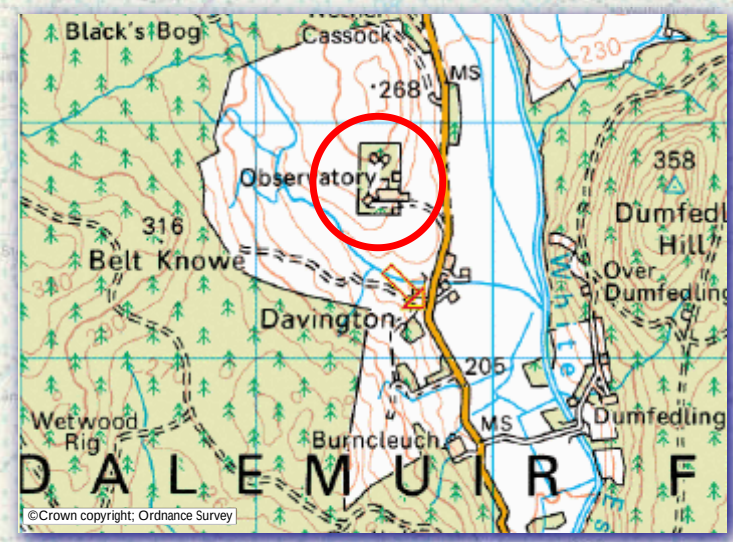
BRITISH GEOLOGICAL SURVEY

Eskdalemuir Observatory

Monthly Magnetic Bulletin

October 2010

IO/IO/ES



British Geological Survey
NATURAL ENVIRONMENT RESEARCH COUNCIL

ESKDALEMUIR OBSERVATORY MAGNETIC DATA

1. Introduction

Eskdalemuir observatory is one of three geomagnetic observatories in the UK operated and maintained by the British Geological Survey (BGS).

This bulletin is published to provide rapid access to the provisional geomagnetic observatory results. The information is freely available for personal, academic, educational and non-commercial research or use. Magnetic observatory data are presented as a series of plots of one-minute, hourly and daily values, followed by tabulations of monthly values, reports of rapid variations and geomagnetic activity indices. The operation of the observatory and presentation of data are described in the rest of this section.

Enquiries about the data should be addressed to:

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Internet: www.geomag.bgs.ac.uk

2. Position

The observatory is situated on a rising shoulder of open moorland in the upper part of the valley of the White Esk River in the Southern Uplands of Scotland. The observatory co-ordinates are:

Geographic:	*55.317 °N*	*356.800 °E*
Geomagnetic:	*57.682 °N*	*83.756 °E*
Height above mean sea level:	*245 m*	

The geomagnetic co-ordinates are approximations, calculated using the 11th generation International Geomagnetic Reference Field (IGRF) at epoch 2010.5. On-line access to models (including IGRF), charts and navigational data are available at www.geomag.bgs.ac.uk/navigation.html

3. The Observatory Operation

3.1 GDAS

The observatory operates under the control of the Geomagnetic Data Acquisition System (GDAS), which was developed by BGS staff, installed in 2002, and became fully operational in January 2003. The data acquisition software, running on QNX operated computers, controls the data logging and the communications.

There are two sets of sensors used for making magnetic measurements. A tri-axial linear-core fluxgate magnetometer, manufactured by the Danish Meteorological Institute, is used to measure the variations in the horizontal (*H*) and vertical (*Z*) components of the field. The third sensor is oriented perpendicular to these, and measures variations, which are proportional to the changes in declination (*D*). Measurements are made at a rate of 1 Hz.

In addition to the fluxgate sensors there is a proton precession magnetometer (PPM) making measurements of the absolute total field intensity (*F*) at a rate of 0.1Hz.

The raw unfiltered data are retrieved automatically via Internet connections to the BGS office in Edinburgh in near real-time. The fluxgate data are filtered to produce one-minute values using a 61-point cosine filter and the total field intensity samples are filtered using a 7-point cosine filter. The one-minute values provide input for various data products in the Geomagnetism Information and Forecast Service (GIFS), available on-line at www.geomag.bgs.ac.uk/on_line_gifs.html

3.2 Back-up Systems

There are two other fully independent identical systems, GDAS 2 and GDAS 3, operating at the observatory. The data from these are also processed in near real-time and used for quality control purposes. They are also used to fill any gaps or replace any corrupt values in the primary system, GDAS 1.

3.3 Absolute Observations

The GDAS fluxgate magnetometers accurately measure variations in the components of the geomagnetic field, but not the absolute magnitudes. Two sets of absolute measurements of the field are made manually once per week. A fluxgate sensor mounted on a theodolite is used to determine *D* and inclination (*I*); the GDAS PPM measurements, with a site difference correction applied, are used for *F*. The absolute observations are used in conjunction with the GDAS variometer measurements to produce a continuous record of the absolute values of the geomagnetic field

elements as if they had been measured at the observatory reference pillar.

4. Observatory Results

The data presented in the bulletin are in the form of plots and tabulations described in the following sections.

4.1 Absolute Observations

The absolute observation measurements made during the month are tabulated. Also included are the corresponding baseline values, which are the differences between the absolute measurements and the variometer measurements of *D*, *H* and *Z* (in the sense absolute–variometer). These are also plotted (markers) along with the derived preliminary daily baseline values (line) throughout the year. Daily mean differences between the measured absolute *F* and the *F* computed from the baseline corrected *H* and *Z* values are plotted in the fourth panel (in the sense measured–derived). The bottom panel shows the daily mean temperature in the fluxgate chamber.

4.2 Summary magnetograms

Small-scale magnetograms are plotted which allow the month's data to be viewed at a glance. They are plotted 16 days to a page and show the one-minute variations in *D*, *H* and Z. The scales are shown on the right-hand side of the page. On disturbed days the scales are multiplied by a factor, which is indicated above the panel for that day. The variations are centred on the monthly mean value, shown on the left side of the page.

4.3 Magnetograms

The daily magnetograms are plotted using one-minute values of *D*, *H* and *Z* from the fluxgate sensors, with any gaps filled using back-up data. The magnetograms are plotted to a variable scale; scale bars are shown to the right of each plot. The absolute level (the monthly mean value) is indicated on the left side of the plots.

4.4 Hourly Mean Value Plots

Hourly mean values of *D*, *H* and *Z* for the past 12 months are plotted in 27-day segments corresponding to the Bartels solar rotation number. Magnetic disturbances associated with active regions and/or coronal holes on the Sun may recur after 27 days: the same is true for geomagnetically quiet intervals. Plotting the data in this way highlights this recurrence. Diurnal variations are also clear in these plots and the amplitude changes throughout the year highlight the seasonal changes. Longer term secular variation is also illustrated.

Full lists of the UK observatory hourly mean values from 1983 to the present day are available at www.geomag.bgs.ac.uk/gifs/hourly_means.html.

4.5 Daily and Monthly Mean Values

Daily mean values of *D*, *H*, *Z* and *F* are plotted throughout the year. In addition, a table of monthly mean values of all the geomagnetic elements is provided. These values depend on accurate specification of the fluxgate sensor baselines. It is anticipated that these provisional values will not be altered by more than a few Tortenths of arcminutes before being made definitive at the end of the year.

4.6 Rapid Variations

Charged particles stream from the Sun in the solar wind. The solar wind interacts with the geomagnetic field to create a cavity, the magnetosphere, in which the field is confined. When a region of enhanced velocity and/or density in the solar wind arrives at the dayside boundary of the magnetosphere (at about 10 earth radii) the boundary is pushed towards the Earth. Currents set up on the boundary of the magnetosphere can cause an abrupt change in the geomagnetic field measured on the ground and this is recorded on observatory magnetograms as a sudden impulse (*si*). If, following an *si*, there is a change in the rhythm of activity, the *si* is termed a storm sudden commencement (*ssc*). A classical magnetic storm exhibiting initial, main and recovery phases (shown by, for instance, the *Dst* ring current index) can often occur after a *ssc*, in which case the start of the storm is taken as the time of the *ssc*.

Solar flares, seen at optical wavelengths as a sudden brightening of a small region of the Sun's surface, are also responsible for increased X-ray emissions. These X-rays cause increased ionisation in the ionosphere, which leads to absorption of short-wave radio signals. A solar flare effect (*sfe*), or "crochet", may be observed on a magnetogram during geomagnetically quiet times. It is a relatively short-term change (tens of minutes) to the normal diurnal variation and can vary in size (tens of nT) depending on local time (LT), geomagnetic latitude and solar zenith angle.

4.7 Local geomagnetic activity indices

The Observatory K index. This summarises geomagnetic activity at an observatory by assigning a code, an integer in the range 0 to 9, to each 3-hour Universal Time (UT) interval. The index for each 3-hour UT interval is determined from the maximum range in *H* or *D* (scaled in nT), with allowance made for the regular (undisturbed) diurnal variation. The conversion from range to an index value is made using a quasi-logarithmic scale, with the scale values dependent on the geomagnetic latitude of the observatory. The lower bounds (in nT) for the classification of each period at Eskdalemuir are:

0	1	2	3	4	5	6	7	8	9
0	8	15	30	60	105	180	300	500	750

The *K* index retains the LT and seasonal dependence of activity associated with the position of the observatory. The 3-hourly *K* indices for the month are tabulated and also plotted as a histogram. All UK observatory *K* indices are available at www.geomag.bgs.ac.uk/gifs/k_indices.html

4.8 Global geomagnetic activity indices

The aa index. A number of 3-hour geomagnetic indices are computed by combining *K* indices from networks of observatories to characterise global activity levels and to eliminate LT and seasonal effects. The simplest of these is the *aa* index, computed using the *K* indices from two approximately antipodal observatories: Hartland in the UK and Canberra in Australia. The *aa* index is calculated from linearisations of the Hartland and Canberra *K* indices, and has units of nT. The 3-hourly *aa* indices are tabulated along with the daily mean value of *aa* (denoted *Aa*), the mean values of *aa* for the intervals 00-12UT (Aa_{am}) and 12-24UT (Aa_{pm}) and the monthly mean value. The 3-hourly *aa* indices for the month are also plotted as a histogram.

Although the *aa* index is based on data from only two observatories, provided averages over 12 hours or longer are used, the index is strongly correlated with the *ap* and *am* indices, which are derived using data from more extensive observatory networks.

The *aa* indices listed in this bulletin are available at www.geomag.bgs.ac.uk/gifs/aaindex.html as well as the full data set from 1868.

Definitive *aa* are published by the International Service for Geomagnetic Indices, LATMOS, 4 Avenue de Neptune, F-94107 Saint Maur Cedex, France.

5. Conditions of Use

The data presented in this bulletin are provided for personal, academic, educational, non-commercial research or other non-commercial use and are not for sale or distribution to third parties without written permission from BGS.

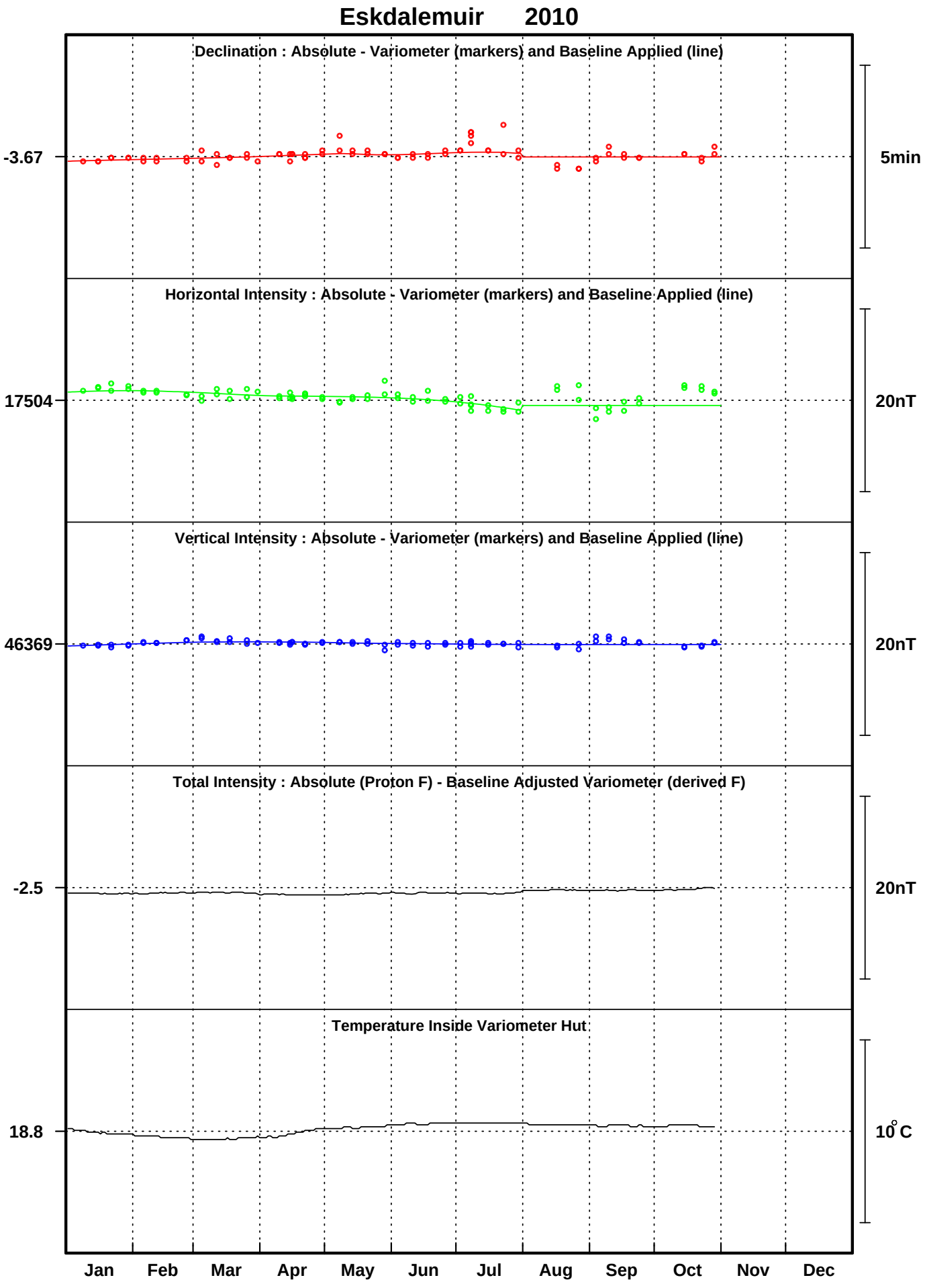
Reproduction of any part of this bulletin should be accompanied by the statement: 'Reproduced with the permission of the British Geological Survey ©NERC. All rights Reserved'. Publications making use of the data should include an acknowledgment statement of the form: 'The results presented in this paper rely on the data collected at Eskdalemuir magnetic observatory, operated by the British Geological Survey.'

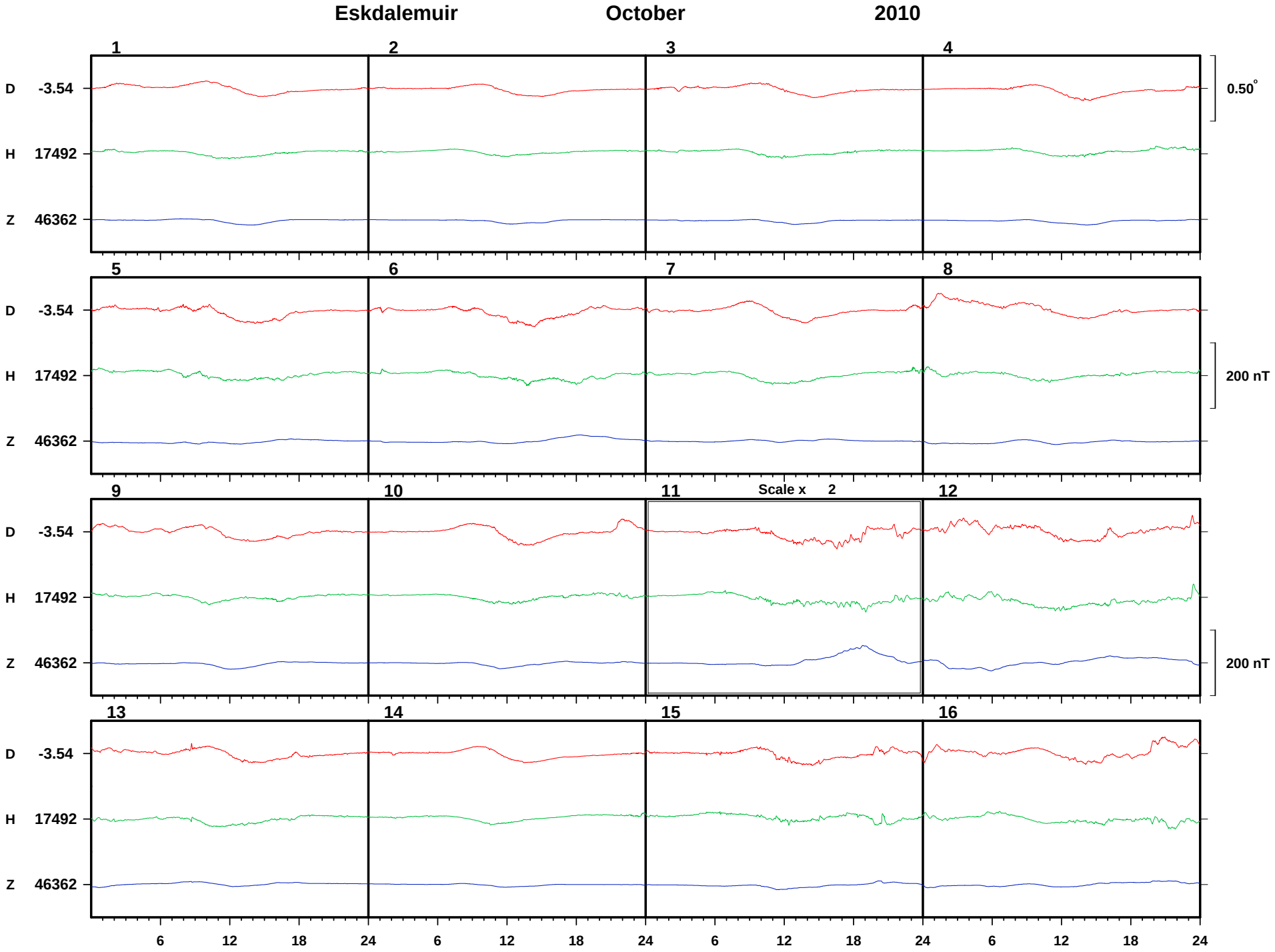
Commercial users can contact the geomagnetism team for information on the range of applications and services offered. Full contact details are available at www.geomag.bgs.ac.uk/staff.html.

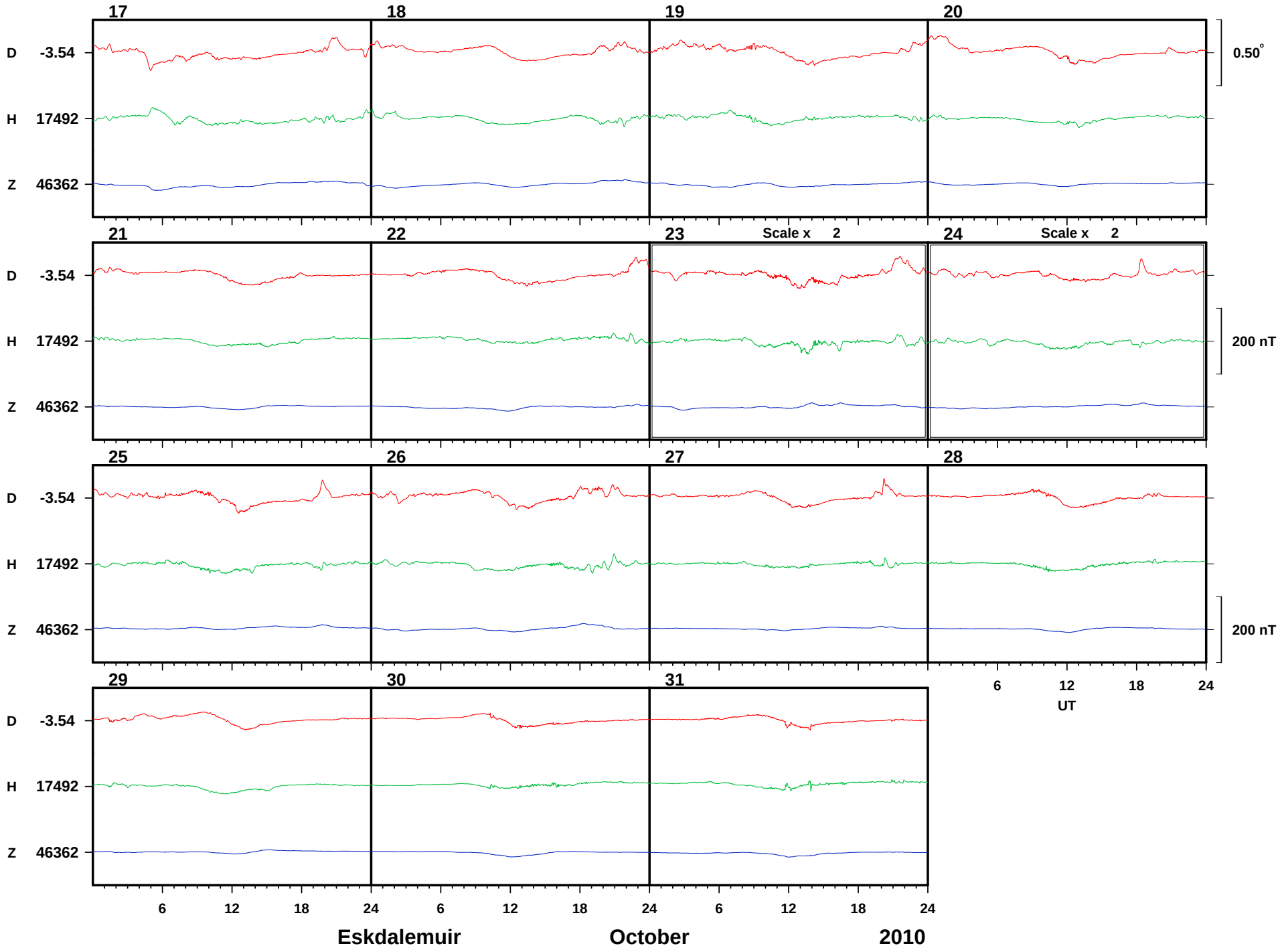
ESKDALEMUIR OBSERVATORY

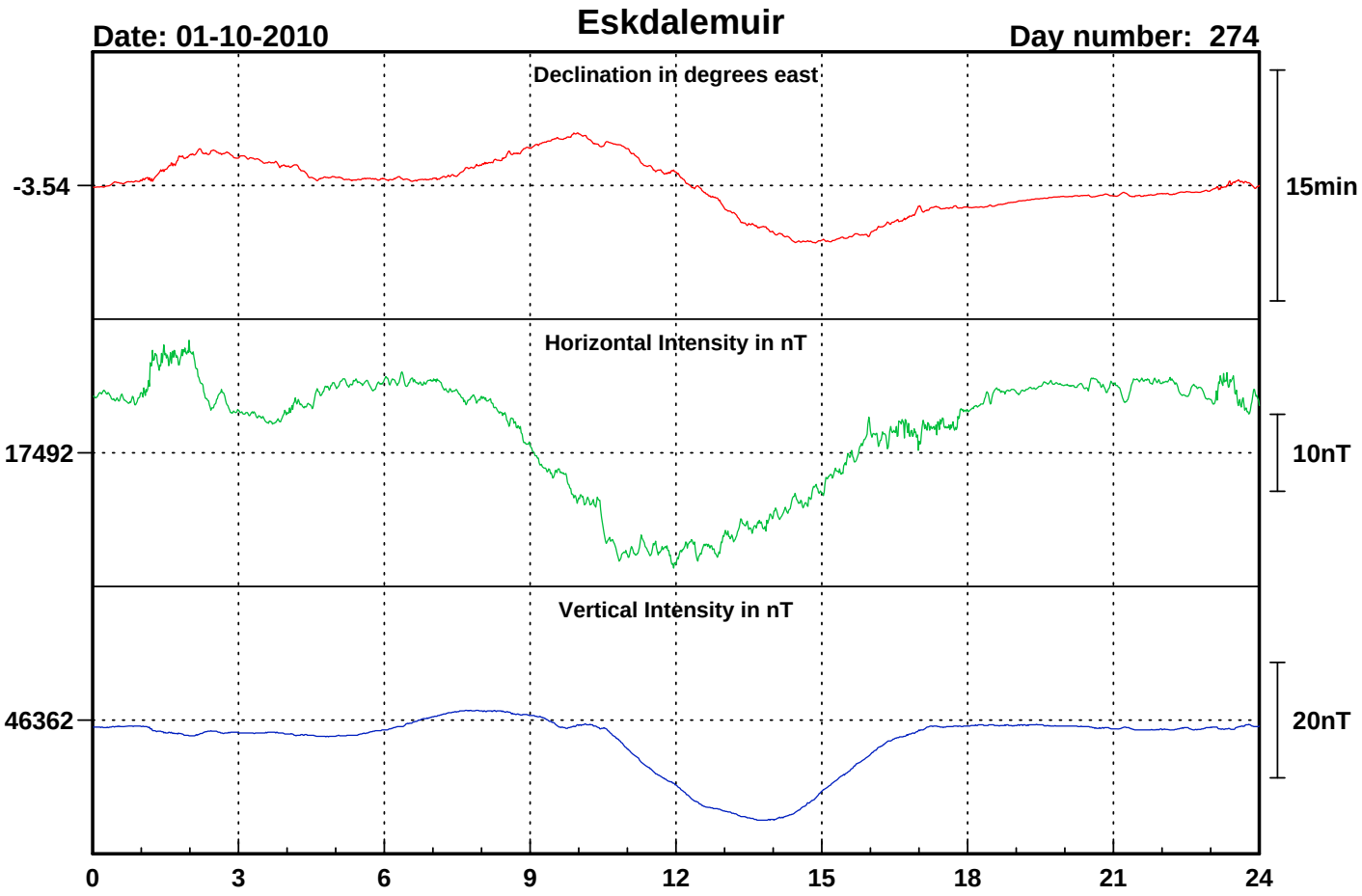
ABSOLUTE OBSERVATIONS

		Declination			Inclination		Total Field		Horizontal Intensity		Vertical Intensity		
Date	**Day Number**	**Time (UT)**	**Absolute (º)**	**Baseline (º)**	**Time (UT)**	**Absolute (º)**	**Site difference (nT)**	**Absolute corrected (nT)**	**Absolute (nT)**	**Baseline (nT)**	**Absolute (nT)**	**Baseline (nT)**	**Observer**
14-Oct-10	287	11:08	-3.5272	-3.6733	11:19	69.3383	3.2	49541.0	17480.5	17505.4	46354.5	46368.8	CP
14-Oct-10	287	11:29	-3.5451	-3.6733	11:39	69.3371	3.2	49540.9	17481.5	17505.1	46354.0	46368.9	CP
22-Oct-10	295	09:46	-3.5097	-3.6750	09:57	69.3233	3.2	49547.3	17494.8	17505.3	46355.9	46368.9	CP
22-Oct-10	295	10:07	-3.5071	-3.6767	10:18	69.3300	3.2	49545.2	17488.7	17504.9	46355.9	46369.0	CP
28-Oct-10	301	14:55	-3.5753	-3.6700	15:06	69.3344	3.2	49555.7	17488.9	17504.5	46367.1	46369.4	CP
28-Oct-10	301	15:17	-3.5654	-3.6733	15:28	69.3325	3.2	49557.7	17491.1	17504.7	46368.4	46369.3	CP


Eskdalemuir 2010
Declination : Absolute - Variometer (markers) and Baseline Applied (line)
-3.67
5min
Horizontal Intensity : Absolute - Variometer (markers) and Baseline Applied (line)
17504
20nT
Vertical Intensity : Absolute - Variometer (markers) and Baseline Applied (line)
46369
20nT
Total Intensity : Absolute (Proton F) - Baseline Adjusted Variometer (derived F)
-2.5
20nT
Temperature Inside Variometer Hut
18.8
10°C
Jan
Feb
Mar
Apr
May
Jun
Jul
Aug
Sep
Oct
Nov
Dec


Eskdalemuir
October
2010
1
2
3
4
D -3.54
H 17492
Z 46362
0.50°
5
6
7
8
200 nT
9
10
11
Scale x 2
12
13
14
15
16
6
12
18
24


17
18
19
20
D -3.54
H 17492
Z 46362
0.50°
21
22
23
Scale x 2
24
Scale x 2
200 nT
25
26
27
28
200 nT
29
30
31
6
12
18
24
UT
Eskdalemuir
October
2010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15min
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10nT
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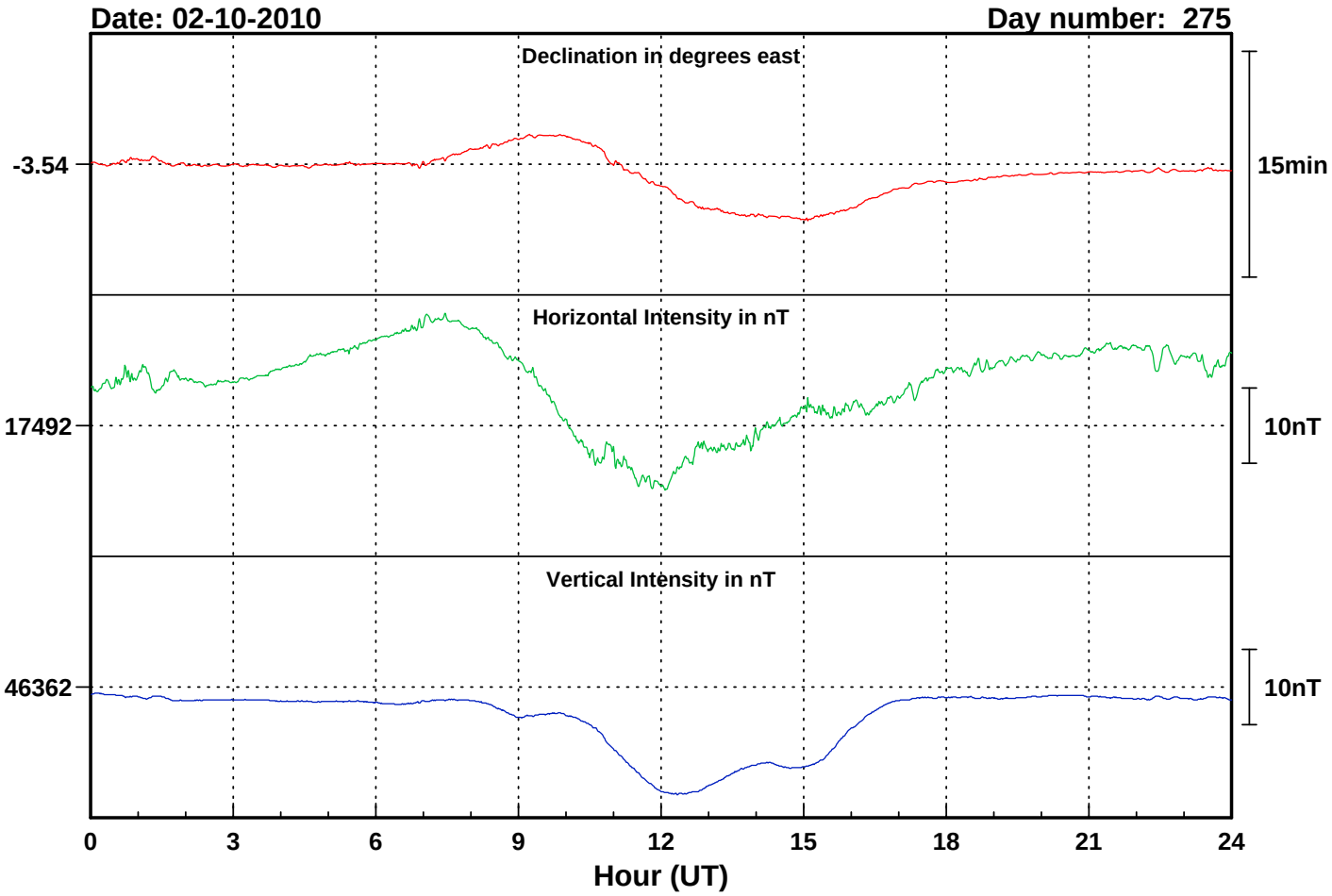
0
3
6
9
12
15
18
21
24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Day number: 275
Declination in degrees east
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15min
Horizontal Intensity in nT
17492
10nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
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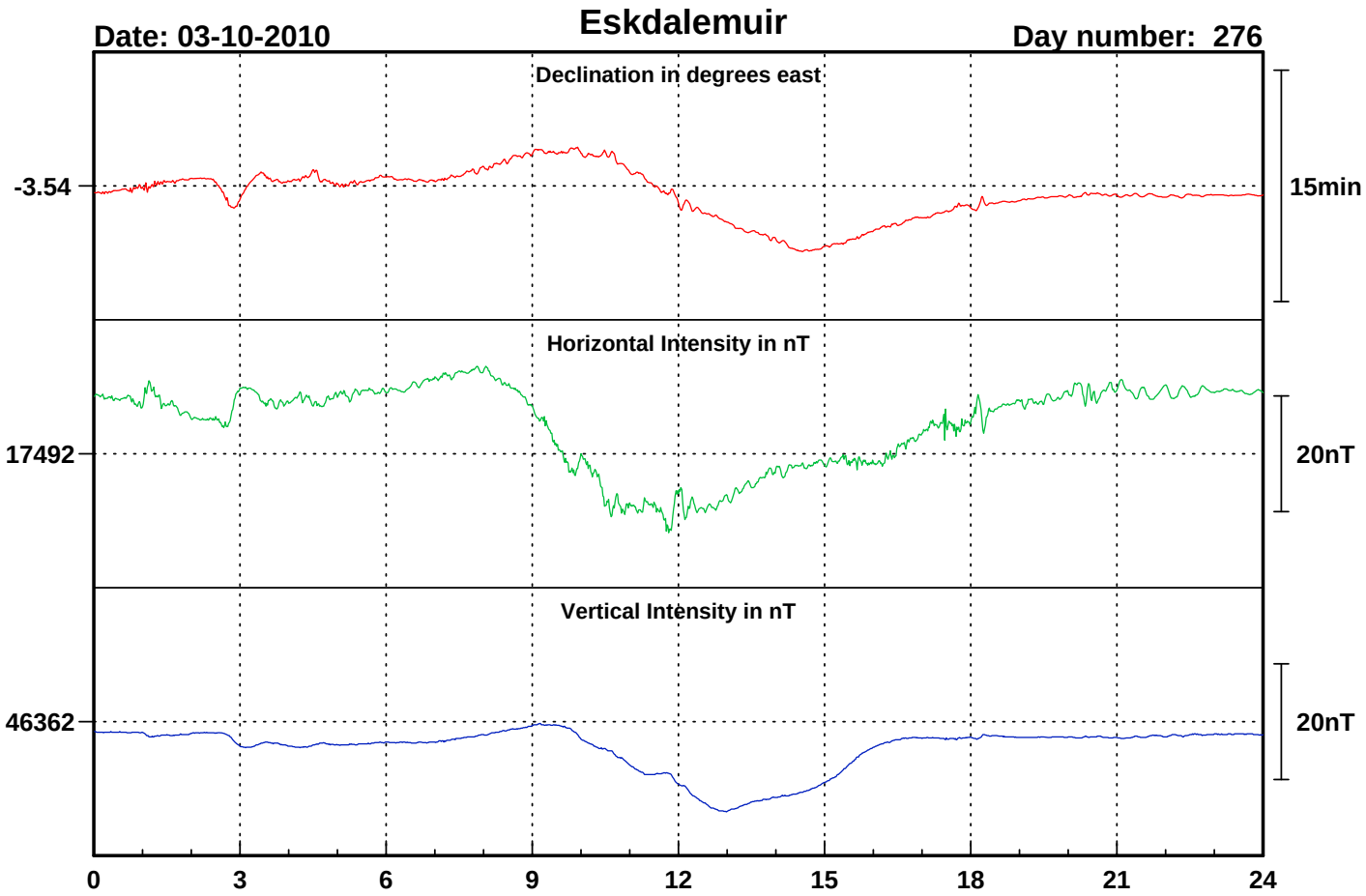
18
21
24
Hour (UT)


Eskdalemuir
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20nT
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0
3
6
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15
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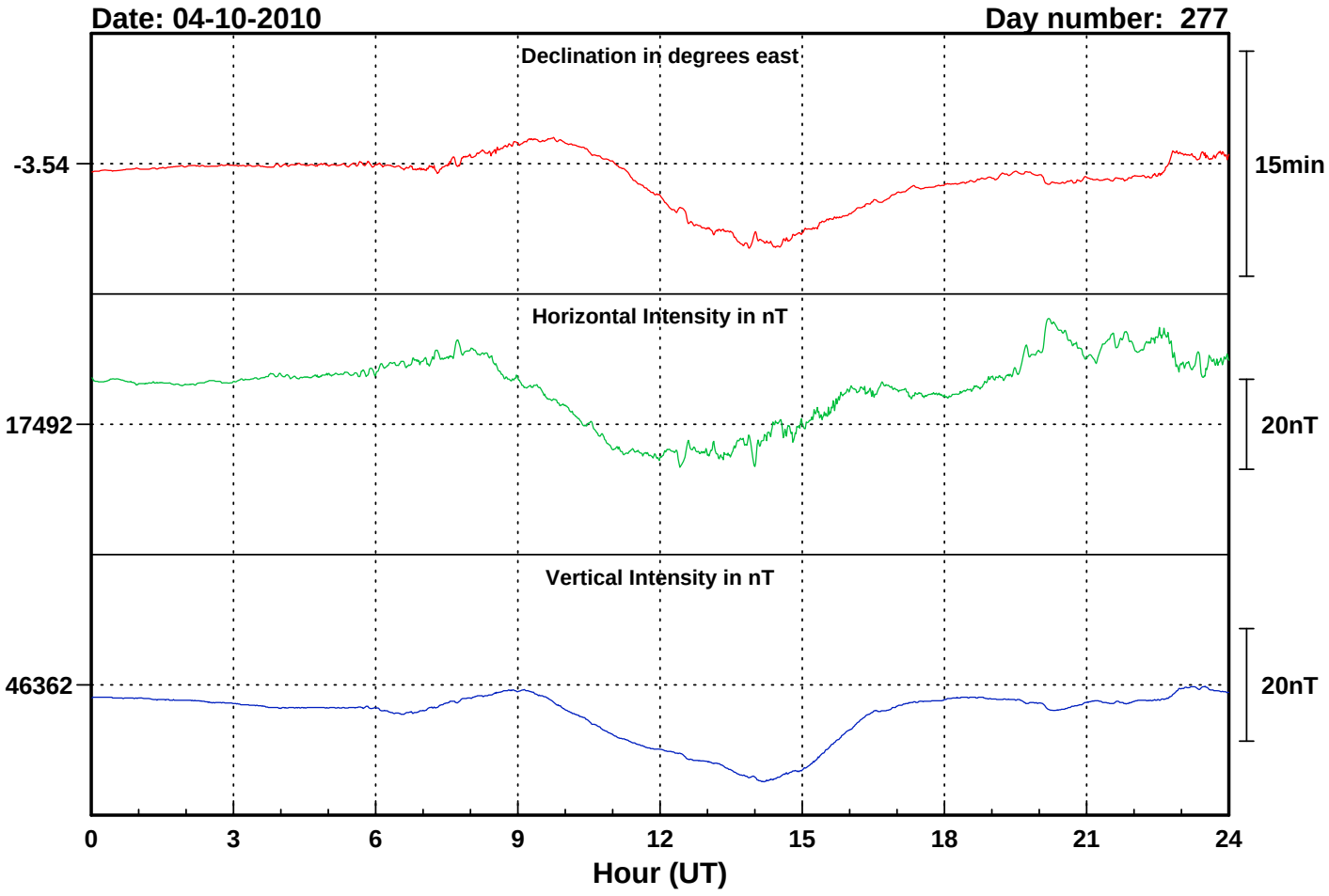
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Vertical Intensity in nT
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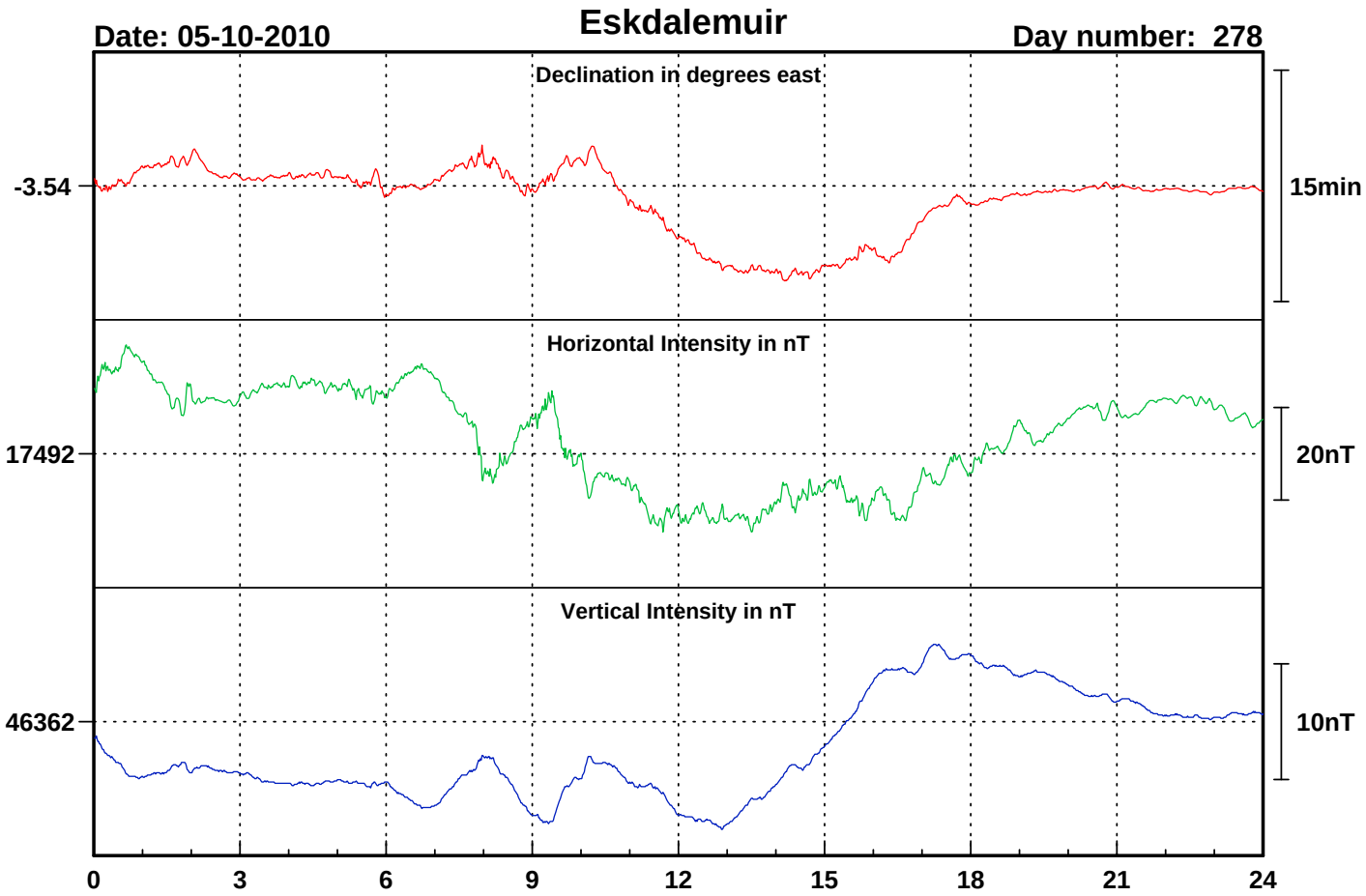
Hour (UT)


Eskdalemuir
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20nT
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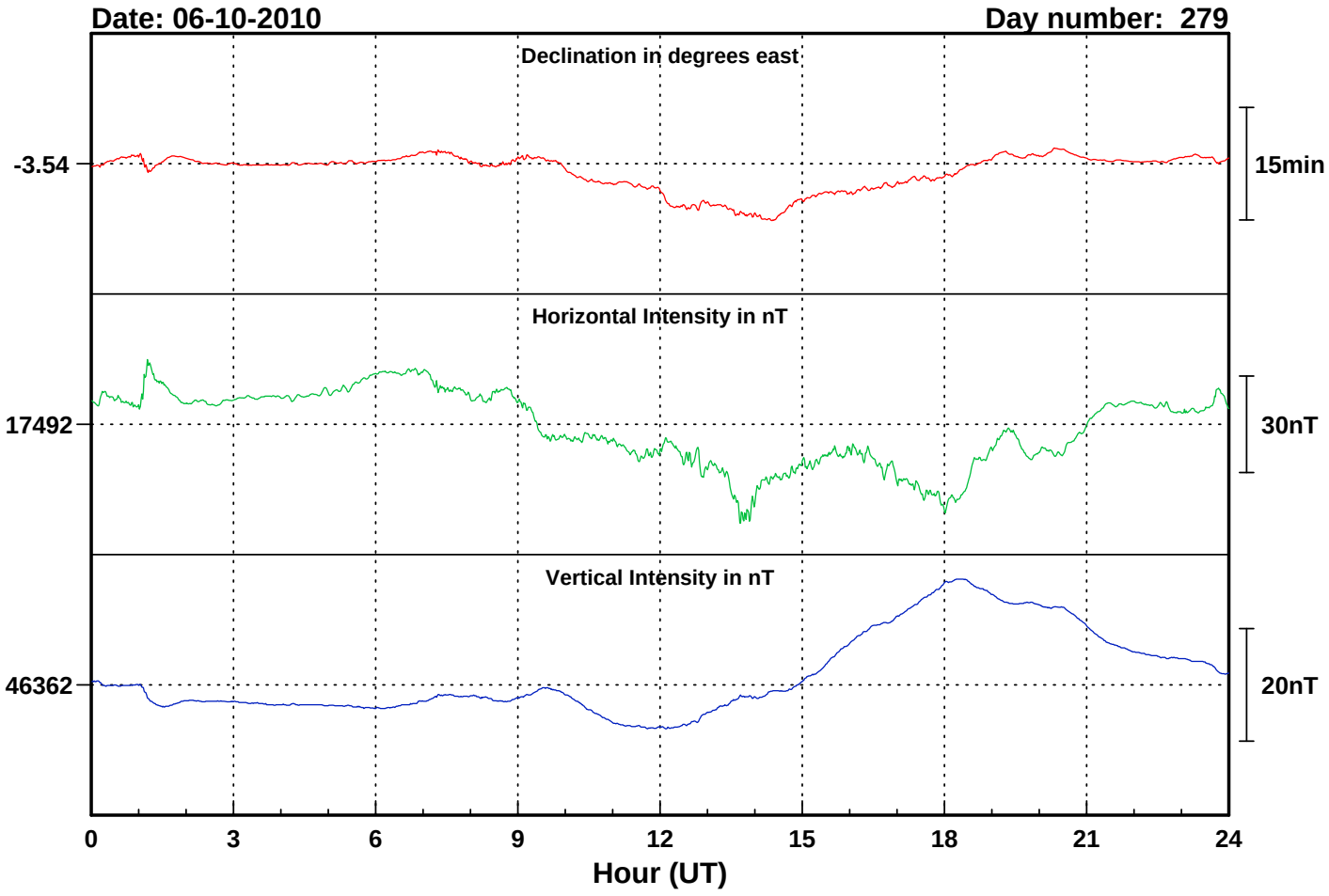
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30nT
Vertical Intensity in nT
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0
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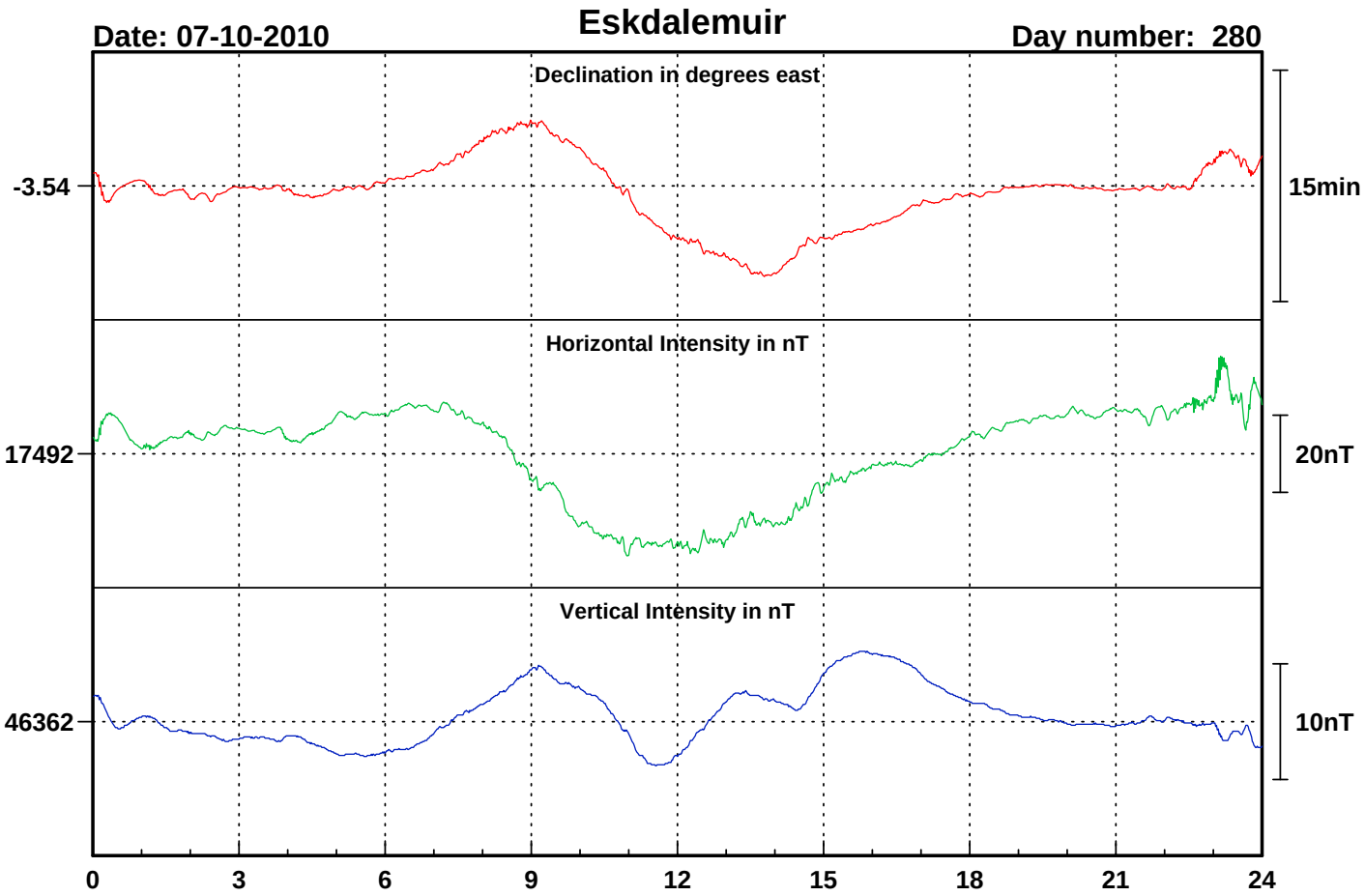
Hour (UT)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20nT
Vertical Intensity in nT
46362
10nT
0
3
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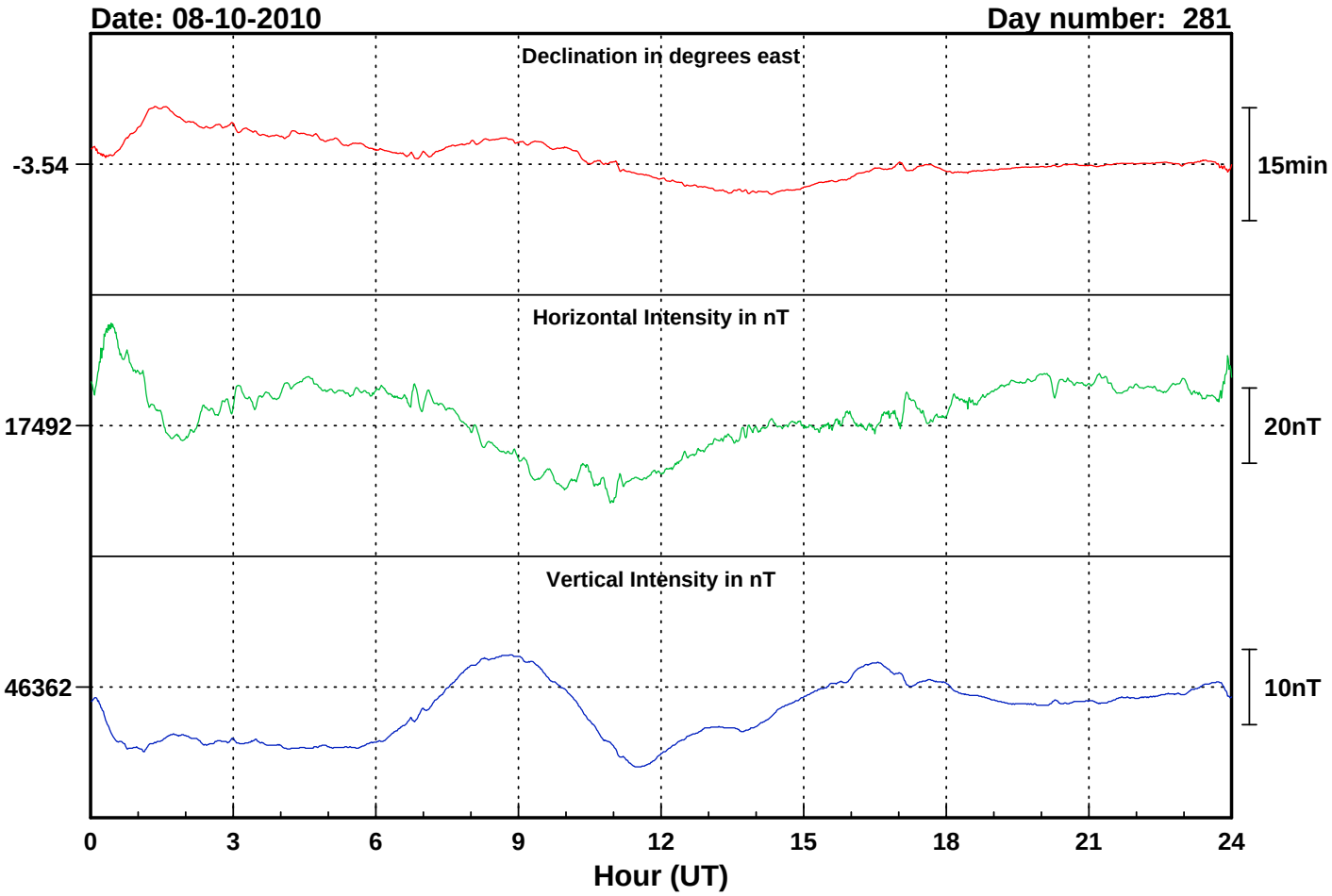
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20nT
Vertical Intensity in nT
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10nT
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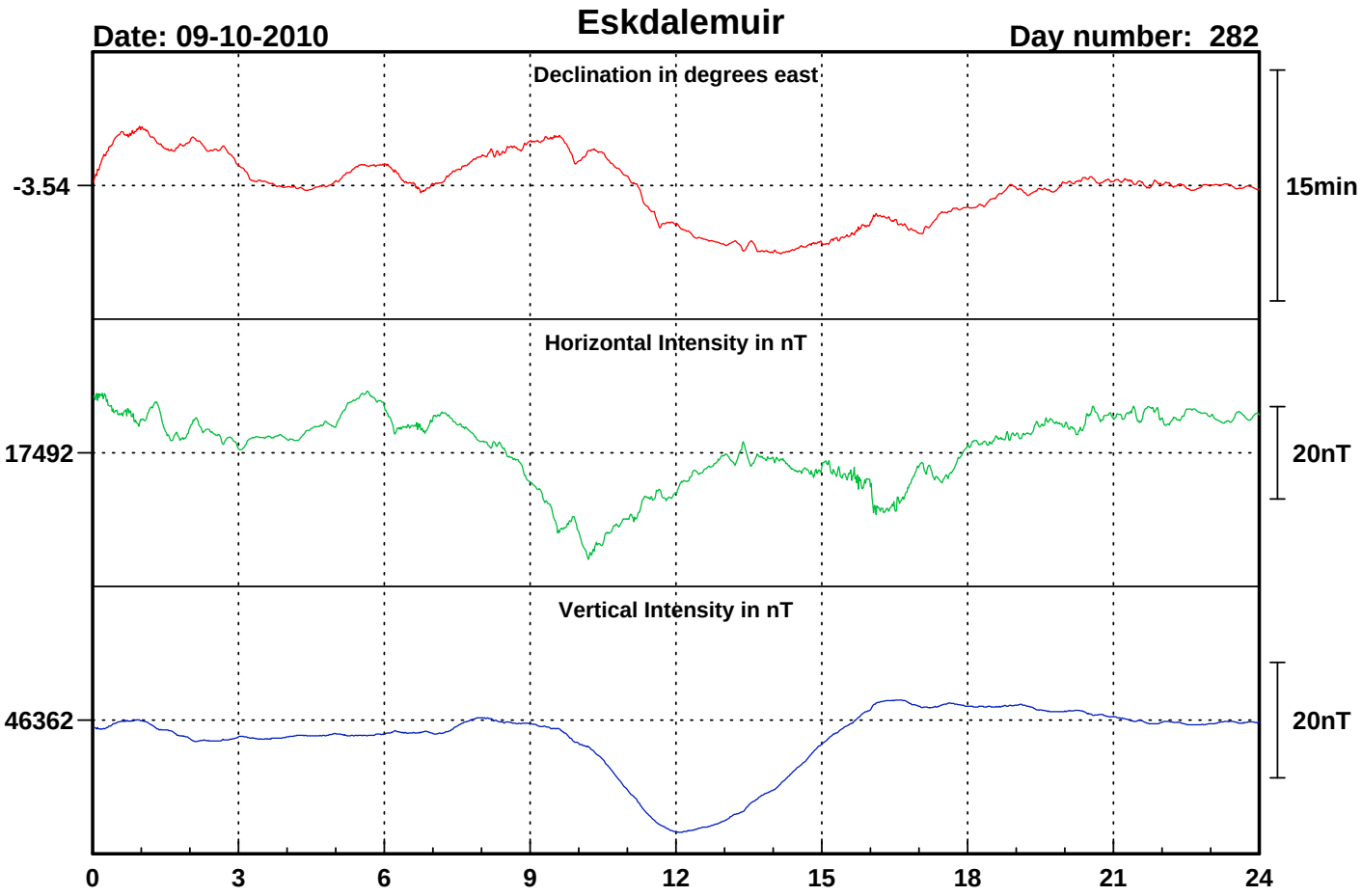
Hour (UT)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20nT
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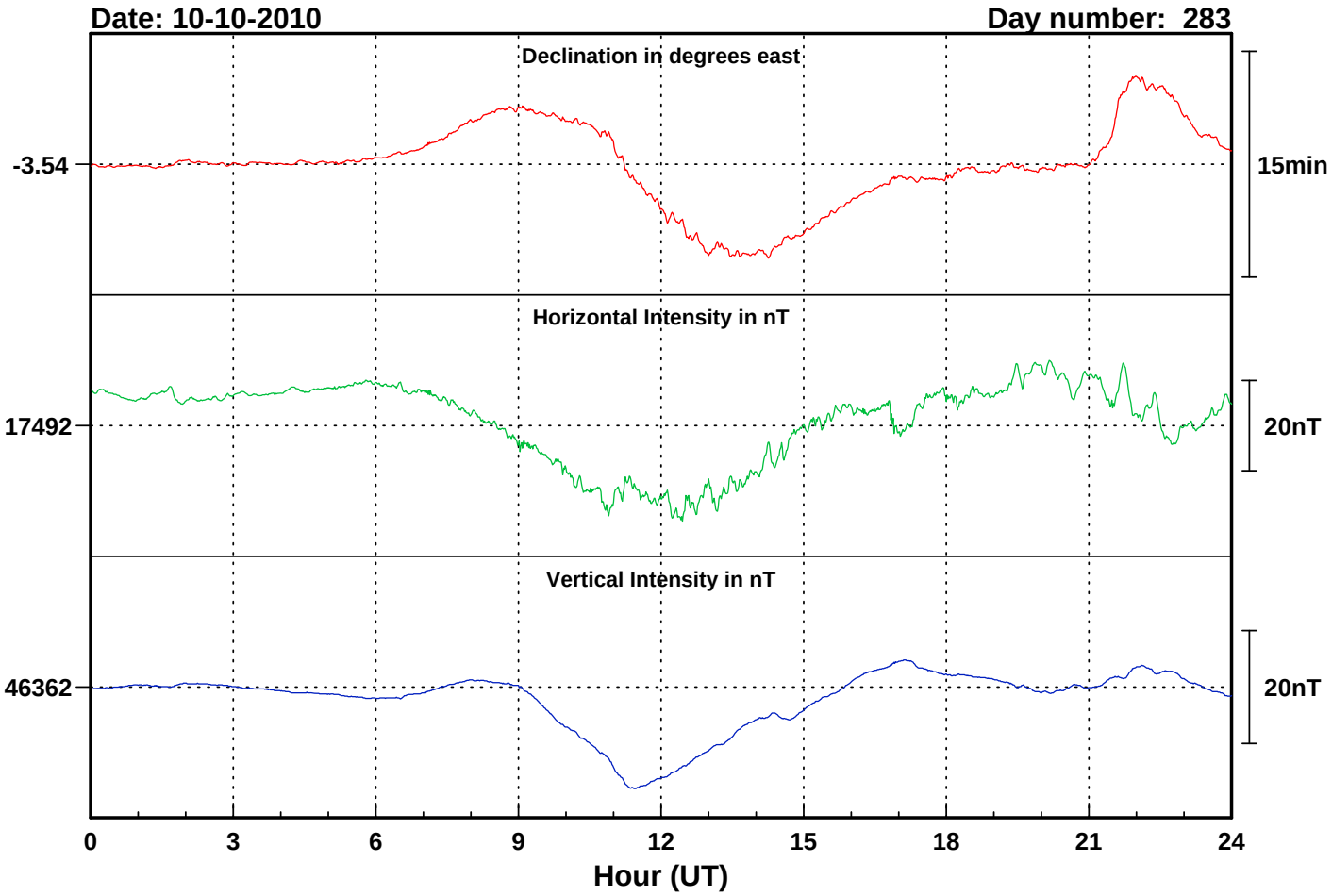
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15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
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20nT
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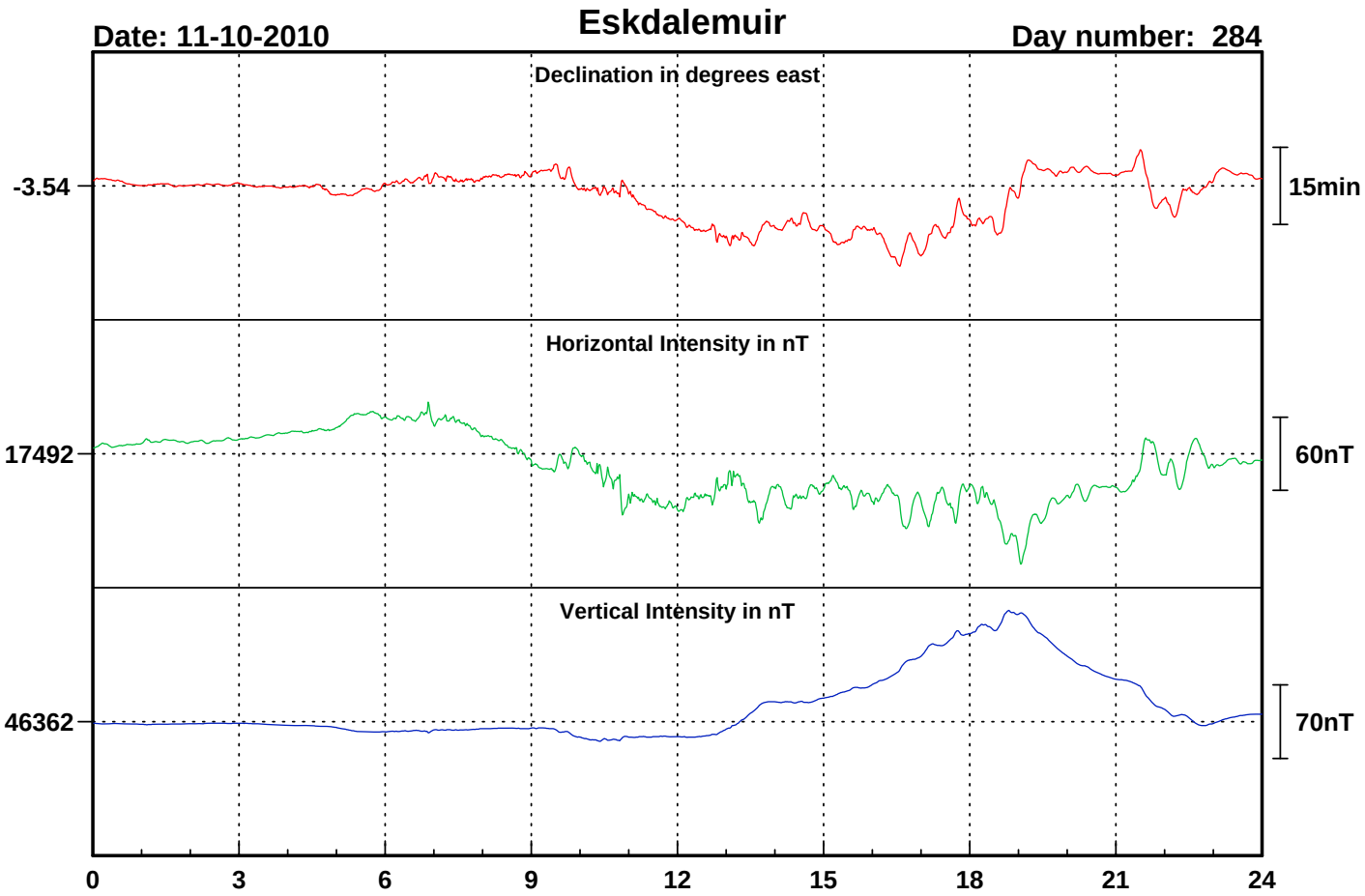
Hour (UT)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15min
Horizontal Intensity in nT
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60nT
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46362
70nT
0
3
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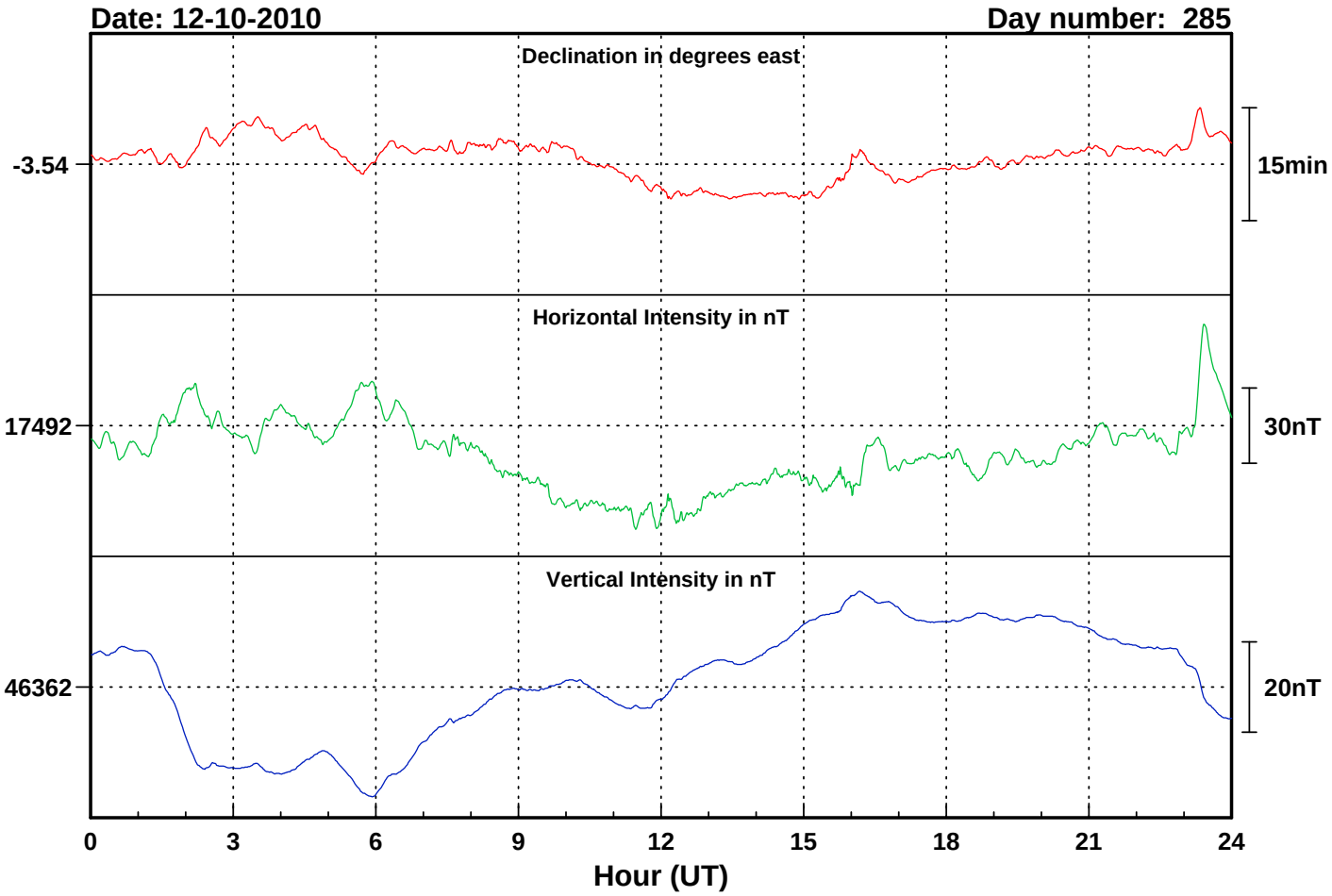
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Declination in degrees east
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15min
Horizontal Intensity in nT
17492
30nT
Vertical Intensity in nT
46362
20nT
0
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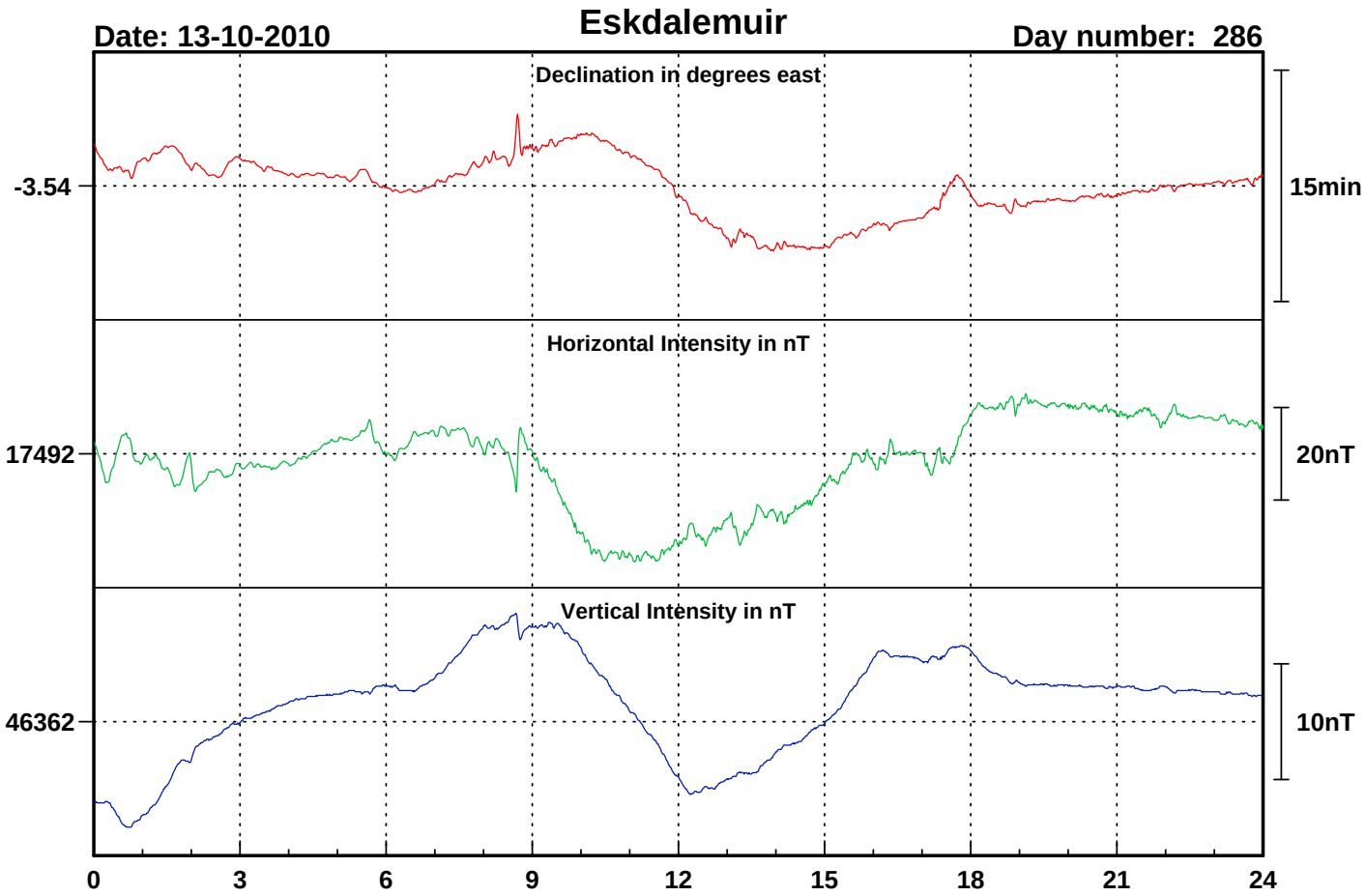
Hour (UT)


Eskdalemuir
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15min
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46362
10nT
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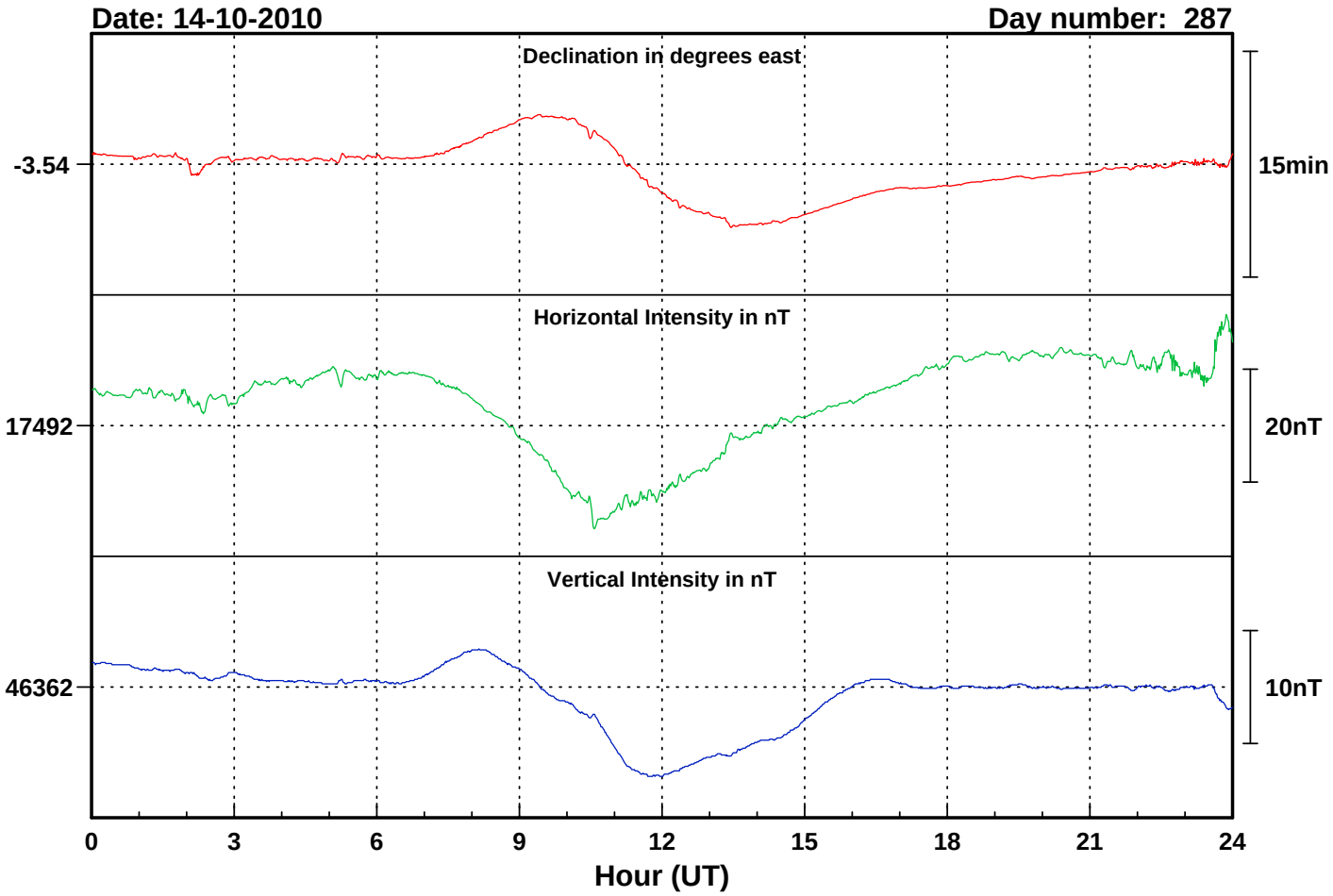
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15min
Horizontal Intensity in nT
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Vertical Intensity in nT
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10nT
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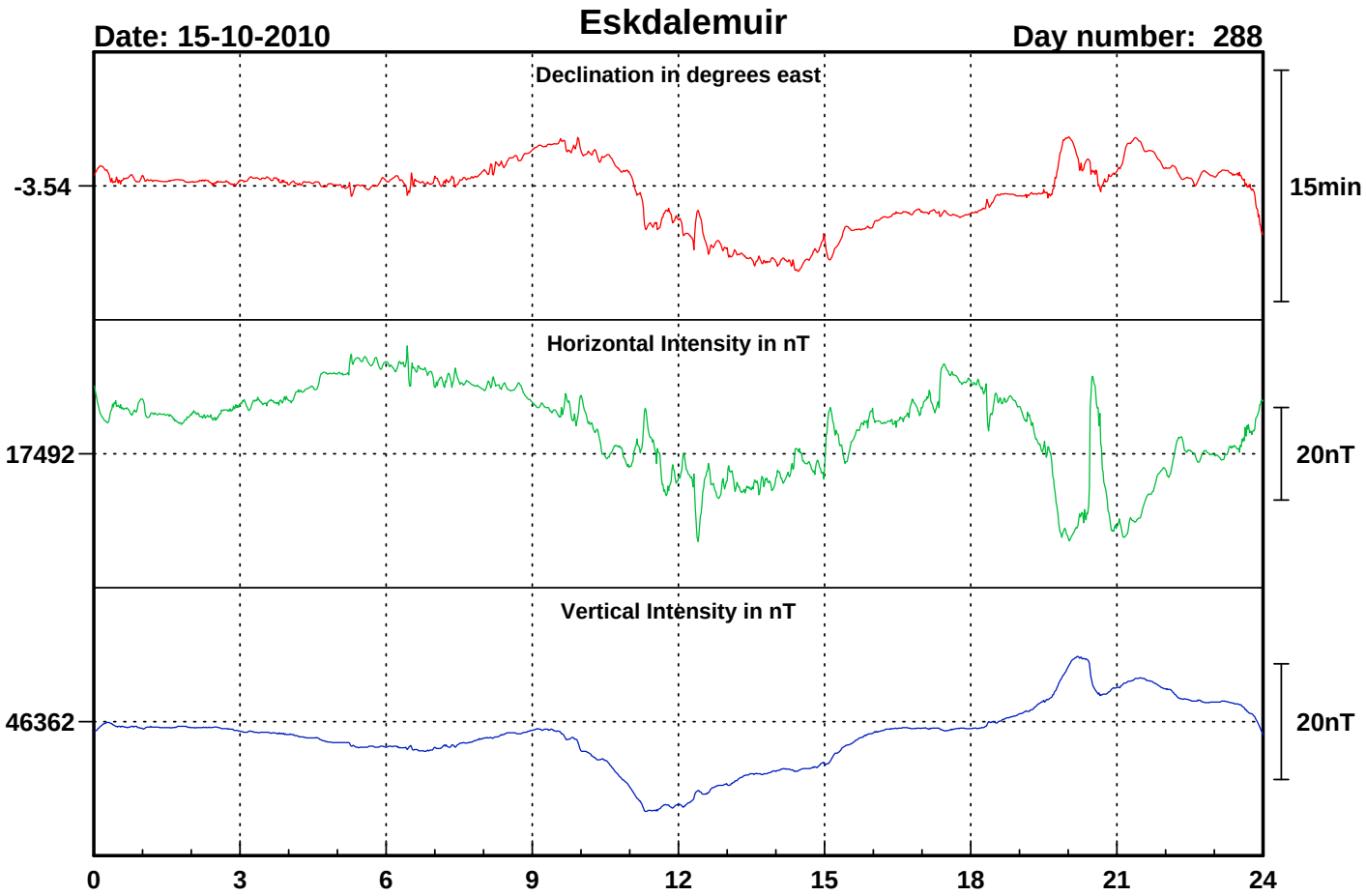
Hour (UT)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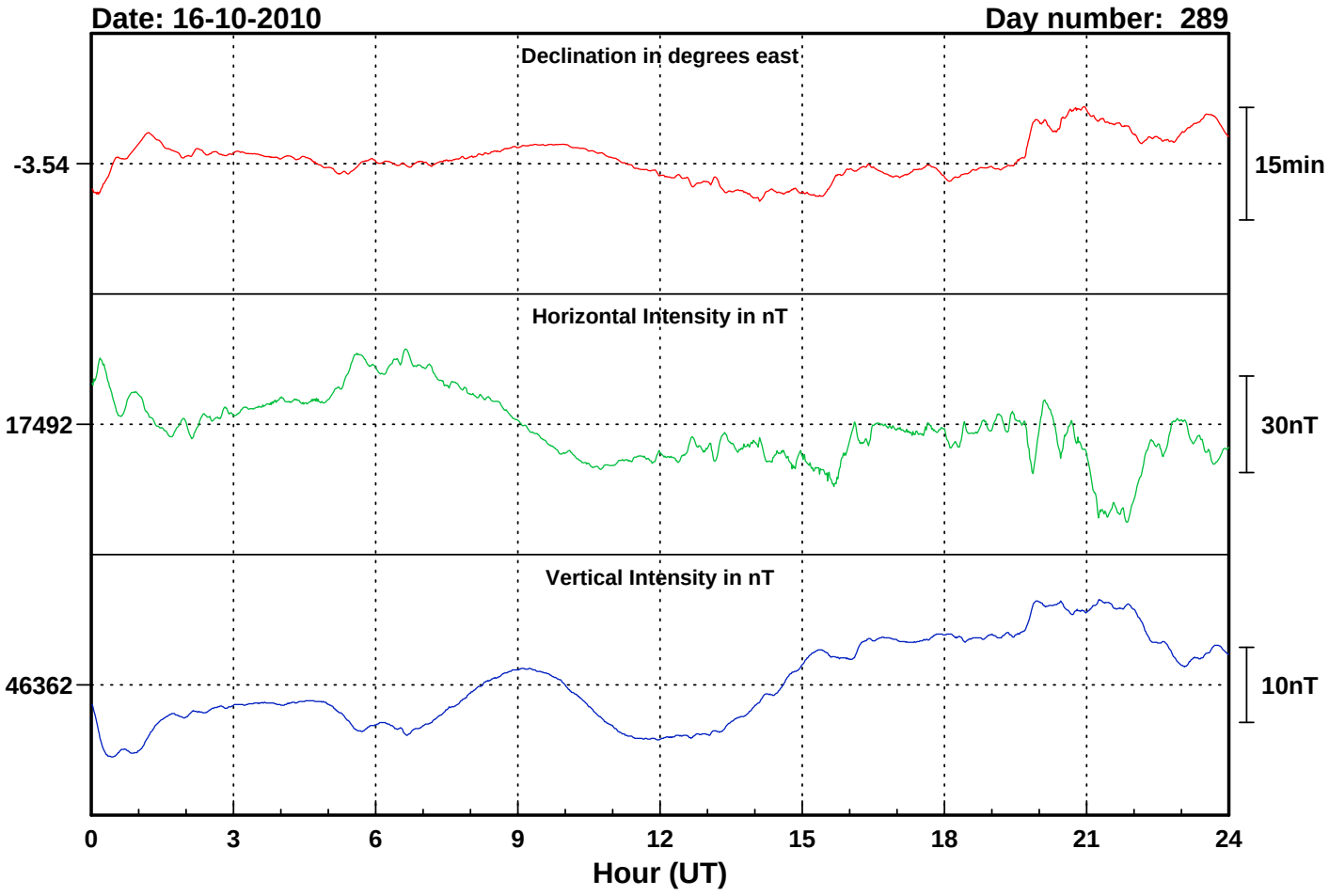
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Vertical Intensity in nT
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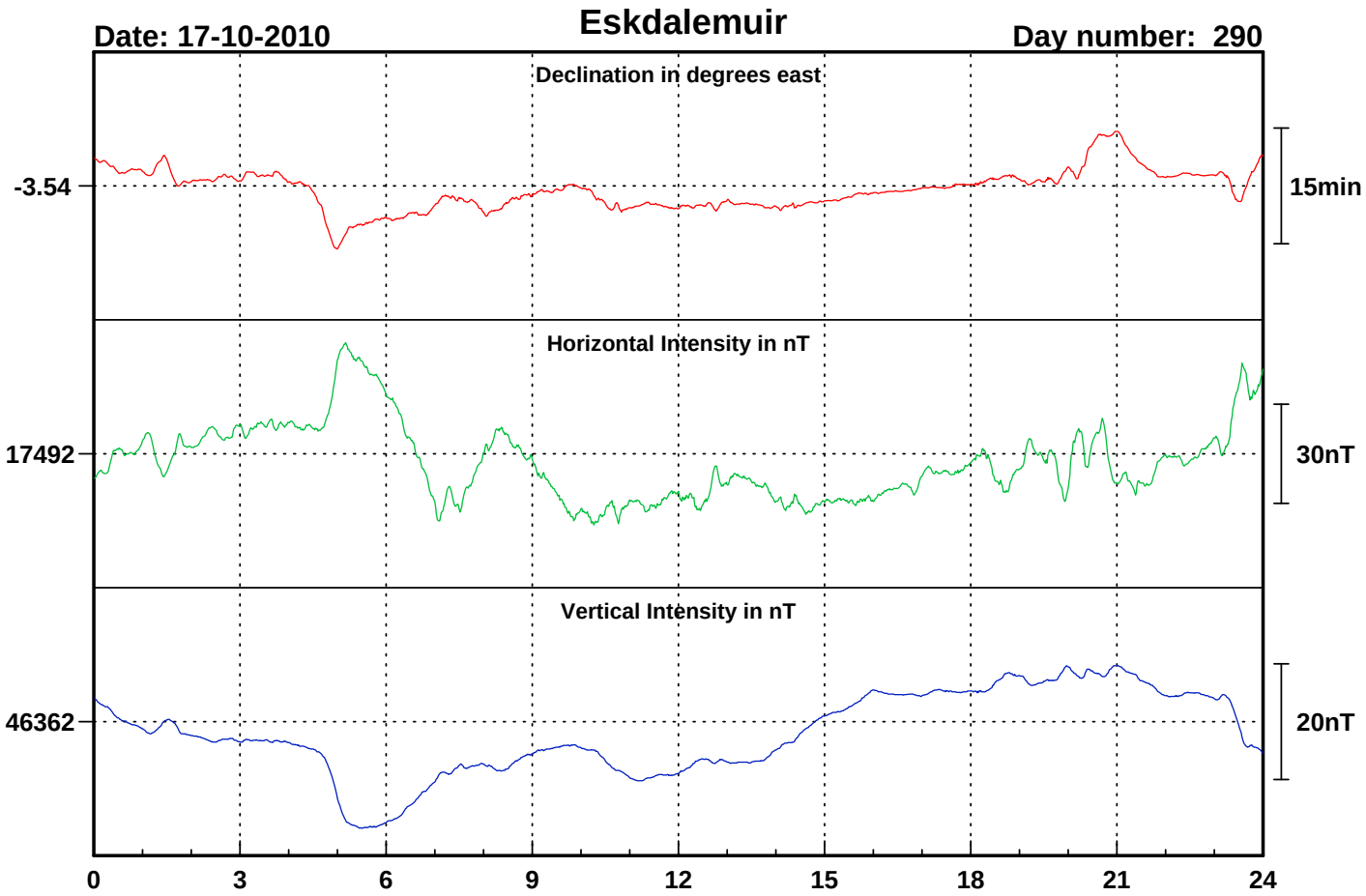
18
21
24
Hour (UT)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15min
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30nT
Vertical Intensity in nT
46362
20nT
0
3
6
9
12
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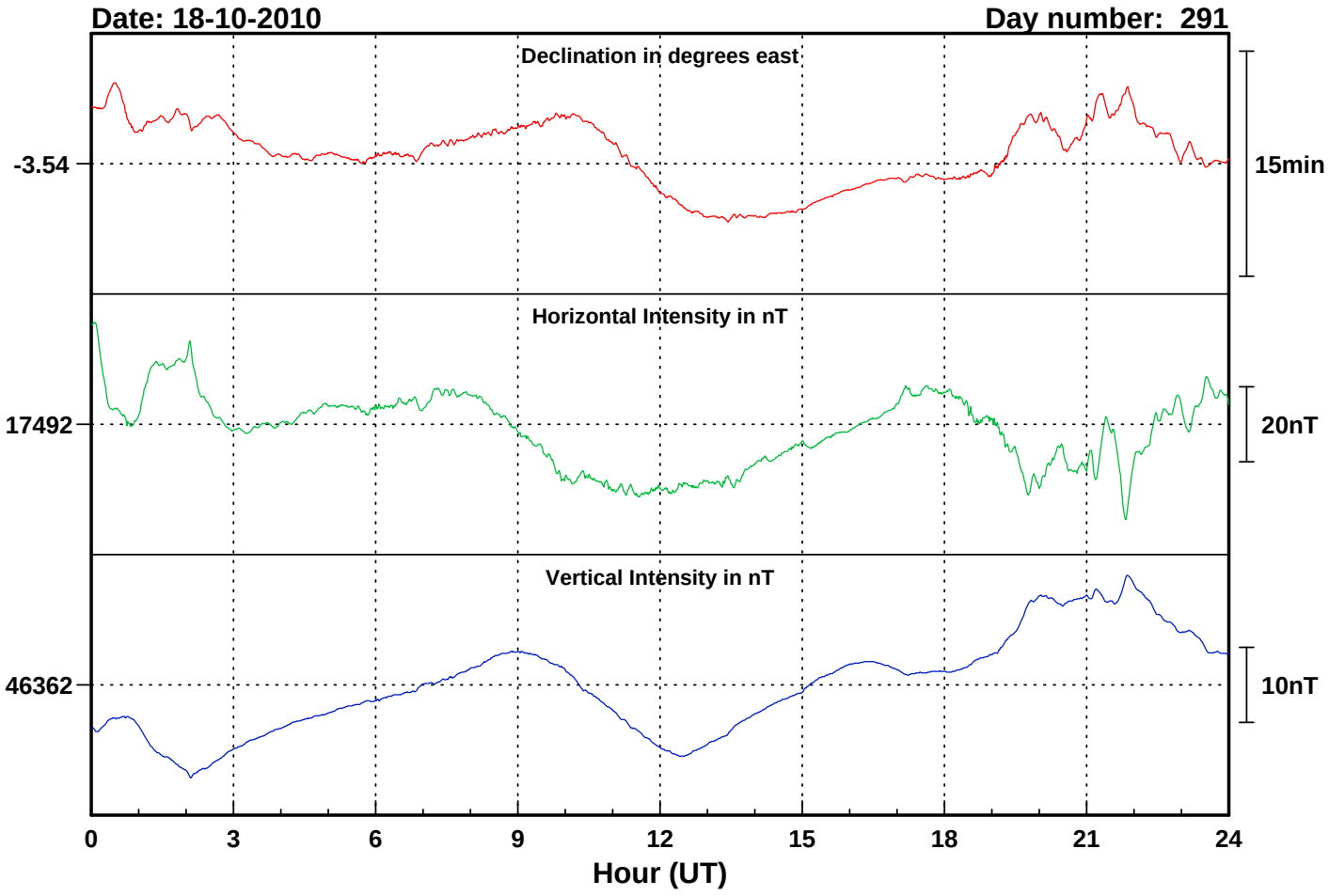
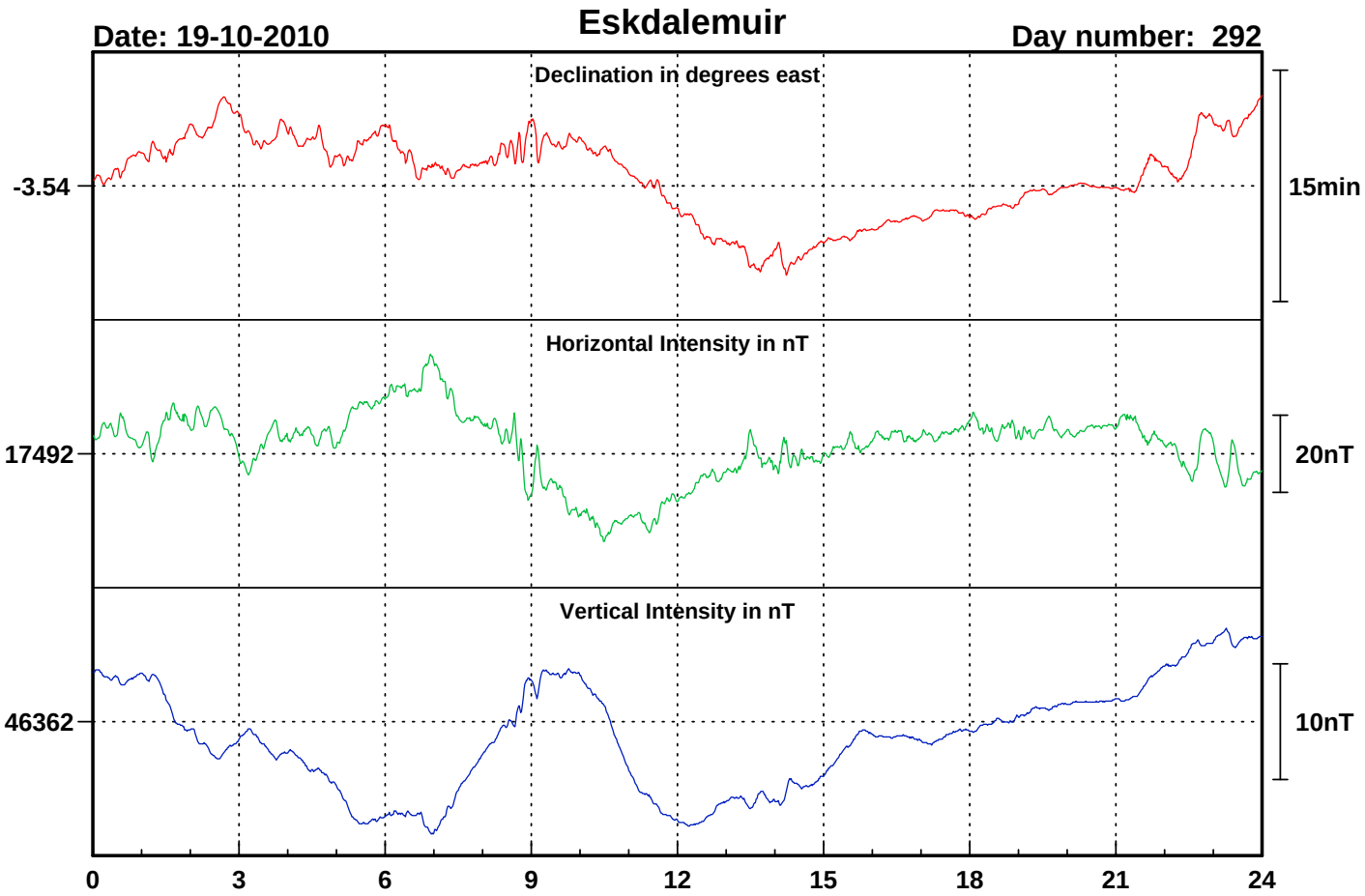
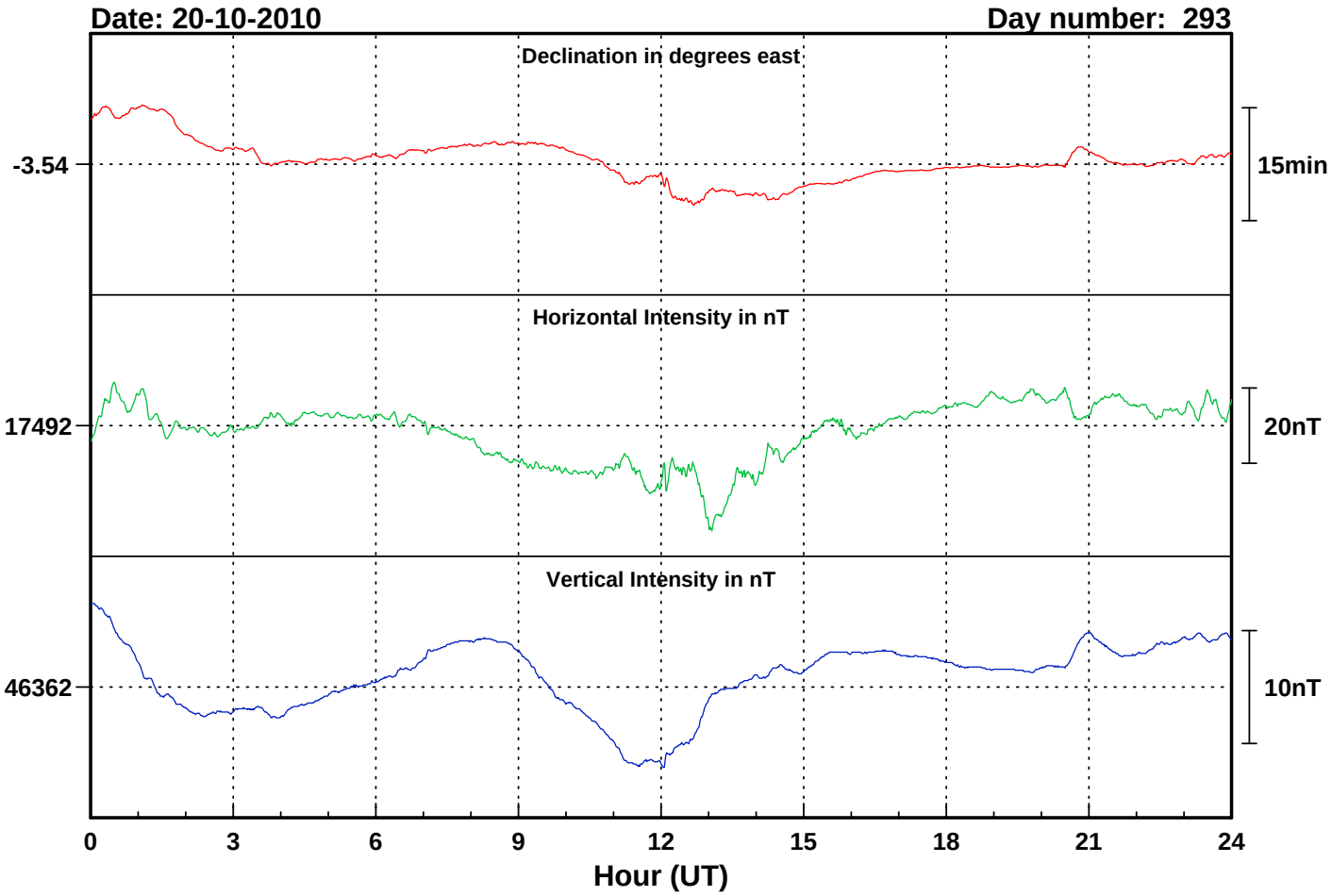
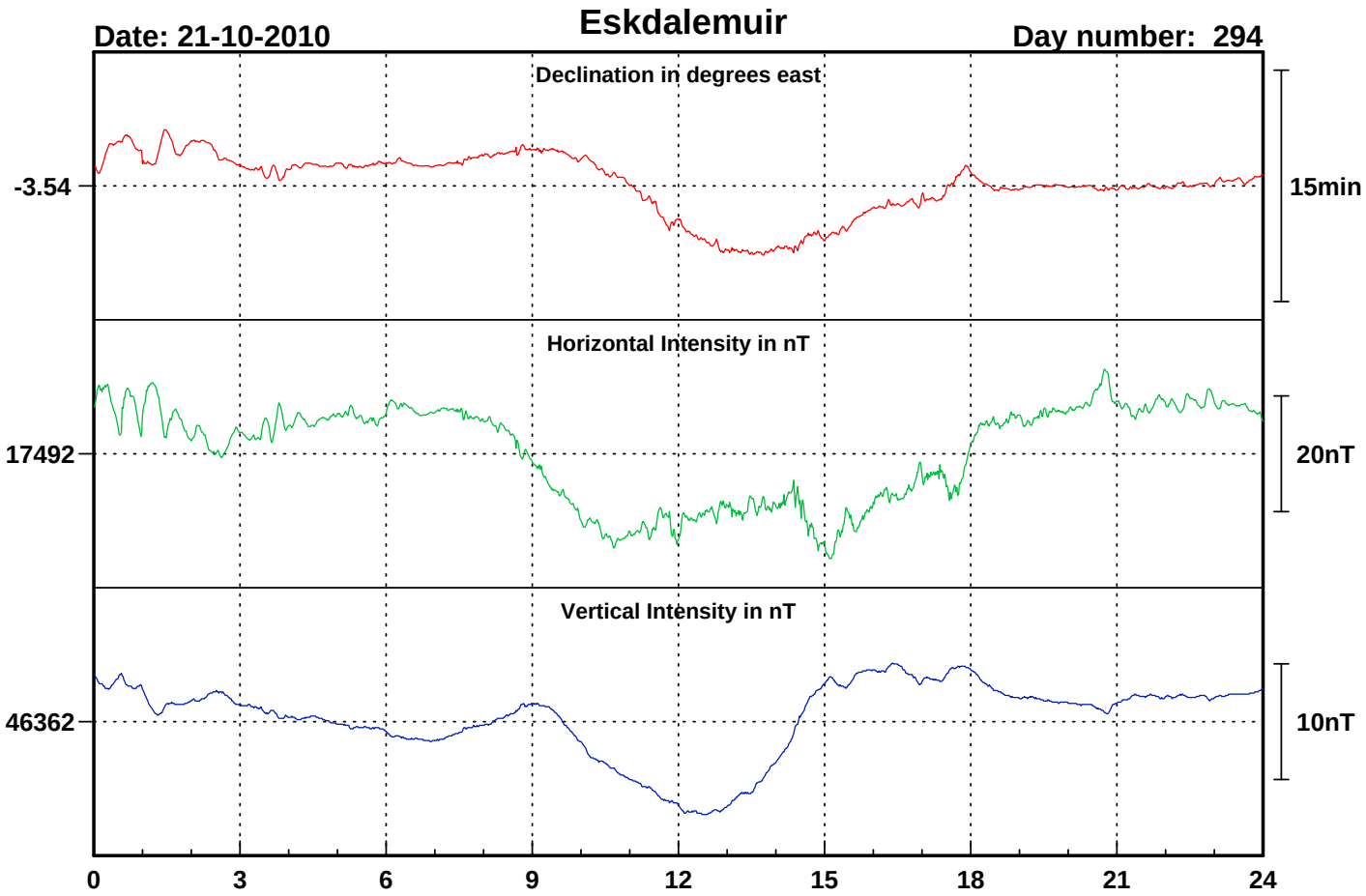
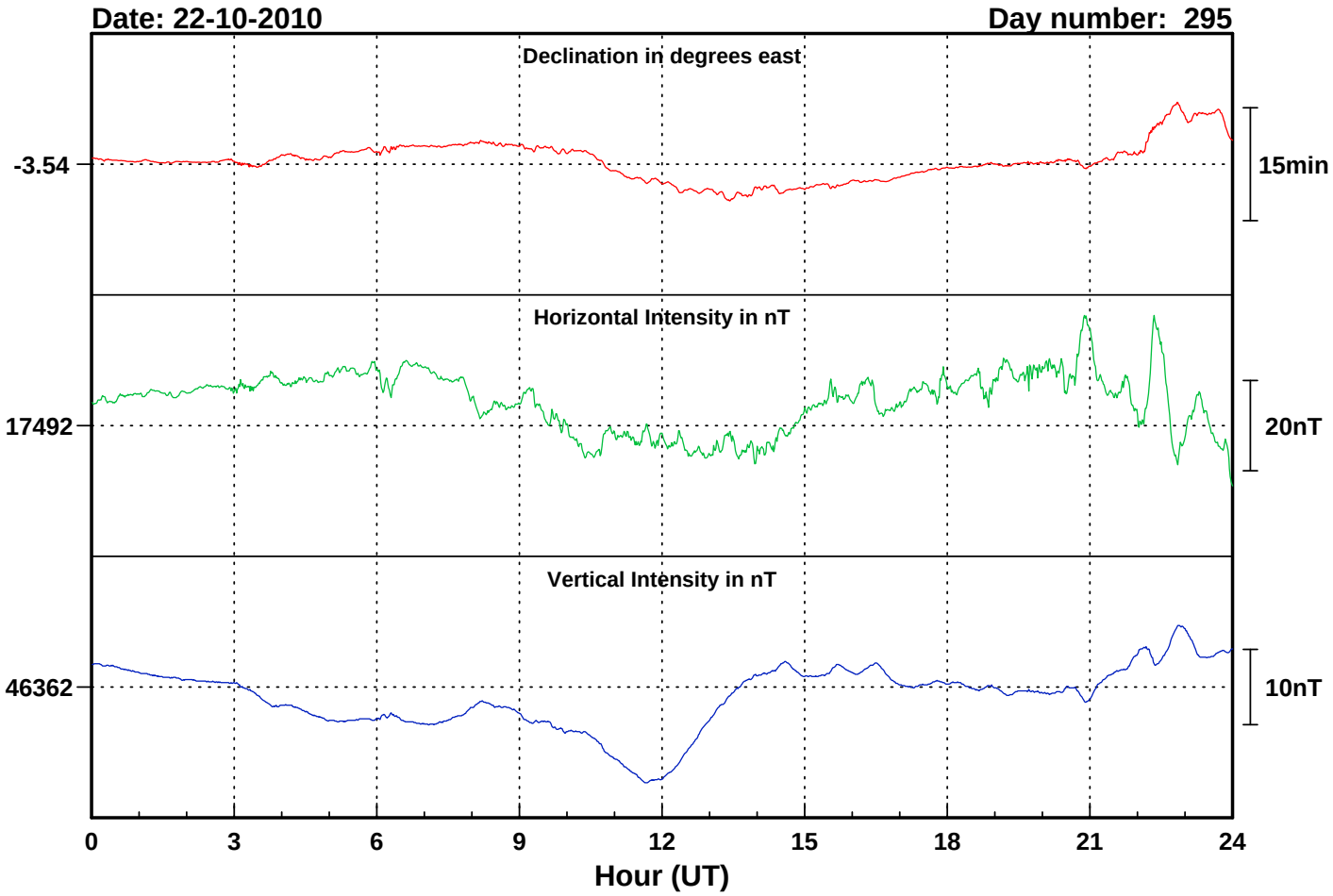
21
24

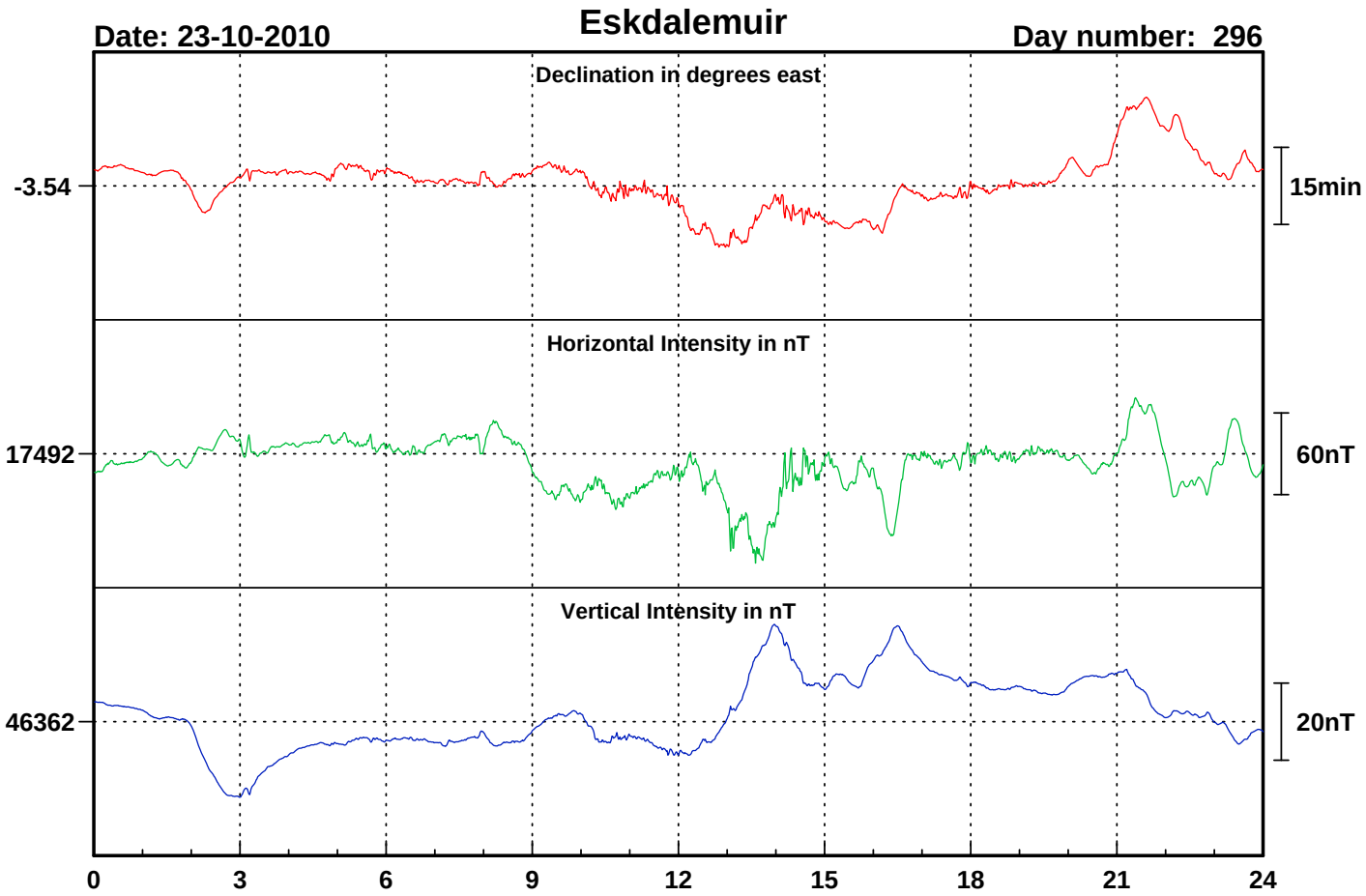

Date: 18-10-2010
Day number: 291
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24
Hour (UT)

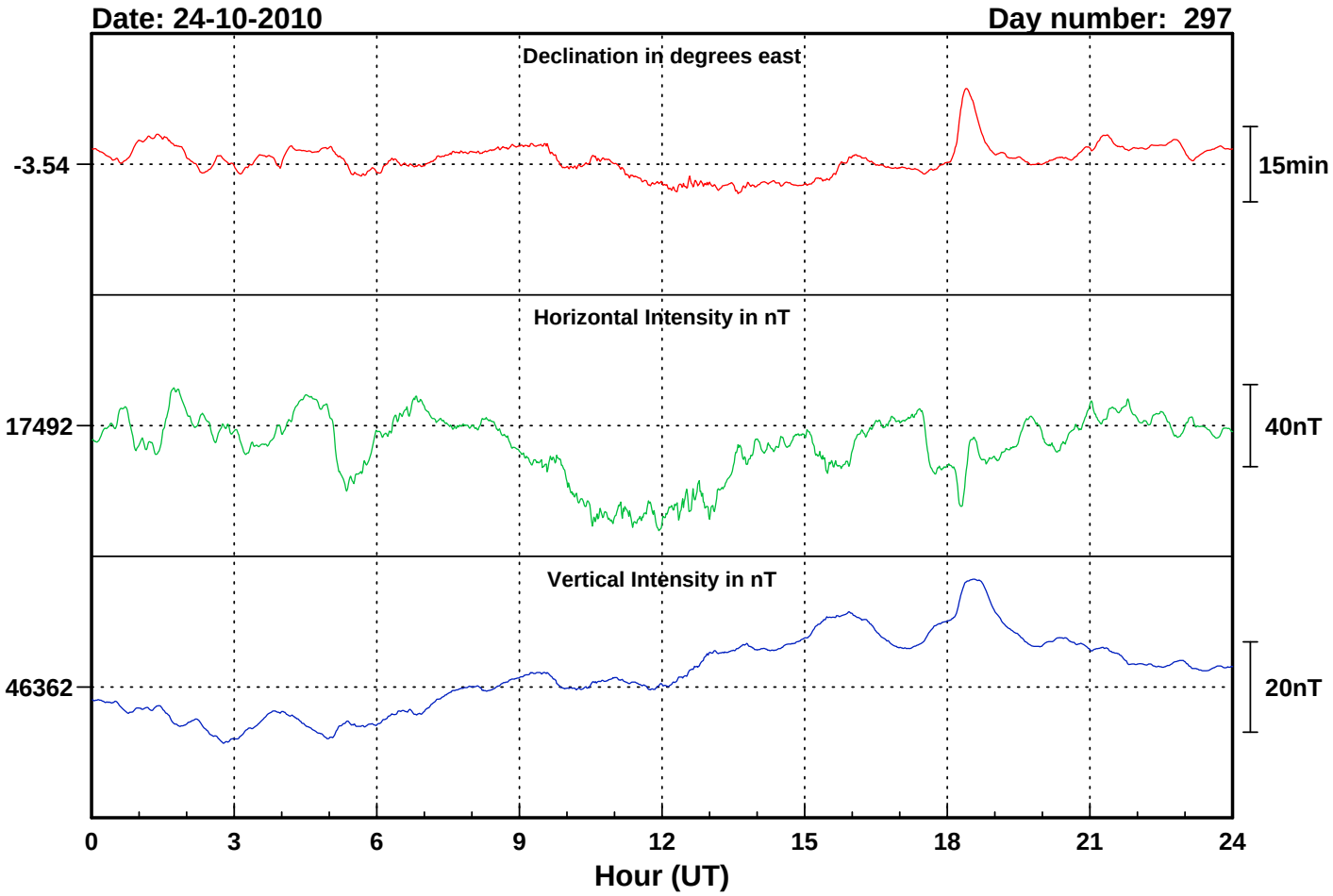

Eskdalemuir
Date: 19-10-2010
Day number: 292
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24

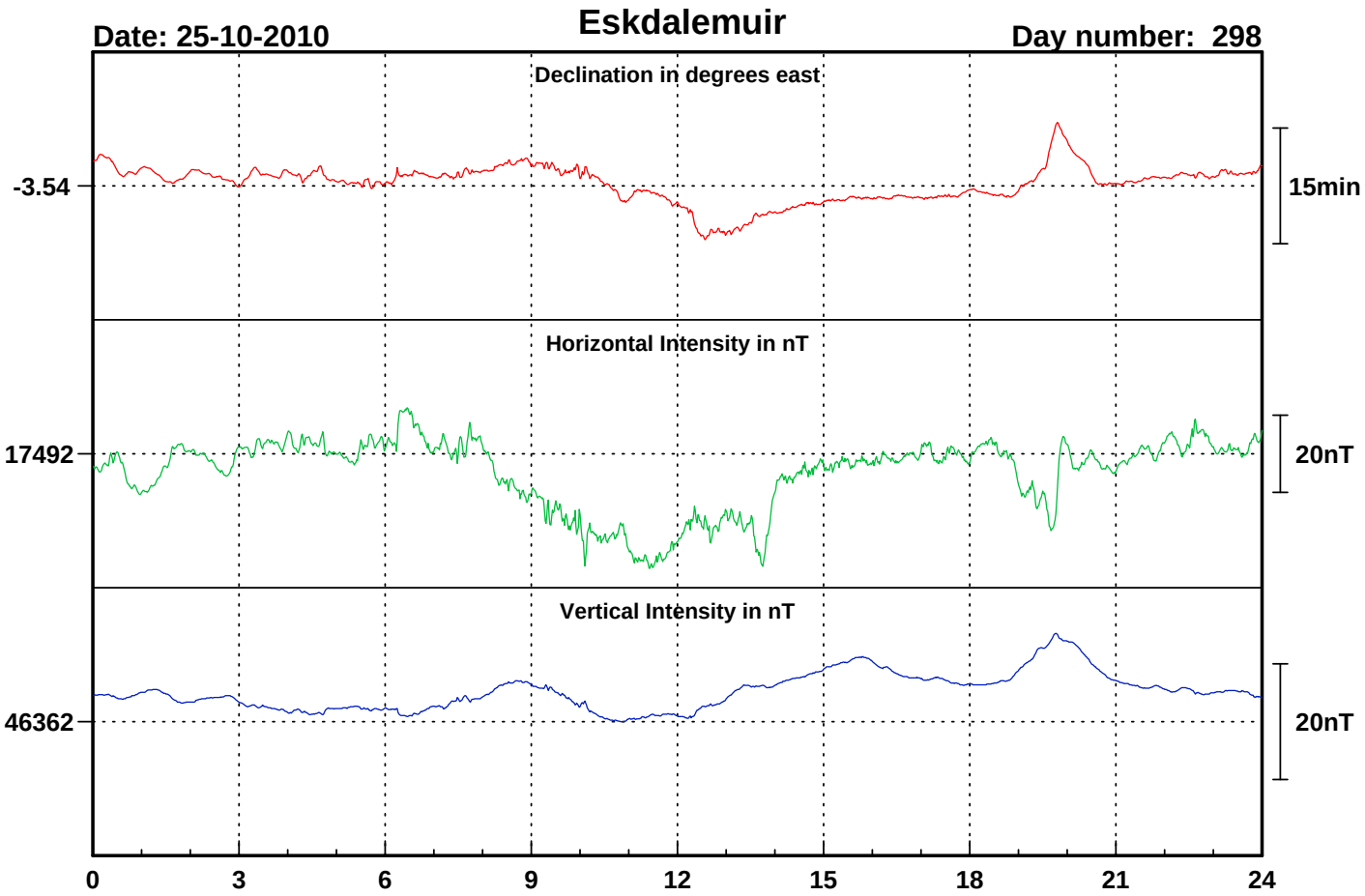

Date: 20-10-2010
Day number: 293
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24
Hour (UT)

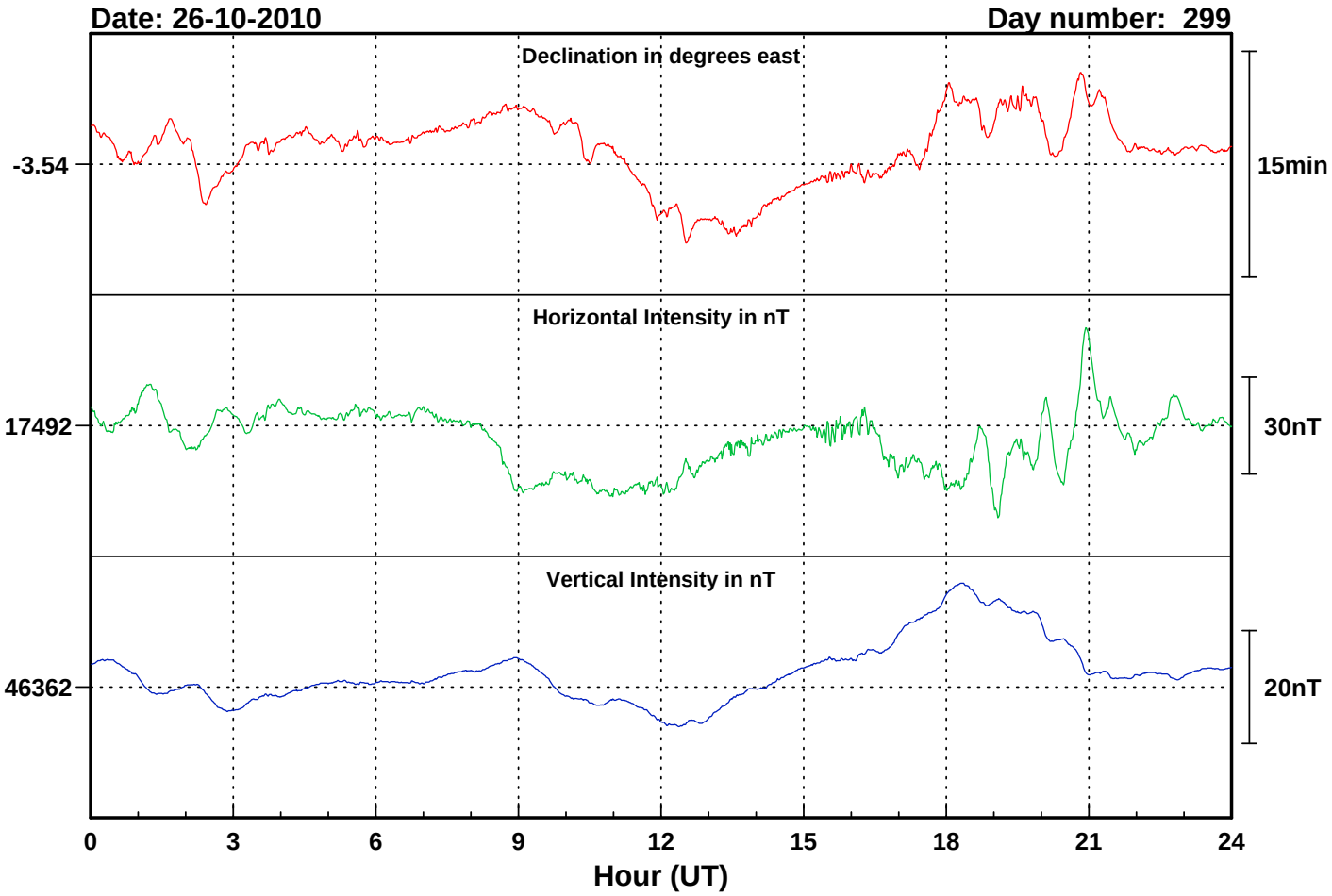

Eskdalemuir
Date: 21-10-2010
Day number: 294
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24

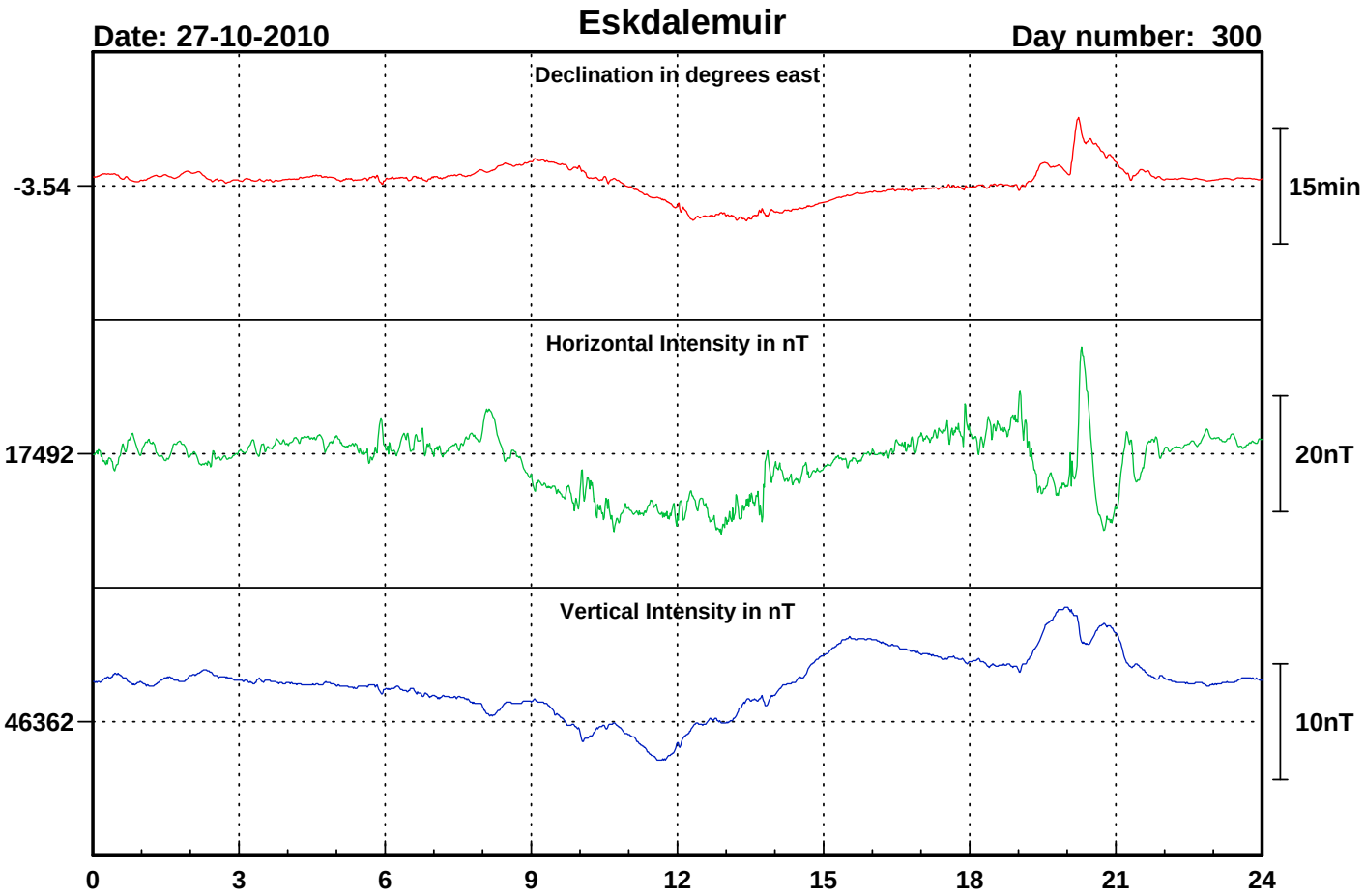

Date: 22-10-2010
Day number: 295
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24
Hour (UT)

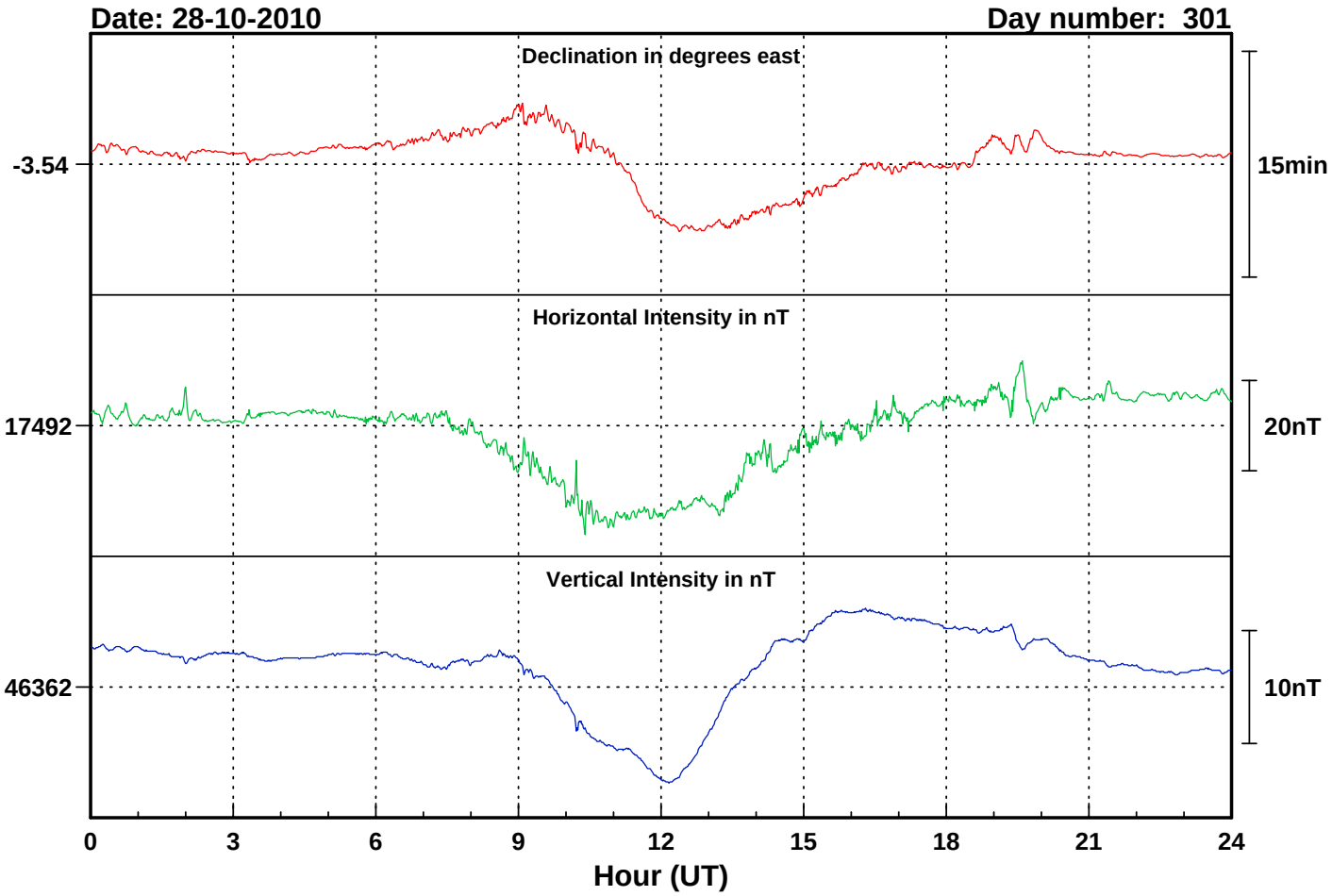

Eskdalemuir
Date: 23-10-2010
Day number: 296
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
60nT
Vertical Intensity in nT
46362
20nT
0
3
6
9
12
15
18
21
24

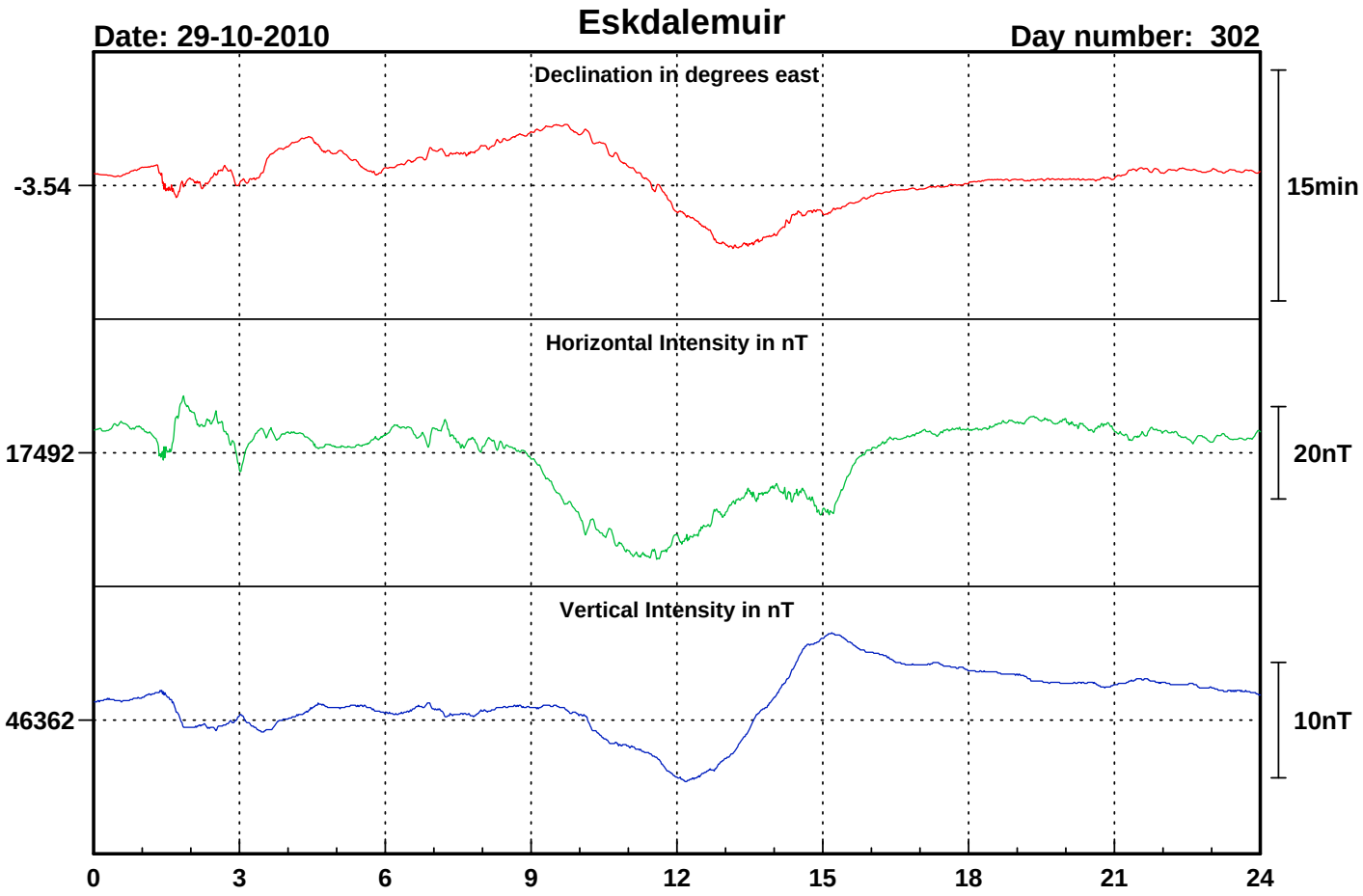

Date: 24-10-2010
Day number: 297
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
40nT
Vertical Intensity in nT
46362
20nT
0
3
6
9
12
15
18
21
24
Hour (UT)

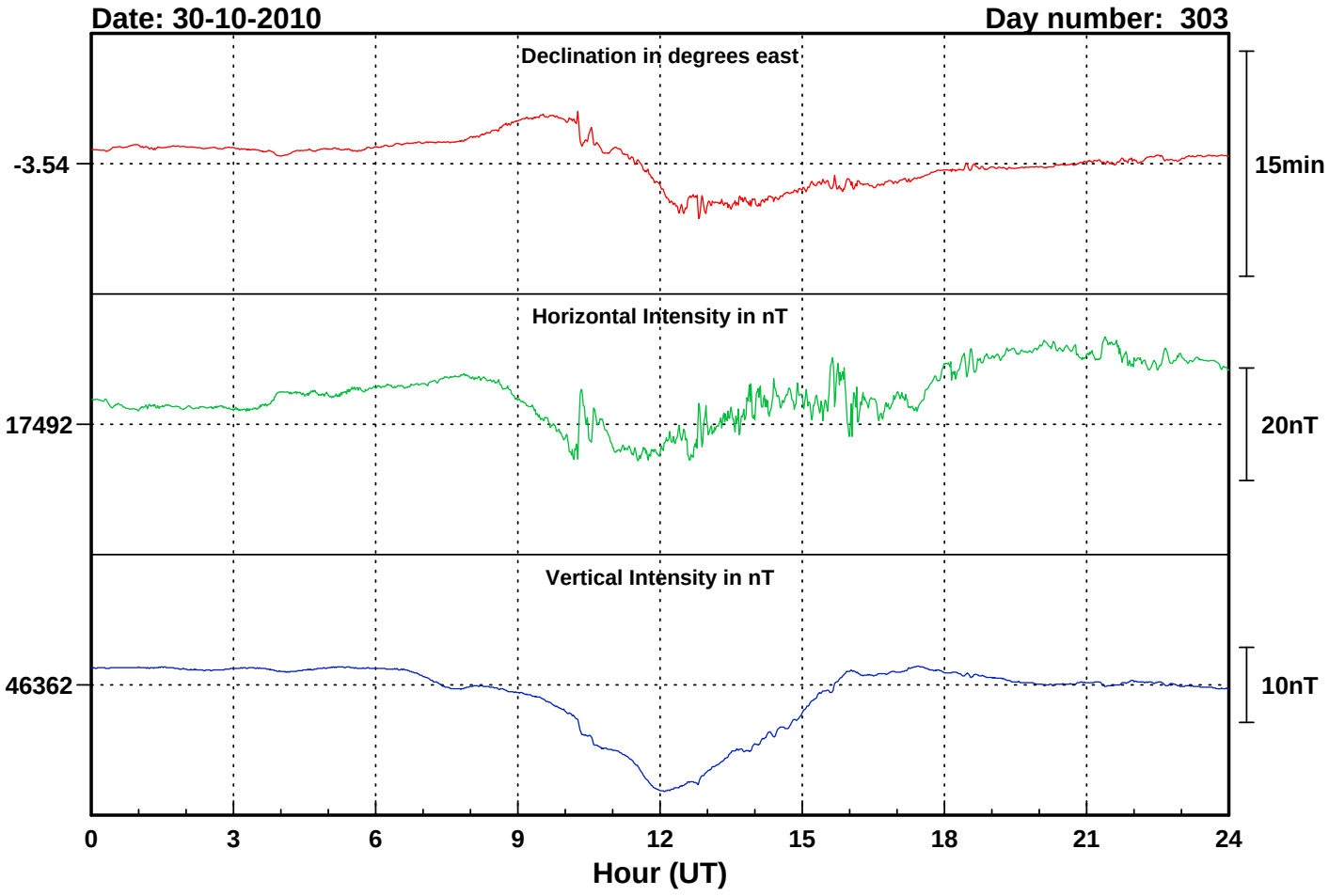

Eskdalemuir
Date: 25-10-2010
Day number: 298
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
20nT
0
3
6
9
12
15
18
21
24

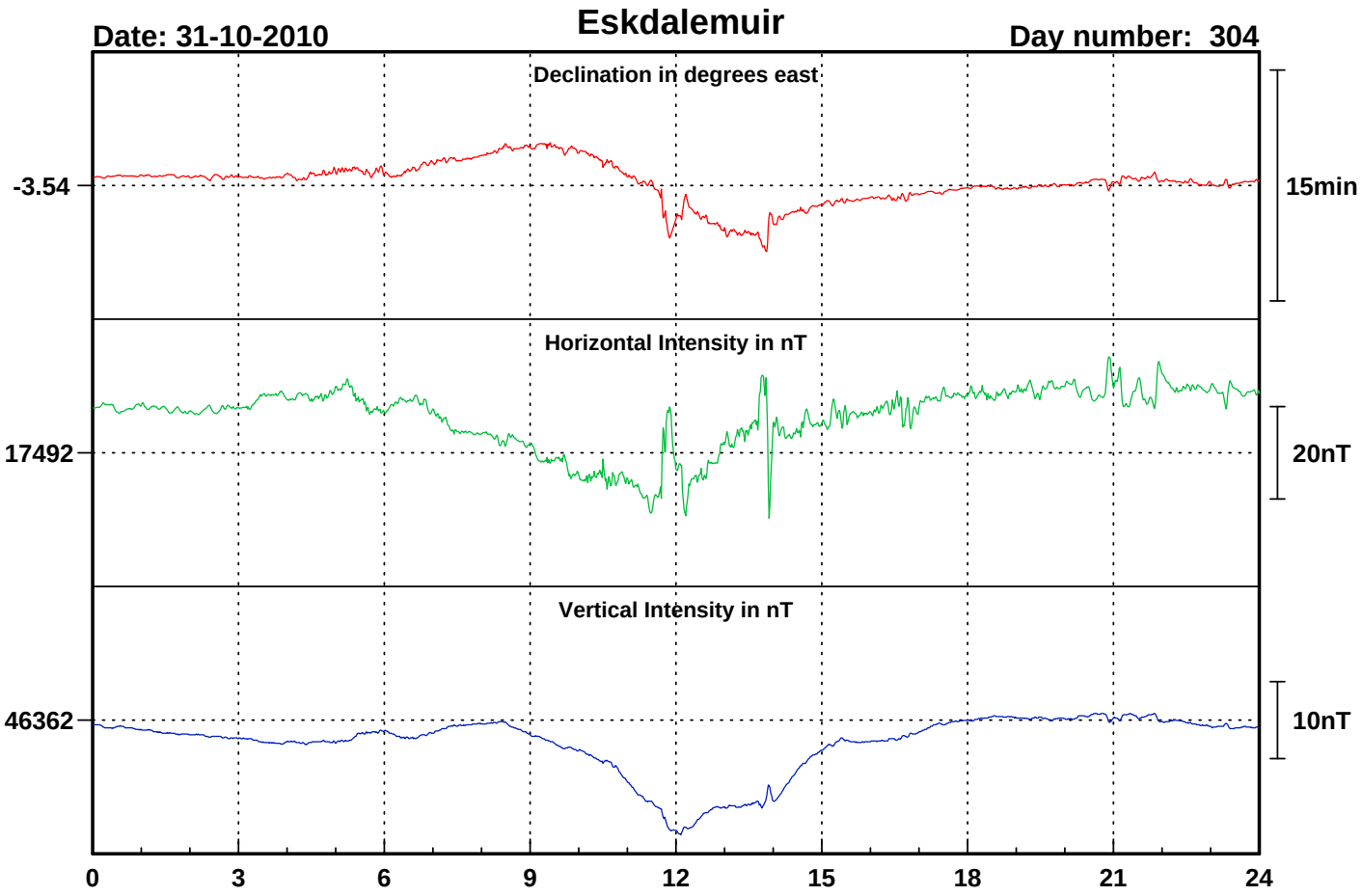

Date: 26-10-2010
Day number: 299
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
30nT
Vertical Intensity in nT
46362
20nT
0
3
6
9
12
15
18
21
24
Hour (UT)


Eskdalemuir
Date: 27-10-2010
Day number: 300
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24

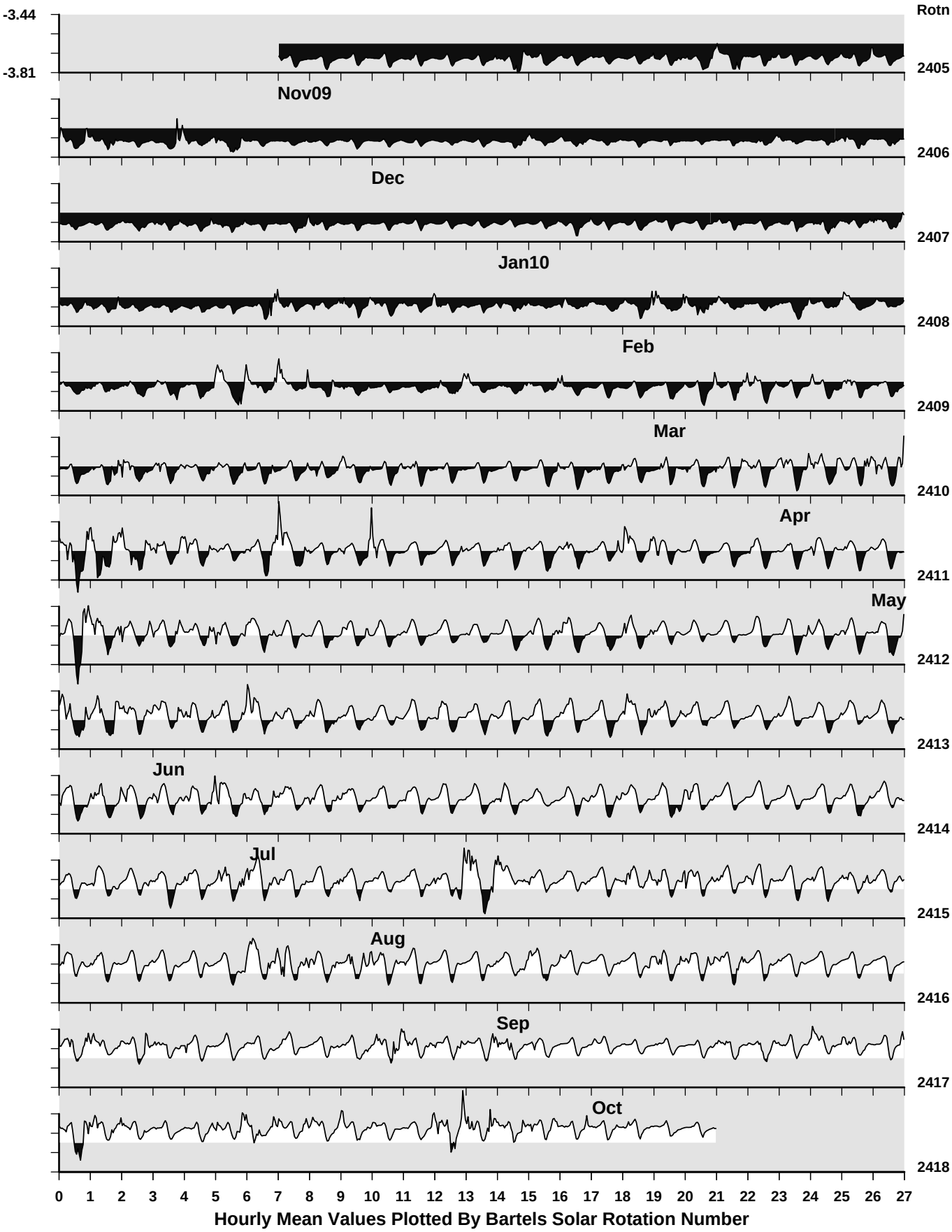

Date: 28-10-2010
Day number: 301
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24
Hour (UT)


Eskdalemuir
Date: 29-10-2010
Day number: 302
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24

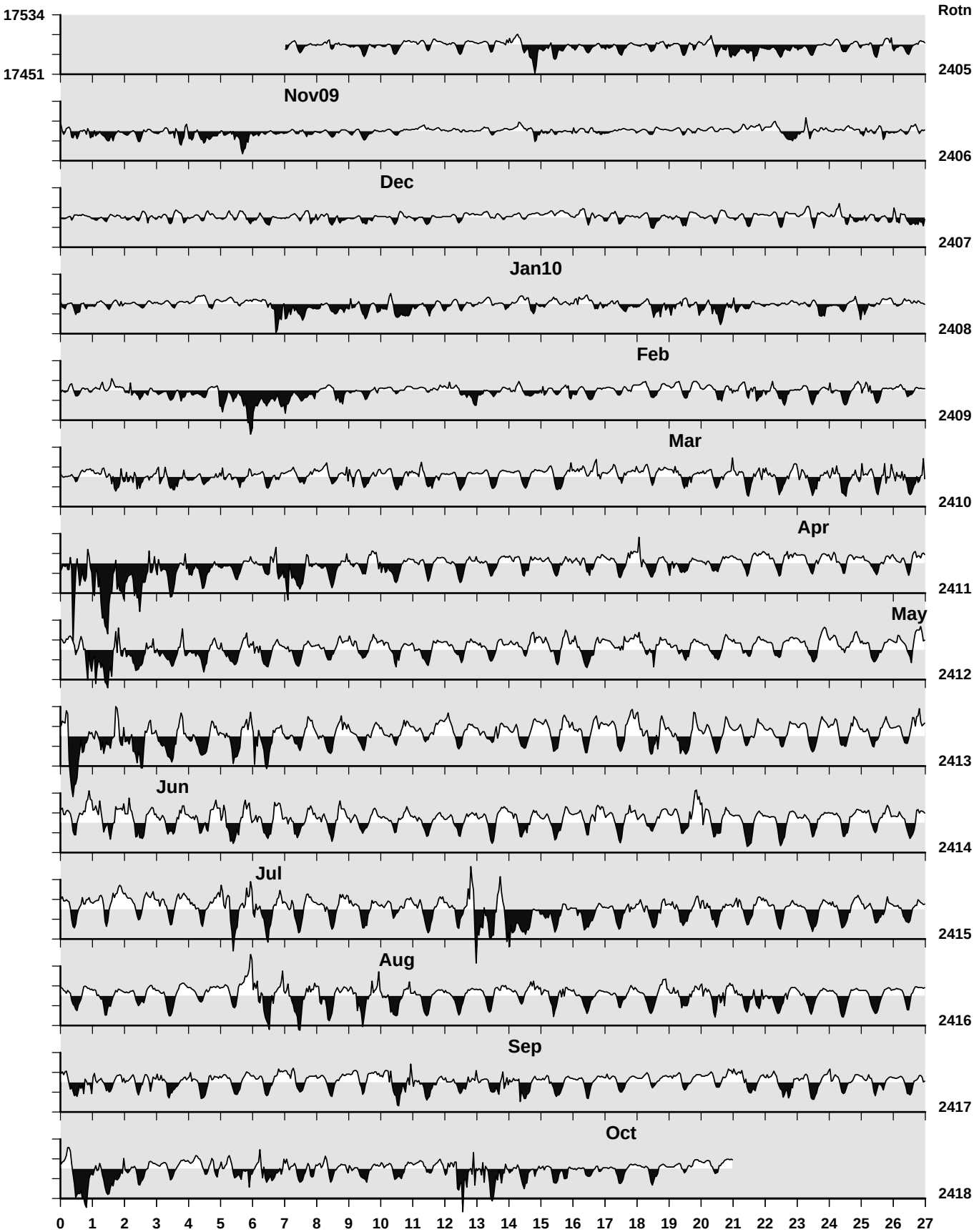

Date: 30-10-2010
Day number: 303
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24
Hour (UT)


Eskdalemuir
Date: 31-10-2010
Day number: 304
Declination in degrees east
-3.54
15min
Horizontal Intensity in nT
17492
20nT
Vertical Intensity in nT
46362
10nT
0
3
6
9
12
15
18
21
24

Eskdalemuir Observatory: Declination (degrees)

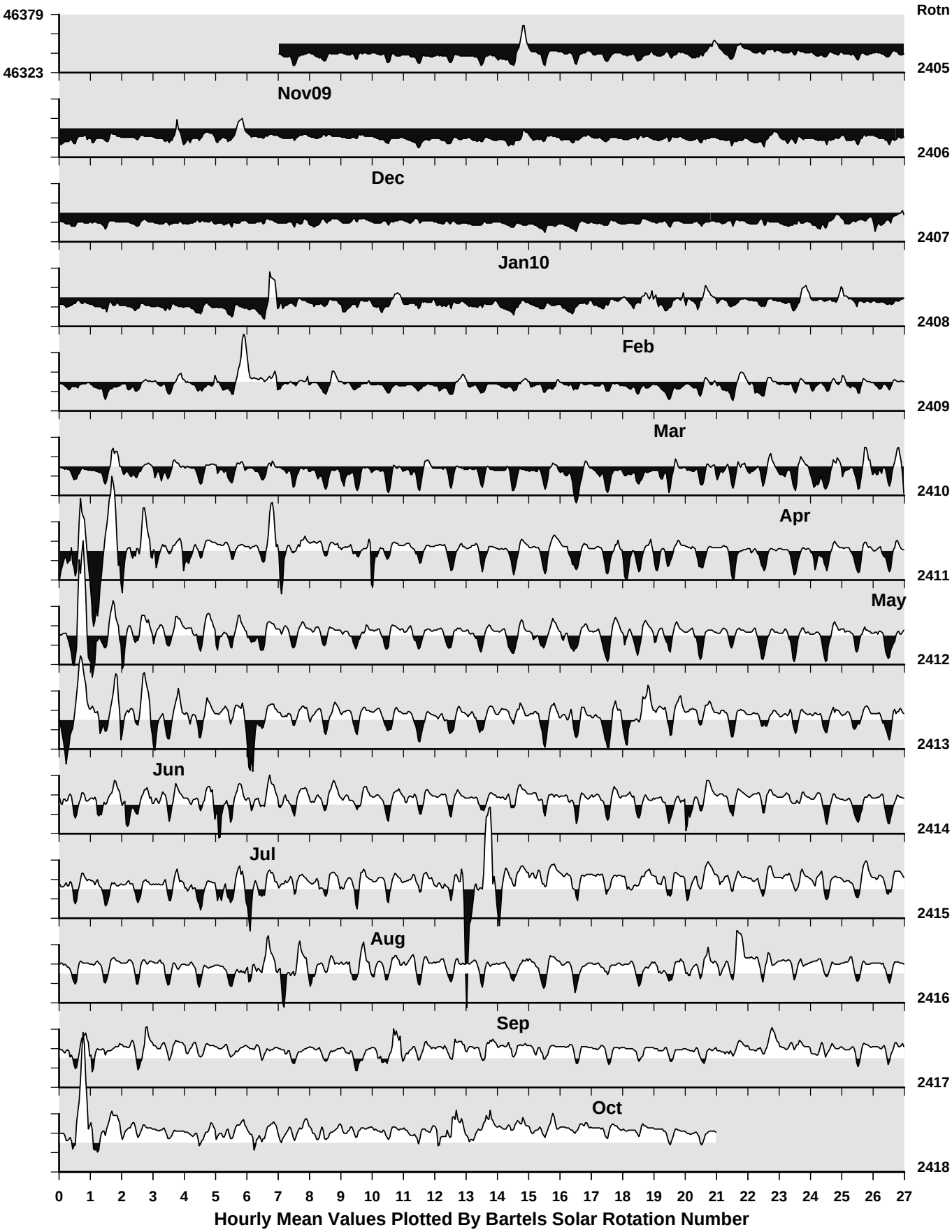


Eskdalemuir Observatory: Horizontal Intensity (nT)



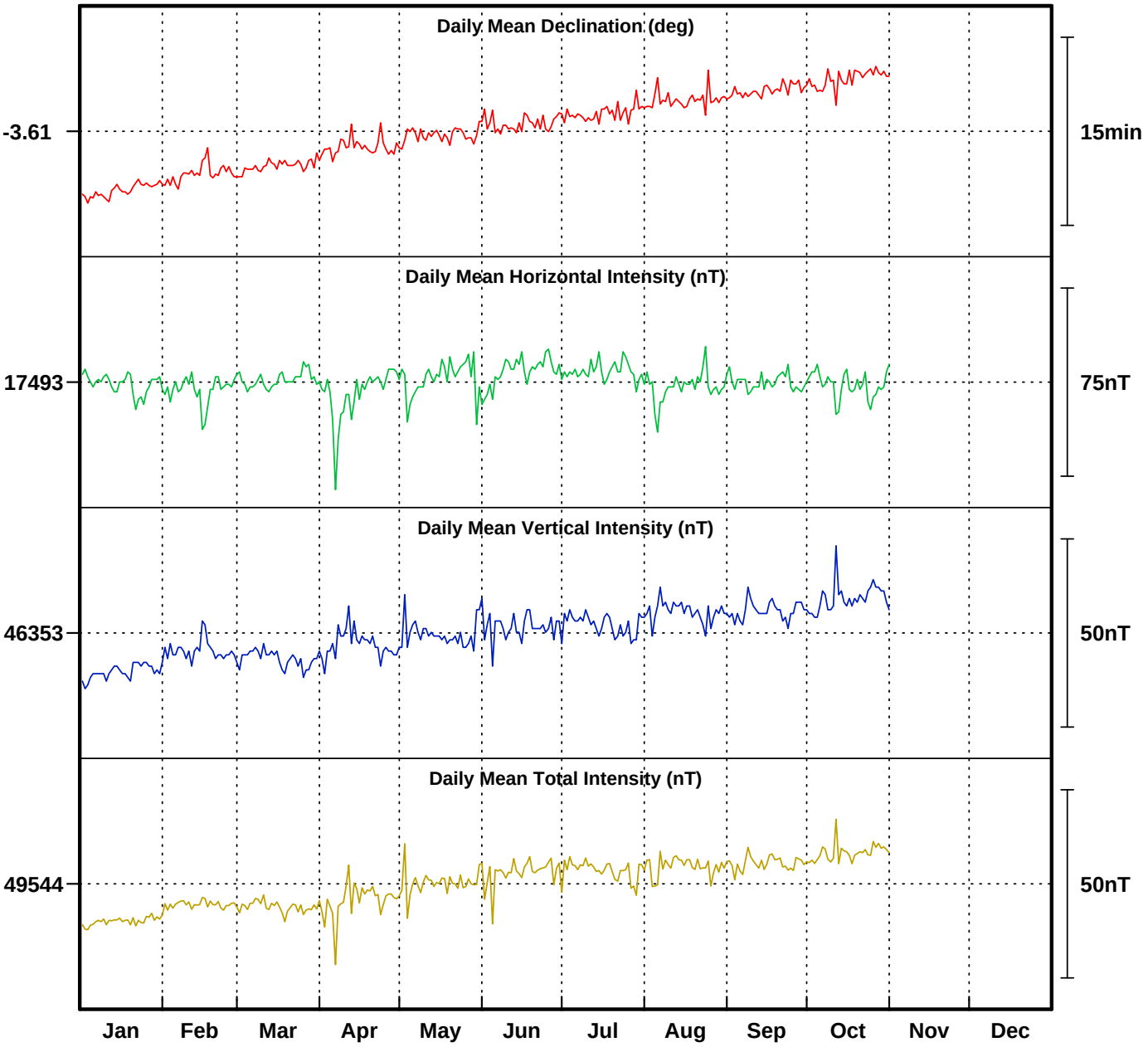
Hourly Mean Values Plotted By Bartels Solar Rotation Number

Eskdalemuir Observatory: Vertical Intensity (nT)



Hourly Mean Values Plotted By Bartels Solar Rotation Number

Eskdalemuir Observatory 2010



Monthly Mean Values for Eskdalemuir Observatory 2010

Month	*D*	*H*	*I*	*X*	*Y*	*Z*	*F*
January	-3° 41.1′	17492 nT	69° 19.2′	17456 nT	-1124 nT	46343 nT	49534 nT
February	-3° 39.7′	17490 nT	69° 19.5′	17454 nT	-1117 nT	46348 nT	49538 nT
March	-3° 39.1′	17494 nT	69° 19.2′	17458 nT	-1114 nT	46346 nT	49538 nT
April	-3° 37.6′	17488 nT	69° 19.7′	17453 nT	-1106 nT	46350 nT	49539 nT
May	-3° 36.6′	17494 nT	69° 19.4′	17459 nT	-1102 nT	46353 nT	49544 nT
June	-3° 35.7′	17497 nT	69° 19.2′	17463 nT	-1097 nT	46354 nT	49547 nT
July	-3° 34.9′	17497 nT	69° 19.2′	17463 nT	-1093 nT	46355 nT	49548 nT
August	-3° 33.9′	17491 nT	69° 19.7′	17457 nT	-1087 nT	46358 nT	49548 nT
September	-3° 33.1′	17493 nT	69° 19.6′	17459 nT	-1084 nT	46359 nT	49549 nT
October	-3° 32.2′	17492 nT	69° 19.7′	17459 nT	-1079 nT	46362 nT	49552 nT

Note

i. The values shown here are provisional.

ESKDALEMUIR RAPID VARIATIONS

SIs and SSCs

Date	Time (UT)	Type	Quality	H (nT)	D (min)	Z (nT)
11-10-10	06 49	SSC*	B	-9.6	-1.80	-1.5
15-10-10	05 12	SSC*	C	4.5	-0.90	-
30-10-10	10 13	SI*	B	12.3	-2.39	-2.0

Notes:
An asterisk (*) indicates that the principal impulse was preceded by a smaller reversed impulse.
The quality of the event is classified as follows:

A = very distinct
B = fair, ordinary, but unmistakable
C = doubtful

The amplitudes given are for the first chief movement of the event.

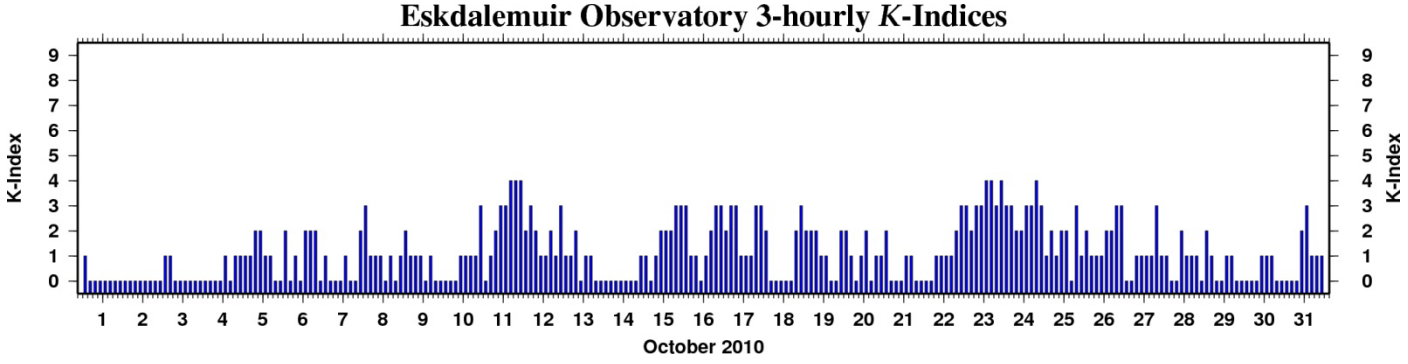
SFEs

Date	Universal Time			H (nT)	D (min)	Z (nT)
	Start	Maximum	End			

Note:
The amplitudes given are for the first chief movement of the event.

INDICES OF GEOMAGNETIC ACTIVITY

Day	*K* - INDICES FOR THREE-HOUR INTERVAL							
	00-03	03-06	06-09	09-12	12-15	15-18	18-21	21-24
1	1	0	0	0	0	0	0	0
2	0	0	0	0	0	0	0	0
3	1	1	0	0	0	0	0	0
4	0	0	0	0	1	0	1	1
5	1	1	2	2	1	1	0	0
6	2	0	1	0	2	2	2	0
7	1	0	0	0	1	0	0	2
8	3	1	1	1	0	1	0	1
9	2	1	1	1	0	1	0	0
10	0	0	0	1	1	1	1	3
11	0	1	2	3	3	4	4	4
12	2	3	2	1	1	2	1	3
13	1	1	2	0	1	1	0	0
14	0	0	0	0	0	0	0	1
15	1	0	1	2	2	2	3	3
16	3	1	1	0	1	2	3	3
17	2	3	3	1	1	1	3	3
18	2	0	0	0	0	0	2	3
19	2	2	2	1	1	0	0	2
20	2	1	0	1	2	0	1	1
21	2	0	0	0	1	1	0	0
22	0	0	1	1	1	1	2	3
23	3	2	3	3	4	4	3	4
24	3	3	2	2	3	3	4	3
25	1	2	1	2	2	0	3	1
26	2	1	1	1	2	2	3	3
27	0	0	1	1	1	1	3	1
28	1	0	0	2	1	1	1	0
29	2	1	0	0	1	1	0	0
30	0	0	0	1	1	1	0	0
31	0	0	0	2	3	1	1	1



The *aa* Index

Date	Day	3-hourly *aa*-indices								Aa_{am}	Aa_{pm}	*Aa*
01-10-10	274	5	5	2	2	2	2	2	2	3.8	2.5	3.1
02-10-10	275	2	2	2	2	2	2	2	2	2.5	2.5	2.5
03-10-10	276	5	8	2	8	2	5	2	2	5.8	3.1	4.5
04-10-10	277	2	2	5	2	8	2	5	8	3.1	5.8	4.5
05-10-10	278	8	12	16	16	8	8	8	5	13	7.2	10.1
06-10-10	279	9	2	8	12	24	12	12	5	7.9	13.3	10.6
07-10-10	280	5	5	5	9	12	2	2	12	6.2	7.2	6.7
08-10-10	281	20	12	5	12	2	8	5	8	12.2	5.8	9
09-10-10	282	12	5	5	12	5	8	2	5	8.5	5.2	6.8
10-10-10	283	5	2	2	8	12	8	8	20	4.5	11.9	8.2
11-10-10	284	8	16	24	81	67	59	45	45	32.2	54.3	43.2
12-10-10	285	20	32	16	20	20	32	8	20	22	20	21
13-10-10	286	8	12	16	8	8	8	8	5	10.9	7.2	9
14-10-10	287	8	2	2	2	2	2	2	8	3.8	3.8	3.8
15-10-10	288	8	5	8	16	16	12	20	20	9.2	16.9	13.1
16-10-10	289	20	8	8	2	12	16	37	24	9.5	22.3	15.9
17-10-10	290	16	24	32	20	12	12	20	20	23.1	15.9	19.5
18-10-10	291	20	5	5	12	5	8	16	32	10.5	15.3	12.9
19-10-10	292	24	24	24	12	8	8	8	12	21.2	8.8	15
20-10-10	293	12	5	2	12	24	5	8	8	7.8	11.3	9.5
21-10-10	294	12	8	2	5	8	8	8	2	6.8	6.5	6.7
22-10-10	295	2	8	12	8	12	8	24	32	7.5	19	13.3
23-10-10	296	24	16	32	59	81	59	37	59	32.8	58.9	45.9
24-10-10	297	32	46	32	32	46	32	45	24	35.5	36.8	36.2
25-10-10	298	9	12	16	16	32	8	24	12	13.3	19	16.2
26-10-10	299	24	12	12	16	12	16	37	20	16	21.3	18.6
27-10-10	300	8	8	8	8	12	8	20	8	7.8	11.9	9.8
28-10-10	301	5	2	8	8	20	8	5	5	5.8	9.6	7.7
29-10-10	302	12	12	5	2	12	12	2	2	7.8	7.2	7.5
30-10-10	303	2	2	2	12	8	16	2	5	4.8	7.9	6.4
31-10-10	304	2	12	9	24	32	8	12	8	12	14.9	13.5
											Monthly Mean Value	13.2

Notes

i. The units of the *aa* index are nT.

ii. The 3-hour *aa* values are rounded to the nearest integer. Where *aa* = *.5, *aa* is rounded down.

iii. Daily values (Aa_{am}, Aa_{pm} and *Aa*) are computed from *aa* values of original resolution.

iv. The monthly mean value is computed from the daily mean values, *Aa*.

v. Definitive *aa* indices are derived and published by the International Service for Geomagnetic Indices.